

# Artifact-suppressing reconstruction of strongly interacting objects in X-ray near-field holography without a spatial support constraint

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**Abstract:** The phase problem is a well known ill-posed reconstruction problem of coherent lens-less microscopic imaging, where only the squared magnitude of a complex wavefront is measured by a detector while the phase information from the wave field is lost. To retrieve the lost information, common algorithms rely either on multiple data acquisitions under varying measurement conditions or on the application of strong constraints such as a spatial support. In X-ray near-field holography however, these methods are rendered impractical in the setting of time sensitive *in situ* and *in operando* measurements. In this paper, we will forego the spatial support constraint and propose a projected gradient descent (PGD) based reconstruction scheme in combination with proper preprocessing that significantly reduces artifacts for refractive reconstructions from only a single acquired hologram without a spatial support constraint. We demonstrate the feasibility and robustness of our approach on different data sets obtained at the nano imaging endstation of P05 at PETRA III (DESY, Hamburg) operated by Helmholtz-Zentrum Hereon.

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## 1. Introduction

The phase problem is a well known ill-posed reconstruction problem of coherent lens-less microscopic imaging, where only the intensity of a complex wave field is measured by a detector while the phase information is lost [1–5]. In those imaging setups, the detector only measures the wavefront as an integral over a certain exposure time which is proportional to the squared magnitude of the wavefront. Naively, one would aim to recover the lost phase information in the detector plane. In practice however, the reconstruction target is not the lost phases at the detector plane but typically the transmission function of the object, that characterizes the interaction with the X rays. This inverse problem is under-determined, since this function is complex, whereas the detector data is real.

Naturally, some approaches alleviate this issue by obtaining more data while slightly changing the experimental setup. This includes for example ptychography [6–8] or full-field imaging [9] with multiple distances [10–12]. However, measuring more data is not always possible or not

always desired, for example in single-pulse imaging at X-ray free electron lasers or *in situ/in operando* studies of material degradation and battery charge cycles at other accelerator-based light sources like synchrotron radiation sources [13–16]. The beamline P05 at PETRA III (DESY, Hamburg) offers a lens-less X-ray in-line holography setup for nanotomography (Fig. 1) [17–20]. Here, the object is fully illuminated by hard X-rays and in-line holograms are acquired in the so-called Fresnel regime. To minimize the X-ray dose on the objects as well as to maximize the measurements throughput, the object transmission function has to be recovered from a single hologram. While the reconstruction of weak interacting samples works well with simple approaches like the direct inversion using the contrast transfer function (CTF) [11] or standard alternating projection based algorithms like the famous Gerchberg-Saxton algorithm [21] and its successors [22–25], the reconstruction complexity increases with the interaction strength of the measured objects. For strongly interacting objects, many algorithms run into phase wrapping problems, take a very long time to converge to a solution or fail to reconstruct the object at all if it is not a phase object but it is also significantly absorbing [26].

These problems can be partly addressed by spatial regularization. This is the finite spatial support of the object, together with a non-negativity constraint that enforces a positive electron density [27, 28]. However, employing a spatial constraint has limits. One is that the object has to possess a finite spatial support. Furthermore, the support has to be either manually set before the actual reconstruction or automatically found during the reconstruction process [13] which adds further computational complexity and refinement parameters. This renders a spatial support constraint unpractical for certain applications like time sensitive *in situ* and *in operando* measurements.

To solve the phase wrapping problem, a projected gradient descent (PGD) based algorithm, denoted as refAP has been developed for the direct reconstruction of the projected refractive index of an object [26]. The advantages of this formulation are recapitulated in Sec. 3. In this work, we further refine this approach by proposing a preprocessing and reconstruction scheme which is able to reconstruct images from single holograms without a spatial support constraint while significantly suppressing artifacts. The paper is structured as follows. In the problem statement Sec. 2 we go into the details of the current reconstruction issues and the sources of possible artifacts. We then give an overview on the current state of preprocessing and reconstruction methods in Sec. 3. In Sec. 4, we propose multiple techniques to reduce reconstruction artifacts and to improve the convergence speed with respect to a standard iterative algorithm (refAP). An overview of the derived algorithms of Sec. 4 is shown in the appendix Sec. 7. In Sec. 5, we show reconstruction results of data from beamline experiments and compare the different enhancements against the current refAP. We will demonstrate that the proposed final algorithm is robust, requires only a few parameters that need to be manually tuned and performs efficiently with a small amount of iterations required, compared to the standard refAP. The data and the software underlying the results presented in this paper are available under DOI: 10.5281/zenodo.8349365.

## 2. Problem statement

Figure 1 shows a setup for X-ray near-field holography, using the divergent illumination of a nanofocusing optics to enable microscopy. The goal of the reconstruction is to retrieve the projected refractive index  $\tilde{O}$  of a measured specimen with respect to its spatial coordinates from a hologram. The object transmission function describes the interaction of an object with the X-ray illumination. Commonly used algorithms are grouped into direct methods, alternating projection (AP) based [21–25] and projected gradient descent (PGD) based algorithms.

To avoid the phase wrapping problem, a PGD-based algorithm, denoted as refAP has been developed [26]. Nevertheless, AP and PGD based algorithms both typically have problems with image artifacts unrelated to phase wrapping as illustrated in Fig. 2. The images in Fig. 2 were obtained by a standard refAP algorithm. These distinct artifacts originate from several different

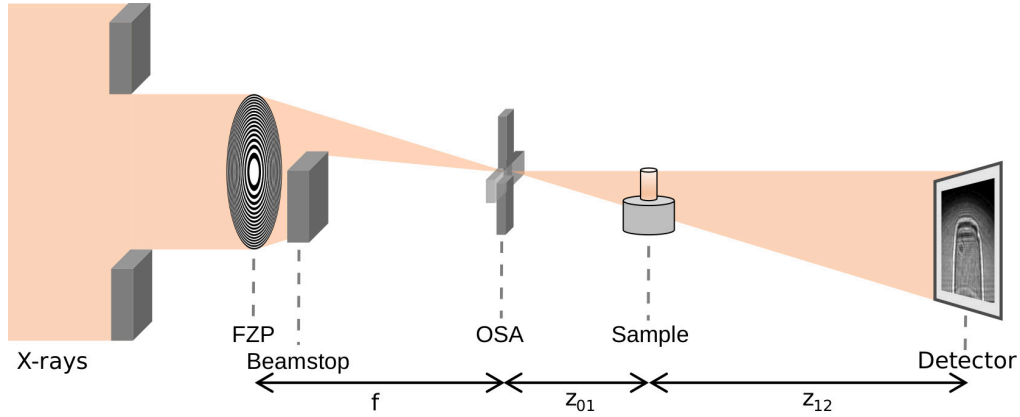


Fig. 1. Sketch of an experimental setup based on a Fresnel zone plate (FZP) for near-field holographic microscopy. The FZP focuses the incoming coherent monochromatic X rays to the focal spot located at  $f$ . There, an order sorting aperture (OSA) is placed that blocks the higher diffraction orders of the FZP. The sample is put into the diverging cone-shaped beam of the FZP at a distance  $z_{01}$ . Behind the sample, the X-rays propagate to the detector that is placed at the sample-to-detector distance  $z_{12}$ . To protect the detector from radiation damage, the direct beam is blocked by a beamstop behind the FZP [20].

95 sources:

- 96 • Non-linear and non-convex optimization problems: These types of inverse problems tend  
97 to have multiple local minima, which can trap [29] or slow down iterative algorithms as  
98 they try to find a path to a global optimum. An algorithm may also converge to different  
99 local minima depending on the initial values.
- 100 • Truncation of information: Due to the limited size of the detector, the hologram measured  
101 at the detector is truncated. If not handled properly, this loss of information can lead to  
102 different kinds of artifacts in the reconstruction. Ring-like or stripe-like artifacts that are  
103 particularly pronounced at the edges are common.
- 104 • Forward model induced reconstruction bias: In X-ray near-field holography, the forward  
105 model is sensitive to the second spatial derivative of the measured object with respect to its  
106 refractive indices. This can be seen by calculating the Taylor expansion of the propagation  
107 kernel in the forward model ([30], Eq. 4.113). The steepness of edges of the reconstructed  
108 object and the residual reconstruction error are directly and proportionally coupled in a  
109 way that steep edges have a higher contribution to the residual reconstruction error than  
110 areas with small changes of the refractive index. Hence, typical iterative reconstruction  
111 algorithms tend to first reconstruct the edges before they reconstruct the object's interior.  
112 This causes a slow reconstruction speed of large structures which have large values for  $\tilde{O}$   
113 but small values for  $\nabla \tilde{O}$  and can lead to artifacts, if the algorithm stops too early.
- 114 • Regularization induced reconstruction bias: Regularization techniques always entail a bias.  
115 The more regularization is applied, the stronger this bias is. This can be exploited to derive  
116 a warm-up phase for the reconstruction but can also manifest in artifacts, e.g. blurred object  
117 edges. The more parameters exist that have to be tuned for the reconstruction, the higher is  
118 the risk for bias and artifacts that depend on the used regularization techniques.
- 119 • Structured noise from illumination: Unstructured measurement noise is a general issue in  
120 inverse problems that can be addressed using regularization strategies such as Tikhonov

regularization. For structured noise, commonly used regularization strategies are less effective. Low spatial frequency artifacts are often observed, especially in object-free areas. In this particular measurement setup, various types of structured noise remain in the data e.g. by the illumination. First, changes in the illumination patterns, for example due to electron refills of the storage ring at the synchrotron radiation source, are not covered by a standard flat-field correction approach. Second, if the illumination of the measured object is larger than the reconstructed FOV, the measured hologram at the detector is partly superimposed by a wave field that has been propagated from outside of the selected FOV.

To counteract this list of problems, many reconstruction methods in X-ray near-field holography rely on the acquisition of diverse data sets such as multi-distance scans [11, 12] or employ additional constraints such as applying a spatial support mask [27, 28]. However, the application of these two particular constraints is either time consuming or object dependent. For *in situ/operando* measurements at the beamline, however, we have high demands on the reconstruction speed, quality, and robustness. Therefore, we have a strong need to find a method that performs high quality reconstruction from only a single hologram. In this paper, we propose a combined preprocessing/reconstruction scheme that addresses most of the above problems and produces artifact-free phase images from a single hologram, without the need for a spatial support constraint.

### 3. Current state of reconstruction

#### 3.1. Refractive forward model for holographic imaging

The physical properties of the object are encoded as complex refractive indices

$$n(x, y, z) = 1 - \delta(x, y, z) + i\beta(x, y, z), \quad (1)$$

along the spatial coordinates  $(x, y, z)$ , where  $\beta \in \mathbb{R}$  describes the absorption, i.e. the attenuation,  $\delta \in \mathbb{R}$  the dispersion, i.e. phase shifting properties and  $i$  is the imaginary unit. In this paper, we assume a sufficiently thin object to consider only the projection of the refractive indices over the object thickness  $d$  in beam direction. The object is illuminated by a monochromatic wave field  $\psi_0$ . We follow the projection approximation of Paganin, section 2.2, equation 2.39 [30] and get the transmission function that describes the exit wave  $\psi_{\text{exit}}$  from the interaction between the illumination and the object

$$\psi_{\text{exit}}(x, y) = \exp\left(-ik \int_0^d \delta(x, y, z) - i\beta(x, y, z) dz\right) \psi_0, \quad (2)$$

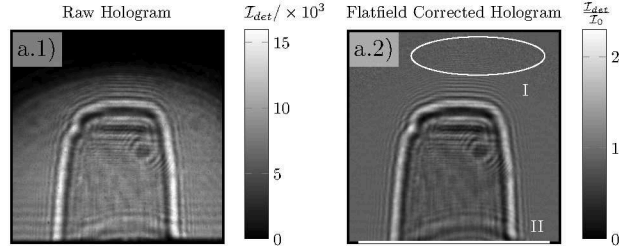
where  $k = \frac{2\pi}{\lambda}$ , with  $\lambda$  the wavelength of the illumination. For the sake of simplicity of these equations, we substitute the exponent of the exponential function and define the refractive object by

$$\tilde{O}(x, y) = -k \int_0^d \delta(x, y, z) - i\beta(x, y, z) dz = \phi(x, y) + i\mu(x, y). \quad (3)$$

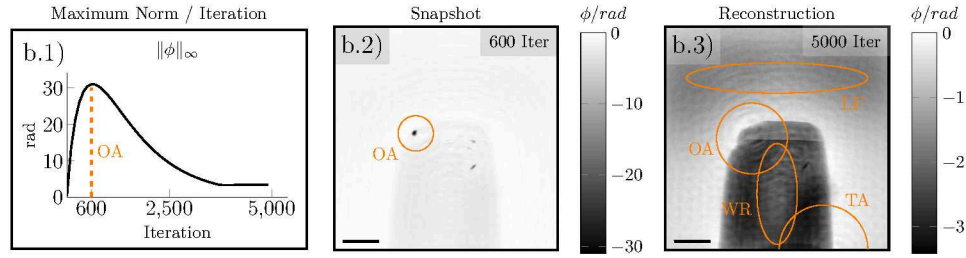
We further omit in the following the notation of the spatial coordinates. In the substitution,  $\phi$  encodes the phase shifting properties and  $\mu$  encodes the attenuation properties of the object. The reconstructions in this paper will aim to recover  $\tilde{O}$ . In idealized conditions, the object is illuminated by an aberration free coherent beam with constant amplitude. In this case, the exit wave behind the object is described by

$$\psi_{\text{exit}}(\tilde{O}) = \exp(i\tilde{O}) A_0 \exp(i\phi_0), \quad (4)$$

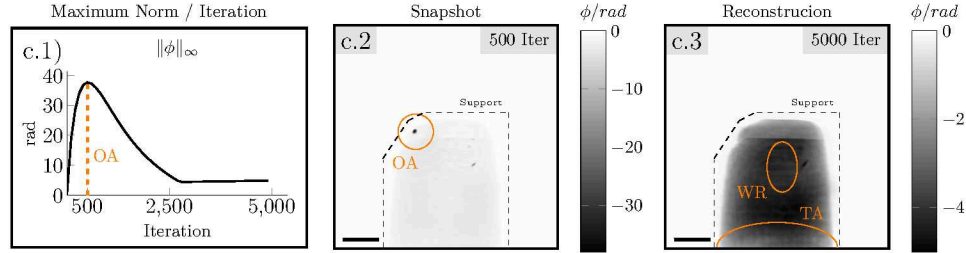
with  $A_0 : \mathbb{R}^2 \mapsto \mathbb{R}$  the illumination amplitude and  $\phi_0 : \mathbb{R}^2 \mapsto \mathbb{R}$  the illumination phase offset. Both parameters are in general unknown and have to either be estimated or reconstructed



(a) Raw hologram a.1) acquired at the experiment. Flat-field corrected hologram a.2) that already shows two sources of reconstruction artifacts: (I) A very noisy area where the illumination fades. This is the boundary of the Fresnel zone plate (FZP). (II) The hologram truncation that also truncates the object at the bottom.



(b) Reconstruction without a spatial support constraint. b.1) Maximum norm of the reconstructed phase shift values as function of the iteration number. It can be seen, that the maximum phase value overshoots during the first hundreds of iterations. The corresponding overshooting artifact (OA) at the maximum, which is iteration 600, can be seen in the marked spot in b.2). The reconstruction result b.3) after 5000 iterations is distorted by low frequency artifacts (LF), ring-like truncation artifacts (TA), an overshooting artifact (OA), and a weak reconstructed area (WR).



(c) Reconstruction with a static spatial support, panels analogue to sub figure (b). An overshooting artifact (OA) can be observed at iteration 500. Due to the support constraint, the algorithm was able to recover from the overshooting at around iteration 2500 cf. c.1). The result does not suffer from low frequency artifacts (LF) in the object free area, also not inside of the support. Remaining issues are a truncation artifact which appears as a weakly reconstructed area at the boundary (TA) and convergence issues (WR) resulting in a weakly reconstructed area in the center of the object.

Fig. 2. Comparison of reconstruction artifacts between a reconstruction from (a) a hologram, (b) without a spatial support constraint and (c) with a spatial support constraint. The images in Fig. 2 were obtained by a standard refAP algorithm (Alg. 2 with and without spatial support constraint). The scale bars indicates  $10\text{ }\mu\text{m}$ .

separately. The measurement process of a hologram is then modeled by propagating  $\psi_{\text{exit}}$  to the detector and taking the squared magnitude of the propagated wave field. We summarize the illumination-interaction-propagation process into an operator  $\mathcal{D}_{\text{Fr}}(\tilde{O})$  and take the squared magnitude as an extra step:

$$\mathcal{D}_{\text{Fr}}(\tilde{O}) = \mathcal{F}^{-1} \circ \exp\left(-i \cdot \pi \frac{(k_x^2 + k_y^2)}{\text{Fr}}\right) \circ \mathcal{F} \circ \psi_{\text{exit}}(\tilde{O}) \quad (5)$$

$$\mathcal{I}_{\text{det}} = |\mathcal{D}_{\text{Fr}}(\tilde{O})|^2 \quad (6)$$

$\mathcal{D}_{\text{Fr}}$  is an approximation of the free space propagation for the near field and is called the Fresnel propagator in operator form [30]. Here, the function  $\mathcal{F}$  describes the 2D Fourier transform of the wave field and  $\mathcal{F}^{-1}$  its inverse respectively with respect to the spatial coordinates  $(x, y)$  and the coordinates in frequency space  $(k_x, k_y)$ . The Fresnel number  $\text{Fr}$  depends on the geometry of the illuminating beam. In a parallel beam setup, the Fresnel number is derived from the pixel size of the detector  $\Delta x$ , the wavelength of the source  $\lambda$  and the propagation distance between sample and detector  $z_{12}$ :

$$\text{Fr} = \frac{\Delta x^2}{\lambda z_{12}}. \quad (7)$$

In order to enable microscopy through holographic imaging, the object has to be placed into a divergent beam in a defocused position. The Fresnel number must then also incorporate the magnification through the cone beam which can be derived from the Fresnel scaling theorem [30]. The theorem states that the cone beam setup can be transformed to a virtual parallel beam setup by calculating a new effective propagation distance  $z_{12}^* = M z_{12}$  from the object magnification  $M$ . The divergent beam magnifies the sample by a factor

$$M = \frac{z_{01} + z_{12}}{z_{01}} \quad (8)$$

that is given by the proportion of the total propagation distance  $z_{01} + z_{12}$  to the focus-object distance  $z_{01}$ . The Fresnel number for a cone beam becomes

$$\text{Fr} = \frac{\Delta x^2}{\lambda M z_{12}}. \quad (9)$$

### 3.2. Preprocessing

Reconstruction algorithms in X-ray near-field holography assume idealized conditions for the source and the setup components. An ideal setup has aberration-free optics and perfect coherent illumination which are not fulfilled in practice. An important tool for any reconstruction is to prepare the acquired data in a preprocessing step such that the assumptions of the forward model are met while respecting issues and limitations of certain implementations. Two important issues that are typically handled in a preprocessing step are non-uniform static illumination and spectral leakage originating from the fast Fourier transform of a non-periodic data set.

#### 3.2.1. Flat-field correction

To sufficiently approximate ideal conditions for a reconstruction, a flat-field correction approach is often applied on the detector data, that corrects the static part of aberrations. For that, a synthetic flat-field image  $\mathcal{I}_{\text{flat}}$  is derived first from measurements without an object in the beam by principle components analysis [31,32]. The raw data  $\mathcal{I}_{\text{raw}}$  is then divided by this flat-field to yield a flat-field corrected hologram  $\mathcal{I}_{\text{corr}}$ :

$$\mathcal{I}_{\text{corr}} = \frac{\mathcal{I}_{\text{raw}}}{\mathcal{I}_{\text{flat}}}. \quad (10)$$

### 185 3.2.2. Padding

186 The acquired hologram consists of a discrete set of detector pixel values forming a 2D image.  
 187 Part of the forward model is the Fourier transform, computed by its discrete implementation, the  
 188 Fast Fourier Transform (FFT). For the FFT, all data and operators have to be sampled properly at  
 189 a sufficient rate. For the Fresnel propagation operator, it has been shown that this function, which  
 190 is a chirp, has to be sampled at least by  $N \geq \frac{1}{\Delta f}$  pixels in each spatial dimension [33]. A possible  
 191 method to match the required sampling rate in the frequency domain is to pad the hologram in  
 192 the spatial domain. Various methods for padding data already exist [34]. Typical methods are  
 193 padding with a constant value, a repetition of the marginal values or mirroring of the existing  
 194 data towards each orthogonal and diagonal directions.

### 195 3.2.3. Windowing

196 Handling spectral leakage issues of the FFT has been widely studied in the field of signal  
 197 processing. An assumption of the FFT is that the input data is periodic and the sampled data  
 198 covers a full period and can therefore be repeated into each spatial direction. If the input data does  
 199 not cover a full period of a signal, spectral leakage will occur in form of a convolution between  
 200 the true spectrum and a sinc shaped function. A common approach to handle this problem is  
 201 to transform the data into a pseudo periodic signal, before a Fourier transform is applied. To  
 202 this end, a Hadamard product of the discrete data and window image fades the data towards the  
 203 margin. Nevertheless, common windows still create a convolution with a sinc shaped function.  
 204 However, common windows also shape the main lobe, i.e. the passband and the side lobes i.e. the  
 205 stopband in their respective Fourier transforms according to desired properties. Some examples  
 206 are the Hamming, Hann or Blackman windows [35, 36].

## 207 3.3. Reconstruction problem and constraints

208 When the X-ray wave field arrives at the detector, the involved processes can be described by  
 209 linear operations (Eq. 5) except the measurement itself (Eq. 6). Due to the measurement process,  
 210 the phase information of the propagated wave field is lost and the object reconstruction a non-  
 211 linear inverse problem. First papers on this topic [10, 21] introduced algorithms that are based on  
 212 alternating projections onto constraint sets for one or more constraints. They attempt to solve the  
 213 feasibility problem

$$\text{find } \psi \in \mathcal{I}_{\text{det}} \cap \Omega \quad (11)$$

214 i.e. to find the intersection of two sets namely the holograms  $\mathcal{I}_{\text{det}}$  and a set of object constraints  
 215  $\Omega$ . These algorithms converge to points of shortest distances to the constraint set and posses the  
 216 property to follow a performance criterion named Summed Distance Error (SDE) [29] stated

$$\text{SDE}(\psi) = \|\mathcal{P}_{\mathcal{I}_{\text{det}}}\psi - \psi\|_2 + \|\mathcal{P}_{\Omega}\psi - \psi\|_2, \quad (12)$$

217 where  $\mathcal{P}_x$  are the projections of  $\psi$  onto the constraint sets.

218 In recent decades huge progress has been made in the field of convex analysis, providing us  
 219 with a mathematical framework and necessary tools to approximate solutions to different kinds  
 220 of inverse problems [37–39]. Many of the reconstruction algorithms used in X-ray holography  
 221 can be associated with relatives [21–23] or have been even actively derived from concepts of  
 222 convex analysis [24, 25].

223 A common approach in convex optimization is to formulate an objective function that has to be  
 224 optimized. The solution of this optimization problem yields the desired reconstruction result. An  
 225 objective function typically consists of multiple terms, where at least one describes the consistency  
 226 of a solution with the set of measurements and some additional terms to include prior knowledge  
 227 and regularization. Here, given a single exposure hologram  $\mathcal{I}_{\text{det}}$ , the reconstruction problem

for the refractive index in X-ray near-field holography can be formulated using a regularized least-squares approach

$$\tilde{O}^* = \underset{\tilde{O}}{\operatorname{argmin}} \frac{1}{2} \left\| |\mathcal{D}_{\text{Fr}}(\tilde{O})| - \sqrt{\mathcal{I}_{\text{det}}} \right\|_2^2 + \chi_{\Omega}(\tilde{O}). \quad (13)$$

The term  $\chi_{\Omega}$  is an indicator functions, defined by

$$\chi_{\Omega}(\tilde{O}) = \begin{cases} 0 & \tilde{O} \in \Omega \\ +\infty & \text{else} \end{cases} \quad (14)$$

that models prior knowledge of the setup and the object. An established choice of  $\Omega$  is a constraint set that consists of the following:

$$\Omega = \Omega_P \cap \Omega_S \quad (15)$$

with  $\Omega_P$  being the set of physically valid reconstruction values that models the interaction between X rays and the object such that

$$\Omega_P = \{ \forall x \in \tilde{O} : \operatorname{Re}(x) \in [-\infty, 0], \operatorname{Im}(x) \in [0, \infty] \}. \quad (16)$$

$\Omega_S$  is the spatial support of the object. As stated in Sec. 2, we aim to omit this constraint.

## 4. Methods

### 4.1. Data Preprocessing

In this section, we aim to reduce truncation artifacts that result from the limited detector size. We also take care of spectral leakage that is caused by the fast Fourier transform on the non-periodic measured hologram.

#### 4.1.1. Padding

The proposed padding scheme for the preprocessing aims to reduce truncation artifacts that result from edges that are introduced at the hologram border. These edges appear for example as propagation fringes when a truncated wavefield is propagated from the detector plane to the object plane. To avoid this, edges have in general to be avoided when the data is continued beyond the truncation edge. In one dimension, this can be easily achieved by a repetition of the marginal values of the hologram. In two dimensions it is more complex since a simple repetition in one direction introduces edges in the orthogonal direction. Our approach here is to extend the data by mirroring the acquired hologram (see Fig. 3a) into each direction (see Fig. 3b). Although mirroring is not a physically correct model and the truncated hologram continues in a way that is unknown, it is a way to avoid truncation artifacts through consistency. At the mirroring border, we avoid sudden edges while in the extended area, instead of truncation artifacts, the data will simply resolve to a copy of the reconstructed object. We then pad the mirrored hologram with the value  $A_0$  of the constant probe model from Eq. 4 to the necessary size to match the sampling rate of the Fresnel propagation operator as described in Sec.3.2. The result is shown in Fig. 3c. We choose  $A_0$  such that it is consistent with the flat-field corrected hologram. An ideal flat field correction would normalize  $A_0$  to one. However, after applying the flat-field approach as described in Sec. 3.2.1, an offset remains. Currently, the offset parameter  $A_0$  has to be tuned manually.

#### 4.1.2. Windowing

Fig. 4 illustrates the steps for the creation of the fading window. To reduce spectral leakage in the frequency domain, we apply a two dimensional window function on the extended detector



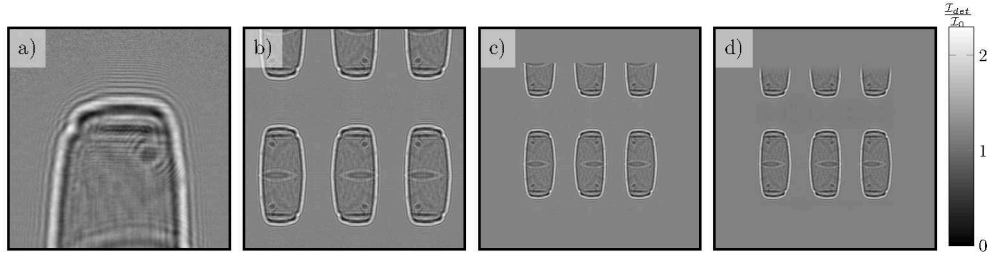


Fig. 3. Preprocessing scheme. The raw hologram is flat-field corrected a) by a PCA approach [31, 32]. To solve the marginal problem, the flat-field corrected hologram is mirrored along each direction, yielding b). To match the required sampling rate of the Fresnel propagation convolution kernel, the result is padded by a constant which is the estimated illumination  $A_0$ , yielding c). A two dimensional fadeout mask (Fig. 4.c) is then applied, yielding d).

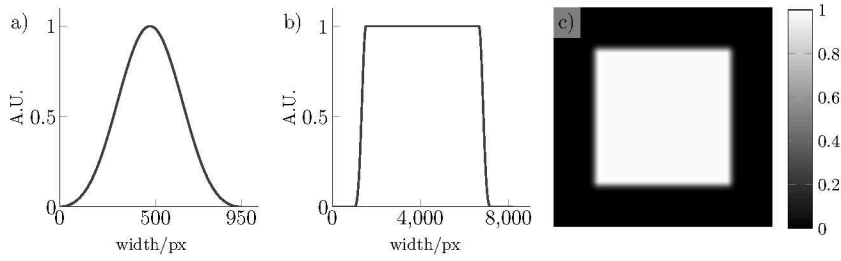


Fig. 4. Window function. a) A one dimensional Blackman function is sampled to the size  $w$  of Eq. 17. b) The window function is split in half and the gap filled with ones, until desired fading start and fading end is reached. The result is then zero padded until the window width matches the size of the input data for the reconstruction. From that, the two dimensional window function c) is derived by a repetition in each direction.

263 data, such that after mirroring and padding, the values of the original hologram in the center are  
 264 unchanged. To create such a window, we modify one of the common fading windows which are  
 265 typically bell shaped. Note that an unmodified window would induce a fading into the original  
 266 hologram that we aim to preserve. We start with a one dimensional Blackman window

$$y(w, x) = 0.42 - 0.5 \cdot \cos\left(\frac{2\pi x}{w-1}\right) + 0.08 \cdot \cos\left(\frac{4\pi x}{w-1}\right). \quad (17)$$

267 Here,  $w$  is the window width and  $x$  the one dimensional spatial coordinate. To preserve a fading  
 268 free area in the center, we split the one dimensional window Fig. 4a at the center and continue  
 269 to fill the gap with the constant value one, until the desired starting points and end points for  
 270 the fading are reached. The result is then zero padded to match the size of the input data for  
 271 the reconstruction. The new one dimensional window is shown in Fig. 4b. From that, the two  
 272 dimensional window function is derived by a repetition of the one dimensional function into each  
 273 spatial direction, yielding Fig. 4c that leaves a rectangular shaped fading free area in the center.  
 274 The respective hologram with the applied window is shown in Fig. 3d.

275 The processing described here in Sec. 4.1 is summarized in Alg. 1 in the appendix.

## 276 4.2. Reconstruction

277 In this section, we aim to increase the convergence rate and reduce the image artifacts discussed  
 278 initially in the problem statement. In the following subsections, we start first with the introduction  
 279 of a basic support-less algorithm for comparison. We then propose two modifications to address  
 280 convergence issues and artifacts and eventually combine these in a final algorithm.

### 281 4.2.1. Basic algorithm without a spatial support constraint

282 We begin with the construction of the basic algorithm that finds a solution for equation 13 under  
 283 the constraint set  $\Omega = \Omega_P$  without the spatial support constraint  $\Omega_S$  discussed in Sec. 3.3.  $\Omega_P$  is  
 284 the non-negativity constraint that enforces a positive electron density in the calculated solution,  
 285 which is a convex set.

$$\Omega_P = \{ \forall x \in \tilde{O} : \text{Re}(x) \in [-\infty, 0], \text{Im}(x) \in [-\log(A_0), \infty] \}. \quad (18)$$

286 As a boundary for the set of physical solution,  $-\log(A_0)$  instead of 0 is used, which is the direct  
 287 consequence of the constant probe model  $A_0$  that has been used in Sec. 4.1.1. We get a projector  
 288  $\mathcal{P}_{\Omega_P}$  of  $\tilde{O}$  onto  $\Omega_P$  by projecting the real and imaginary part separately and pointwise onto the  
 289 defined intervals:

$$\mathcal{P}_{\Omega_P}(\tilde{O}) = \min(0, \text{Re}(\tilde{O})) + i \max(-\log(A_0), \text{Im}(\tilde{O})). \quad (19)$$

290 For gradient descent methods, it is in general not recommended to use a standard gradient  
 291 descent step for the optimization of a non-convex function  $f$  in form of  $x_{i+1} = x_i - \alpha \nabla f(x_i)$  [40,41].  
 292 This standard algorithm is known for having trouble with optimizing functions that contain local  
 293 minima, flat surfaces or ravines. Instead, we make use of the Nesterov accelerated gradient  
 294 (NAG) [42,43] with momentum  $\gamma$  and step size  $\eta$ . The modified gradient  $\nabla g$  is given by:

$$\nabla g(\tilde{O}_i) = \gamma \nabla g(\tilde{O}_{i-1}) + \eta \nabla f(\tilde{O}_i - \gamma \nabla g(\tilde{O}_{i-1})). \quad (20)$$

with  $\nabla g(\tilde{O}_0) = 0$ . Here,  $\nabla f(\tilde{O})$  is the non-accelerated gradient of the data fidelity term, given  
 by [26,44]:

$$f(\tilde{O}) = \frac{1}{2} \| |\mathcal{D}_{\text{Fr}}(\tilde{O})| - \sqrt{\mathcal{I}_{\text{det}}} \|_2^2 \quad (21)$$

$$\nabla f(\tilde{O}) = -i \cdot \overline{\exp(i\tilde{O})} \cdot \mathcal{D}_{\text{Fr}}^{-1} \left( \mathcal{D}_{\text{Fr}}(\tilde{O}) - \sqrt{\mathcal{I}_{\text{det}}} \odot \text{sgn}(\mathcal{D}_{\text{Fr}}(\tilde{O})) \right) \quad (22)$$

$$\text{sgn}(x) = \begin{cases} 0 & \text{if } x = 0 \\ \frac{x}{|x|} & \text{else} \end{cases} \quad (23)$$

295 Recalling the preprocessing scheme Alg. 1, we also model the previous extension of the hologram  
 296 as a prior knowledge into the reconstruction. At the beginning of each iteration  $k$ , we apply  
 297 the same preprocessing steps on the current solution  $\tilde{O}_k$ . The complete PGD algorithm with a  
 298 Nesterov momentum accelerated gradient step is shown in Alg. 2 in the appendix.

### 299 4.2.2. Regularization for $\text{Im}(\tilde{O})$

300 In order for Alg. 2 to reach a minimum of the cost function Eq. 13 fast, the Nesterov acceleration  
 301 weight needs to be as high as possible. However, high momentum values entail the risk of  
 302 overshooting in the gradient descent step, that causes overshooting artifacts (problem statement  
 303 (Fig. 2)). Here, we aim to counteract the overshooting in single pixels by the following strategy.  
 304 In the forward model Eq. 5, one single pixel of the calculated exit wave can be described as a  
 305 vector with an amplitude inversely proportional to  $\text{Im}(\tilde{O})$  and an angle proportional to  $\text{Re}(\tilde{O})$ .

For a vanishing vector amplitude, the reconstruction of the angle becomes numerically unstable. Furthermore, high Nesterov momentum values entail a risk of overshooting  $\text{Im}(\tilde{O})$  in single pixels and  $\text{Re}(\tilde{O})$  such that vanishing amplitudes and simultaneously large angles are produced, which could be observed in Fig. 2 b.1 and c.1. To stabilize the reconstruction, we propose a Tikhonov-like regularization only for the amplitude and extend the target cost function by an  $L_2$  term for the absorption values, i.e. the imaginary part of  $\tilde{O}$ :

$$\tilde{O}^* = \underset{\tilde{O}}{\text{argmin}} \frac{1}{2} \left\| |\mathcal{D}_{\text{Fr}}(\tilde{O})| - \sqrt{\mathcal{I}_{\text{det}}} \right\|_2^2 + \mathcal{X}_{\Omega}(\tilde{O}) + \beta \|\text{Im}(\tilde{O})\|_2. \quad (24)$$

We can account for this additional term by a slight modification of the gradient in Alg. 2

$$f(\tilde{O})' = \frac{1}{2} \left\| |\mathcal{D}_{\text{Fr}}(\tilde{O})| - \sqrt{\mathcal{I}_{\text{det}}} \right\|_2^2 + \beta \|\text{Im}(\tilde{O})\|_2 \quad (25)$$

$$\nabla f(\tilde{O})' = -i \cdot \overline{\exp(i\tilde{O})} \cdot \mathcal{D}_{\text{Fr}}^{-1} \left( \mathcal{D}_{\text{Fr}}(\tilde{O}) - \sqrt{\mathcal{I}_{\text{det}}} \odot \text{sgn}(\mathcal{D}_{\text{Fr}}(\tilde{O})) \right) + \mathcal{R}(\tilde{O}) \quad (26)$$

$$\mathcal{R}(\tilde{O}) = \begin{cases} 0 & \text{if } \tilde{O} = 0 \\ i\beta \frac{\text{Im}(\tilde{O})}{\|\text{Im}(\tilde{O})\|_2} & \text{else} \end{cases}. \quad (27)$$

The corresponding algorithm is the same basic approach as Alg. 2, extended with a modified gradient and a parameter update loop, which yields Alg. 3 in the appendix.

#### 4.2.3. Suppression of high spatial frequencies

Here, we propose two approaches to counteract the slow reconstruction of low spatial frequencies. In general, this can be done either by increasing the contribution of low spatial frequencies or by decreasing the contribution of high spatial frequencies to the objective function.

First, we suppress the acceleration of high spatial frequencies by the Nesterov accelerated gradient, which causes overshooting of pixel groups down to single pixels for high momentum values. Here, we aim to compensate for the Fresnel propagation kernel's property of being sensitive to the second derivative, which makes a momentum based gradient descent acceleration unstable for high spatial frequencies. Our goal is to reduce the momentum for high spatial frequencies while maintaining a high momentum for low spatial frequencies. To this end, we replace the momentum  $\gamma$  by an operator  $\tau$  which first transforms the Nesterov accelerated gradient into the Fourier domain and then applies frequency-dependent weights with Gaussian distribution:

$$\tau = \gamma \mathcal{F}^{-1} \circ \exp\left(-2\pi^2(k_x^2 + k_y^2)\sigma^2\right) \circ \mathcal{F}, \quad (28)$$

which yields then the Nesterov accelerated gradient

$$\nabla g(\tilde{O}_i) = \tau \nabla g(\tilde{O}_{i-1}) + \eta \nabla f(\tilde{O}_i - \tau \nabla g(\tilde{O}_{i-1})). \quad (29)$$

with the parameters  $\gamma$  and  $\sigma$  included in  $\tau$ . The second approach is a multigrid method. Here an approximate low resolution solution is obtained on coarse grid first. Formally, this can be done by solving the optimization problem

$$\tilde{O}^* = U^{-1} \left[ \underset{\tilde{O}}{\text{argmin}} \frac{1}{2} \left\| |\mathcal{D}_{\text{Fr}}(\tilde{O})| - \sqrt{U \mathcal{I}_{\text{det}}} \right\|_2^2 + \mathcal{X}_{\Omega}(\tilde{O}) \right], \quad (30)$$

where  $U$  and  $U^{-1}$  are the corresponding down- and up-sampling matrices. The approximate solution to this problem is then used as an initial guess in the above optimization problem on a refined grid, represented by new down- and up-sampling matrices. This process is iterated until the original resolution is restored. The stopping criterion is a maximum iteration number. A detailed breakdown of the algorithm can be found in Alg. 4 in the appendix.

#### 335 4.2.4. Artifact-Suppressing Reconstruction Method

336 We create the final algorithm by combining all proposed methods to a global artifact-suppressing  
 337 reconstruction method which we call ASRM. We start again with the basic algorithm Alg. 2 and  
 338 extend the algorithm by first apply the weighted regularization term  $\beta\|\text{Im}(\tilde{O})\|_2$  of Sec. 4.2.2.  
 339 We also add the high frequency suppression and the multigrid method from Sec. 4.2.3, which  
 340 is the extension of the objective function by a down sampling matrix  $U$ , an up sampling matrix  
 341  $U^{-1}$  and the modified momentum operator  $\tau$  in the Nesterov accelerated gradient. The objective  
 342 function to reconstruct the complex refractive object is then

$$\tilde{O}^* = U^{-1} \left[ \underset{\tilde{O}}{\text{argmin}} \frac{1}{2} \left\| |\mathcal{D}_{\text{Fr}}(\tilde{O})| - \sqrt{U \mathcal{I}_{\text{det}}} \right\|_2^2 + \mathcal{X}_{\Omega}(\tilde{O}) + \beta \|\text{Im}(\tilde{O})\|_2 \right]. \quad (31)$$

In each iteration, a Nesterov accelerated gradient  $\nabla g(\tilde{O})$  is calculated by

$$\tau = \gamma \mathcal{F}^{-1} \circ \exp\left(-2\pi^2(k_x^2 + k_y^2)\sigma^2\right) \circ \mathcal{F} \quad (32)$$

$$\nabla g(\tilde{O}_i) = \tau \nabla g(\tilde{O}_{i-1}) + \eta \nabla f(\tilde{O}_i - \tau \nabla g(\tilde{O}_{i-1})). \quad (33)$$

where  $\nabla f(\tilde{O})$  is the analytical gradient of the data fidelity term combined with the  $L_2$  regularization of the imaginary part of  $\tilde{O}$ :

$$f(\tilde{O}) = \underset{\tilde{O}}{\text{argmin}} \frac{1}{2} \left\| |\mathcal{D}_{\text{Fr}}(\tilde{O})| - \sqrt{\mathcal{I}_{\text{det}}} \right\|_2^2 + \beta \|\text{Im}(\tilde{O})\|_2 \quad (34)$$

$$\nabla f(\tilde{O}) = -i \cdot \overline{\exp(i\tilde{O})} \cdot \mathcal{D}_{\text{Fr}}^{-1} \left( \mathcal{D}_{\text{Fr}}(\tilde{O}) - \sqrt{\mathcal{I}_{\text{det}}} \odot \text{sgn}(\mathcal{D}_{\text{Fr}}(\tilde{O})) \right) + \mathcal{R}(\tilde{O}) \quad (35)$$

$$\text{sgn}(x) = \begin{cases} 0 & \text{if } x = 0 \\ \frac{x}{|x|} & \text{else} \end{cases} \quad (36)$$

$$\mathcal{R}(\tilde{O}) = \begin{cases} 0 & \text{if } \tilde{O} = 0 \\ i\beta \frac{\text{Im}(\tilde{O})}{\|\text{Im}(\tilde{O})\|_2} & \text{else} \end{cases} \quad (37)$$

343 The summarized algorithm is shown in Alg. 5 in the appendix.

## 344 5. Experiments

### 345 5.1. Experimental Setup

346 We obtained four datasets at the beamline P05 at PETRA III, located at DESY in Hamburg,  
 347 operated by Helmholtz-Zentrum Hereon [17–19]. A Fresnel-zone-plate-based setup for NFH as  
 348 shown in Figure 1 was used [20]. The detector is a scintillator (10  $\mu\text{m}$  Gadox) sCMOS camera  
 349 (Hamamatsu C12849-101U) with 6.5  $\mu\text{m}$  pixel size at 16 bit image depth and  $2048 \times 2048$  pixels.

350 We chose the following samples for our demonstration: A spider attachment hair (Fig. 5)  
 351 [45–47], the tip of a cactus needle (Fig. 6), a sample of a human tooth prepared by focused ion  
 352 beam milling (Fig. 7) and a partly corroded biodegradable magnesium wire (Fig. 8) [14, 48–50].  
 353 The measurement parameters for each of the samples are shown in Appendix B, Tab. 2.

### 354 5.2. Reconstruction Setup

355 We chose the samples with increasing complexity for the reconstruction. The samples increase  
 356 each in their thickness and their interaction strength with the X-ray illumination possessing  
 357 consecutively larger phase shifts and stronger attenuation properties. The interaction strength  
 358 increases with respect to the given sample order above. The detector data was preprocessed

by Algorithm 1 before the actual reconstruction. The synthetic flat-field  $I_{\text{flat}}$  for the flat-field correction according to Eq. 10 was generated from 50 empty images with a PCA approach [31,32]. The preprocessed data was reconstructed with the reference method and each of the different proposals outlined above:

Alg. 2: Standard PGD method.

Alg. 3: Additional  $L_2$  regularization for the absorption values of  $\tilde{O}$ .

Alg. 4: Suppression of high spatial frequencies.

Alg. 5: The combination of Alg. 3 and Alg. 4.

We implemented our algorithms in Python using the PyTorch library for GPU acceleration [34, 51, 52]. The reconstruction was performed on the Maxwell computing cluster at DESY on a NVIDIA A100 GPU with 40GB memory [53].

### 5.3. Analysis of Reconstruction Quality

In the following, we show the reconstruction results for each variant after 2000 iterations and horizontal cross sections. We measure the reconstruction quality in terms of the ability to reconstruct large phase shifts in the object’s interior while maintaining sharp object edges and low background noise in the object free area. The reconstruction parameters for each sample are shown in Appendix B. Three out of the available parameters remained sample dependent and had to be tuned beforehand: (i) the source intensity value  $A_0$  of the constant source model, (ii) the Gaussian filter width of  $\mathcal{P}_{\Omega_{\text{Nest}}}$  and (iii) the optimal Fresnel number  $\text{Fr}$  of the forward model. The resulting figures 5 to 8 have the same structure and show the reconstruction variants as listed above in their respective panels. We classify the remaining artifacts into categories, according to the respective artifacts in the problem statement Sec. 2:

LF: Low frequency noise  
 TA: Truncation artifacts  
 OA: Overshooting artifacts  
 WR: Weak reconstruction

## 6. Results

For all samples and every algorithm variant, truncation artifacts TA are significantly less visible in the reconstruction results. The remaining artifacts are, depending on the algorithm variants, LF, OA and WR that we summarized into Tab. 1, where we compared the different samples for the algorithm variants and occurring artifacts. The proposed algorithm variants perform as follows:

Algorithm 2, the basic algorithm, yields the worst results in all cases. It failed to reconstruct the object’s interior of the cactus needle, the tooth and the magnesium wire. The results of the spider hair and the magnesium wire possess clearly visible overshooting artifacts. Also, low frequency artifacts are very prominent.

Algorithm 3 reduced the overshooting artifacts of Alg. 2 that were visible in the spider hair and magnesium wire results. The algorithm still failed to completely reconstruct the object’s interior of the cactus needle, the tooth and the magnesium wire. The results also suffer from low frequency artifacts, either in form of a static offset for the spider hair or background variations for the cactus needle, tooth and magnesium wire.

Algorithm 4 significantly suppressed all overshooting artifacts of Alg. 2. Compared to Alg. 3, the algorithm also improved the background variations as well as the reconstruction of the object’s interior. For the cactus needle, the algorithm succeeded to suppress artifacts in the reconstructed phase image and almost reconstructed the object’s interior of the tooth.

Algorithm 5 offers the best reconstruction quality if compared to the other approaches. The combination of both approaches significantly reduced overshooting artifacts and low frequency

398 noise. The algorithm also succeeded in the reconstruction of the object's interior for all tested  
 399 objects.

Sample	Alg. 2			Alg. 3			Alg. 4			Alg. 5		
	LF	OA	WR	LF	OA	WR	LF	OA	WR	LF	OA	WR
Spider hair	X	X	↓	X	↓	↓	X	↓	↓	↓	↓	↓
Cactus needle	X	↓	X	X	↓	X	X	↓	↓	↓	↓	↓
Tooth	X	↓	X	X	↓	X	↓	↓	X	↓	↓	↓
Magnesium wire	X	X	X	X	X	X	↓	↓	X	↓	↓	↓

Table 1. Reconstruction artifacts that are present in the respective result panels of Fig. 5, 6, 7, 8. The truncation artifact (TA) issue was already resolved by preprocessing and is left out in this table. The abbreviations under the algorithm numbers correspond to the artifact types defined as (LF) Low frequency noise, (OA) Overshooting artifacts and (WR) Weak reconstruction. An arrow downwards ↓ indicates a significant suppression of the artifact, an X indicates that the artifact is visibly present.

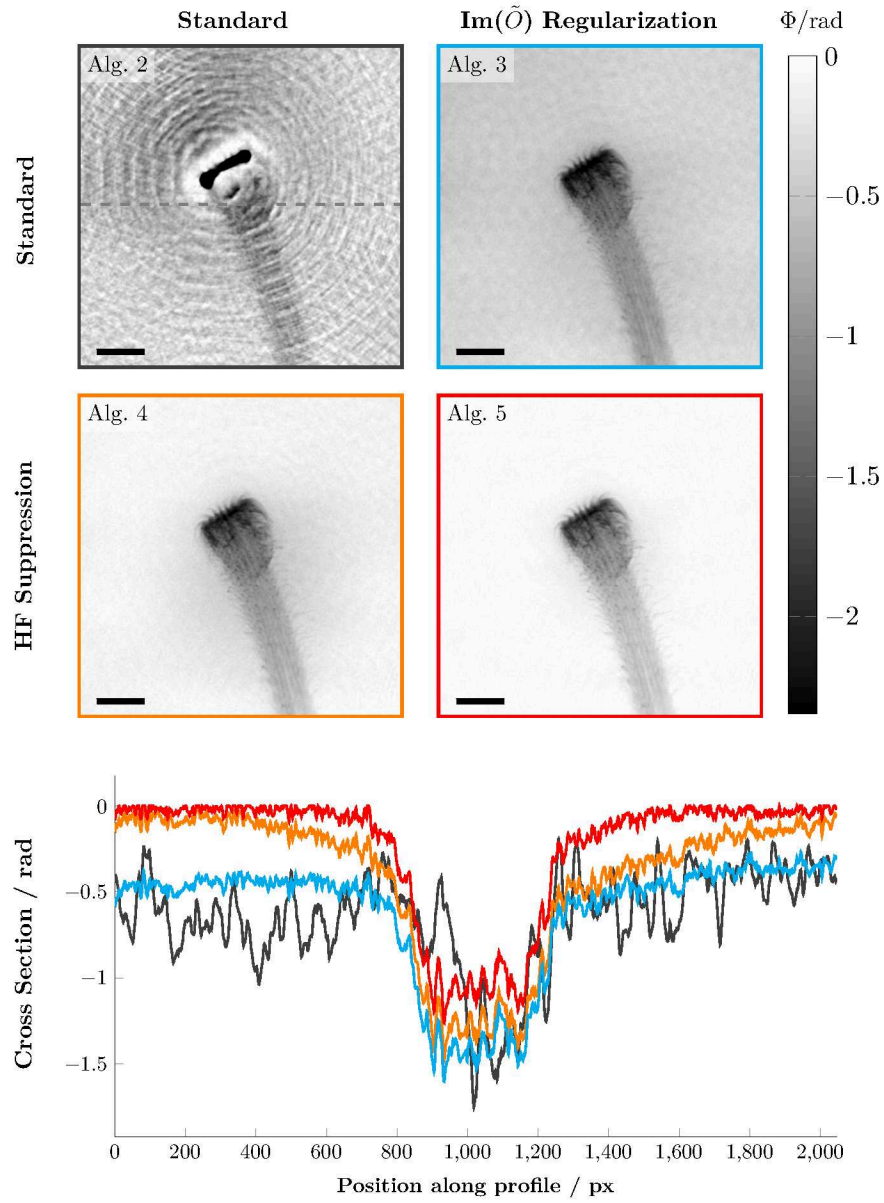


Fig. 5. Spider hair. The result of Alg. 2 shows a clearly visible overshooting artifact in the dense area in the middle that emits fringes into the background. The object itself is optically well reconstructed in panels of Alg. 3 and Alg. 4. However, both approaches still show LF artifacts across the whole phase image. The panel of Alg. 5 shows the best reconstruction. The dashed line indicates the position of the cross section. The scale bar indicates 8  $\mu\text{m}$ .

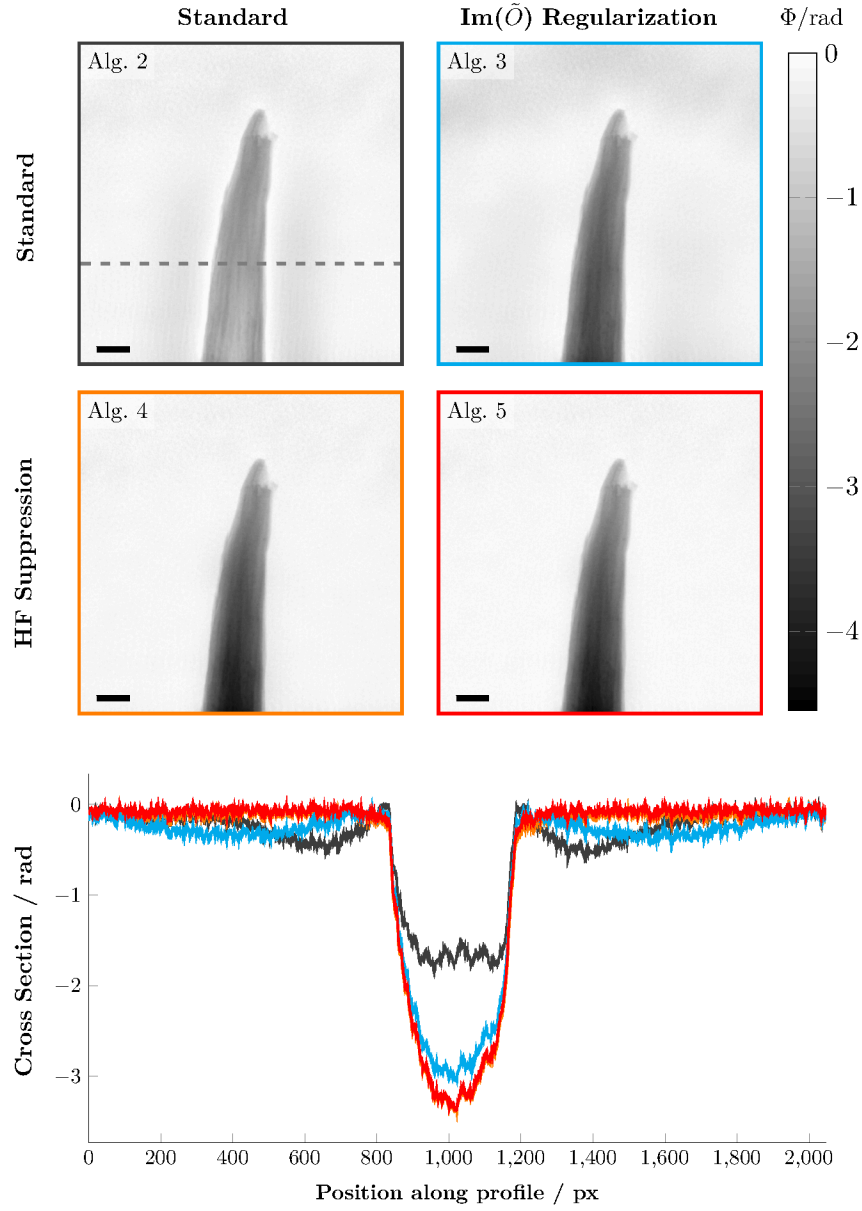


Fig. 6. Cactus Needle. The algorithm Alg. 2 has not completely reconstructed the object's interior. In the background around the object are some LF variations visible. The panel of Alg. 3 offers a better phase reconstruction but suffers from LF artifacts in the top area and LF variations in the background. The panel of Alg. 4 and Alg. 5 reconstructed the object almost identical, with significantly reduced LF variations. The dashed line indicates the position of the cross section. The scale bar indicates 20  $\mu\text{m}$ .



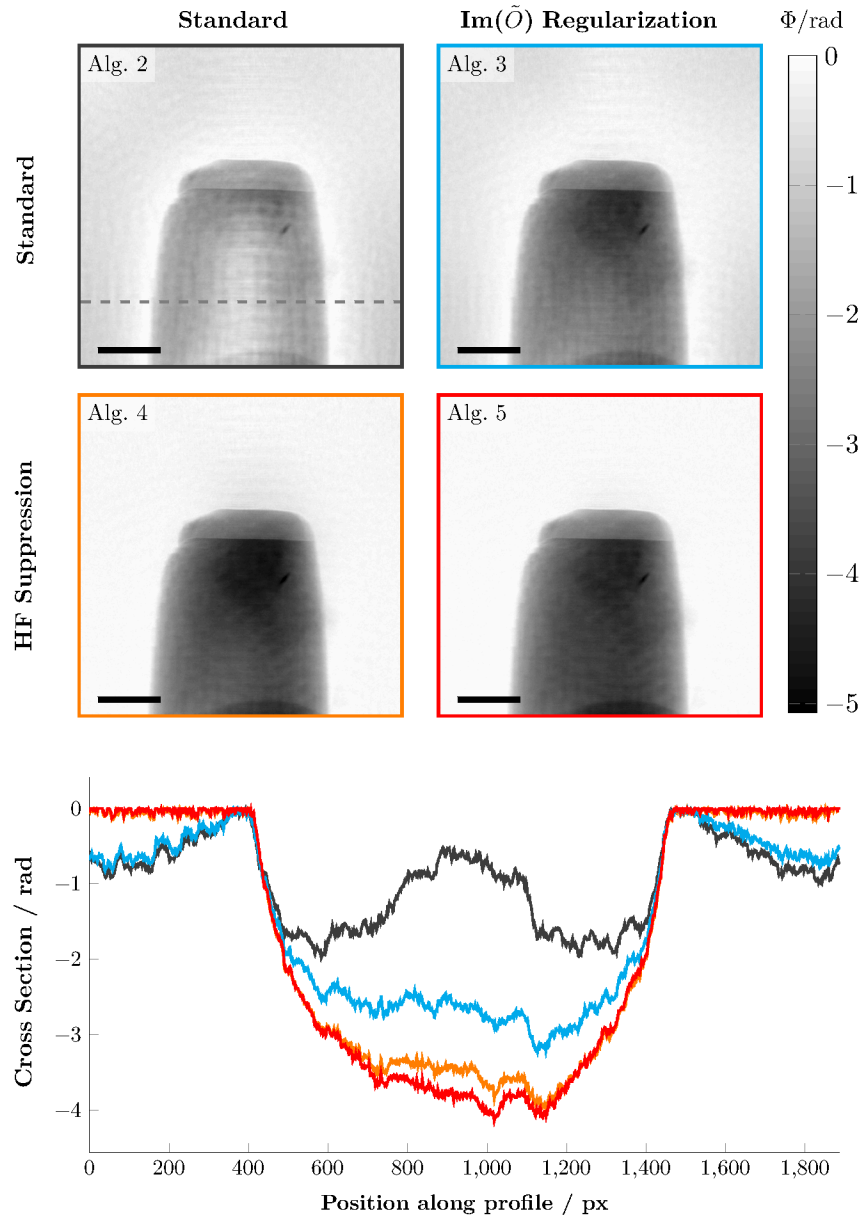


Fig. 7. Tooth. The algorithms Alg. 2 and Alg. 3 have not completely reconstructed the object's interior. In the background are some LF variations visible. In the cross section, the error increases towards the image border. Alg. 4 and Alg. 5 significantly reduce the LF variation in the cross section. Alg. 5 possesses the highest rate of maximum reconstructed phase shift to noise in the object free area. The dashed line indicates the position of the cross section. The visible dense structure at the bottom of the image belongs to the sample and is Gallium from a focused ion beam. The scale bar indicates  $10\mu\text{m}$ .

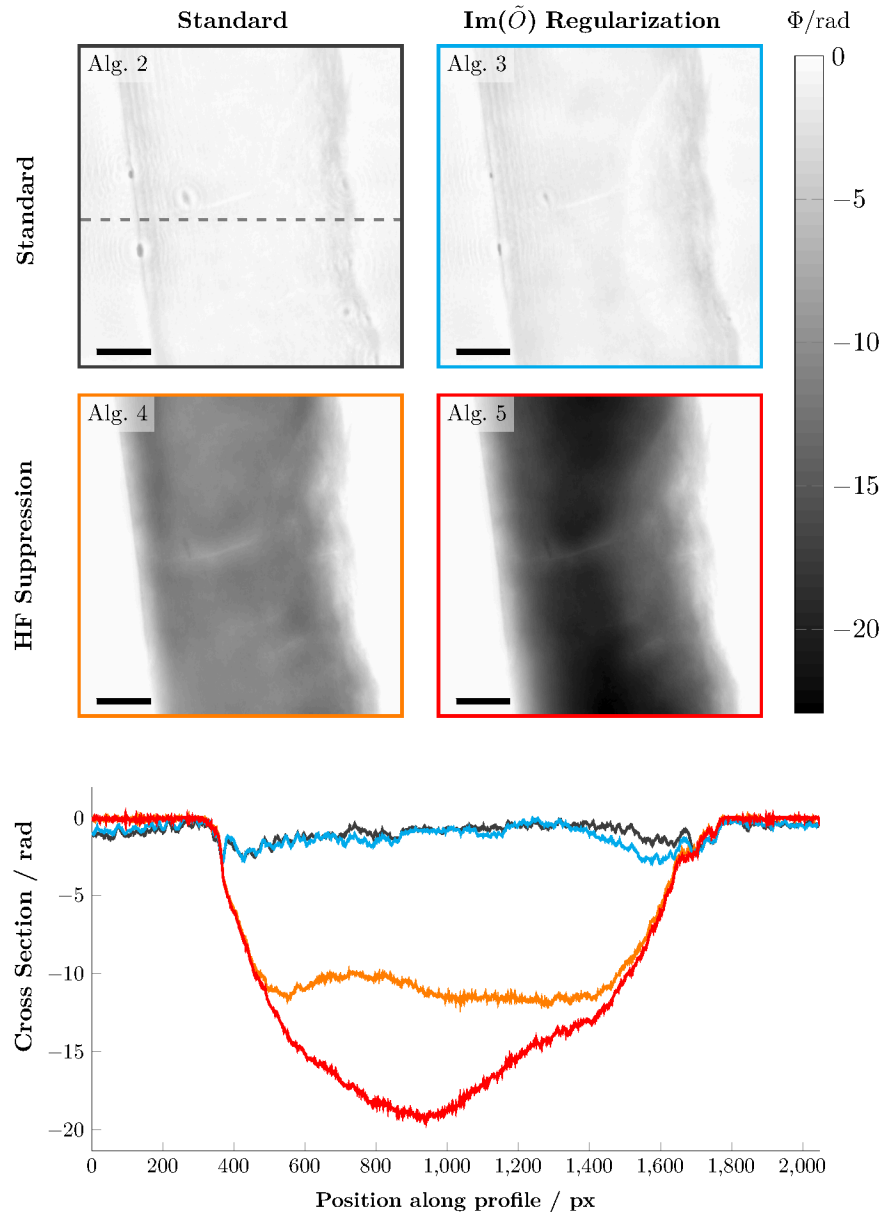


Fig. 8. Magnesium wire. Here, the effects of the regularization techniques are particularly visible. The quality of reconstructed phase values increases gradually in the panel order Alg. 2, Alg. 3, Alg. 4, Alg. 5. Alg. 3 reduced the overshooting artifacts of Alg. 2, especially in the right area of the object but was unable to reconstruct the object's interior. The result of Alg. 4 is free of overshooting artifacts but the object's interior is still not completely reconstructed. Only the combination of all approaches, Alg. 5, yields a well reconstructed phase image with the completely reconstructed object's interior and significantly reduced artifacts. The dashed line indicates the position of the cross section. The scale bar indicates  $50\mu\text{m}$ .

## 400 7. Summary

401 In the problem statement Sec. 2, we investigated the reconstruction of objects with a state-of-the-  
402 art projected gradient descent approach of [26], without a support constraint and from a single  
403 hologram. The reconstruction result suffered from various artifacts that we sorted into different  
404 categories. We found that the algorithm overshoots the reconstructed phase values during the  
405 first few hundred iterations. It could also be seen that the algorithm is had weak reconstruction  
406 capabilities of small image gradients. Due to the limited field of view of the detector, the truncated  
407 hologram also caused artifacts at the image border. In Sec. 4, we proposed methods to reduce the  
408 artifacts that were identified in Sec. 2:

- 409 i. To avoid truncation artifacts, we improved the preprocessing of measured holograms. This  
410 included the derivation of an appropriate padding scheme by mirroring the hologram into  
411 each direction and by adjusting the constant padding value with respect to the flat-field  
412 correction error. Eventually, we also applied a proper window function to the padded data.  
413 To suppress overshooting artifacts and to increase the reconstruction speed, we introduced  
414 a warm-up scheme for the projected gradient descent reconstruction. The new algorithm  
415 employs multiple regularization techniques.
- 416 ii. To suppress the overshooting, we regularized the absorption values of the reconstructed  
417 refractive object with a weighted  $L_2$  regularization.
- 418 iii. To increase the reconstruction capability of low spatial frequencies, we combined a multi-  
419 grid reconstruction with a suppression of high frequencies in the Nesterov accelerated  
420 gradient.

421 In Sec. 5, we tested our approaches on different kind of real objects that where measured at  
422 the beamline P05 at PETRA III, located at DESY in Hamburg. The objects covered interaction  
423 strengths weakly interacting samples which do not exceed phase shifts of  $2\pi$  to multi material  
424 samples that produce a phase range beyond  $6\pi$ . In general, each of our approaches show quality  
425 improvements when compared to a simple approach with a standard projected gradient descent  
426 algorithm. Our preprocessing scheme (i) successfully suppressed truncation artifacts. The  $L_2$   
427 regularization of the absorption values (ii) reduced significantly overshooting artifacts and offered  
428 an improved reconstruction of the object envelope for three of four objects. The high frequency  
429 suppressing method (iii) resulted in stronger improvements of the reconstruction quality than the  
430 regularization approach (ii). In all examples, no overshooting artifacts where visible and the low  
431 frequency artifacts were reduced. For two of four objects, this standalone approach was still not  
432 successful to reconstruct the complete object envelope. It could be seen that the results of the HF  
433 Suppression approach can be further improved by adding the  $L_2$  regularization of the absorption  
434 values. The combination of all approaches successfully suppressed artifacts in all cases with a  
435 good reconstructed object envelope.

436 In this paper, we have demonstrated that the proposed preprocessing and reconstruction scheme  
437 enables a significant suppression of artifacts in the reconstruction of phase images without a  
438 spatial support using only a single hologram. The combination of the  $L_2$  regularization for the  
439 absorption values and the high frequency suppression offers a superior reconstruction quality  
440 than each of the individual approaches. We showed that the implemented algorithm is robust  
441 with respect to a wide range of real objects, that strongly differ in their interaction strength with  
442 the X-ray illumination. We could also identify a common set of parameters that were well suited  
443 and could be generalized over the tested objects. Two object dependent parameters,  $\sigma$  for the  
444 Nesterov accelerated gradient weights and the illumination parameter  $A_0$  in the forward model  
445 still need to be explored in future work.

446 With the advent of fourth generation synchrotron radiation sources, coherent full-field imaging  
447 techniques are gaining more interest in new application areas, for example in high pressure

448 physics [54], while for established science fields, the trend is aiming for dynamic studies of  
 449 specimen, i. e. *in situ* and *in operando* studies. The latter is followed by increasing requirements  
 450 on the measurement process in terms of acquisition parameters. Consequently this increases also  
 451 the requirements on the reconstruction process. This work is a step further towards simplified  
 452 and automatized processing of large amounts of holographic data sets, which could enable an  
 453 online reconstruction to monitor and control the state of dynamic measurements.

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#### 467 Disclosure

468 The authors declare no conflicts of interest.

#### 469 Data Availability

470 Data and software underlying the results presented in this paper are available under DOI:  
 471 10.5281/zenodo.8349365.

#### 472 Appendix

##### 473 A. Algorithms

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#### **Algorithm 1:** Preprocessing of reconstruction input $\tilde{\mathcal{I}}_{\text{det}}$

---

**Input:** Detector data  $\mathcal{I}_{\text{raw}} \in \mathbb{R}^{N \times M}$ , synthetic flat field  $\mathcal{I}_{\text{flat}} \in \mathbb{R}^{N \times M}$ , Fading window width  
 $w \in \mathbb{R}$ , Source intensity  $A_0$

**Output:** Pre-processed detector data  $\tilde{\mathcal{I}}_{\text{det}}$

474  $\tilde{\mathcal{I}}_{\text{det}} = \frac{\mathcal{I}_{\text{raw}}}{\mathcal{I}_{\text{flat}}}$  ▷ Flat-field correction  
 $\tilde{\mathcal{I}}_{\text{det}} = \text{mirr}(\tilde{\mathcal{I}}_{\text{det}})$  ▷ Extend the hologram by mirroring  
 $\tilde{\mathcal{I}}_{\text{det}} = \text{pad}_{A_0}(\tilde{\mathcal{I}}_{\text{det}})$  ▷ Padding with  $A_0$   
 $\tilde{\mathcal{I}}_{\text{det}} = \text{fade}(\tilde{\mathcal{I}}_{\text{det}})$  ▷ Apply fading window  
**return**  $\tilde{\mathcal{I}}_{\text{det}}$

---

---

**Algorithm 2:** Basic reconstruction of  $\tilde{O}$ 

---

**Input:** Pre processed measurements  $\tilde{I}_{\text{det}}$ , update rate  $\eta$  for the gradient step, Nesterov momentum  $\gamma$ , filter coefficients  $\Omega_F$ , initial guess  $\tilde{O}_0$

**Output:** Approximated solution  $\tilde{O}^*$

$g_0 = 0$

► Nesterov accelerated Gradient

**repeat**

$\tilde{O}'_k = \text{gauss}_{\Omega_F}(\text{fade}(\text{mirr}(\tilde{O}_k)))$

$y_{k+1} = \tilde{O}'_k - \gamma g_k$

$g_{k+1} = \gamma g_k + \eta \nabla \left[ \frac{1}{2} \left\| |\mathcal{D}_{\text{Fr}}(y_{k+1})| - \sqrt{\tilde{I}_{\text{det}}} \right\|_2^2 \right]$

$n_{k+1} = \mathcal{P}_{\Omega_P}(\tilde{O}'_k - g_{k+1})$

$\tilde{O}_{k+1} = \text{cutfov}(n_{k+1})$

**until** *Stopping criterion reached*;

$\tilde{O}^* = \tilde{O}_{k+1}$

**return**  $\tilde{O}^*$

---

---

**Algorithm 3:** Reconstruction of  $\tilde{O}$  with  $L_2$  regularization for  $\text{Im}(\tilde{O})$ 

---

**Input:** Preprocessed measurements  $\tilde{I}_{\text{det}}$ , update rate list  $\eta$  for the gradient step, Nesterov momentum list  $\gamma$ , filter coefficients  $\Omega_F$ ,  $L_2$  regularization weight list  $\beta$ , initial guess  $\tilde{O}_0$

**Output:** Approximated solution  $\tilde{O}^*$

$v_0 = 0, j = 0$

► Nesterov accelerated gradient  $v$  and index  $j$  of parameter update loop

// Regularization levels

**repeat**

// Iterate in given level

**repeat**

$\tilde{O}'_{k+1} = \text{gauss}_{\Omega_F^j}(\text{fade}(\text{mirr}(\tilde{O}_k)))$

$y_{k+1} = \tilde{O}'_{k+1} - \gamma^j g_k$

$g_{k+1} = \gamma^j g_k + \eta^j \nabla \left[ \frac{1}{2} \left\| |\mathcal{D}_{\text{Fr}}(y_{k+1})| - \sqrt{\tilde{I}_{\text{det}}} \right\|_2^2 + \frac{\beta^j}{\eta^j} \| \text{Im}(y_{k+1}) \|_2 \right]$

$n_{k+1} = \mathcal{P}_{\Omega_P}(\tilde{O}'_{k+1} - g_{k+1})$

$\tilde{O}_{k+1} = \text{cutfov}(n_{k+1})$

**until** *Stopping criterion reached*;

$j = j + 1$

**until**  $j > j_{\text{max}}$ ;

$\tilde{O}^* = \tilde{O}_{k+1}$

**return**  $\tilde{O}^*$

---

---

**Algorithm 4:** Reconstruction of  $\tilde{O}$  with high frequencies suppression

---

**Input:** Pre processed measurements  $\mathcal{I}_{\text{det}}$ , update rate list  $\eta$  for the gradient step, Nesterov momentum operators  $\tau$ , filter coefficients  $\Omega_F$ , down-sampling operator list  $U$ , initial guess  $\tilde{O}_0$

**Output:** Approximated solution  $\tilde{O}^*$

$v_0 = 0, j = 0$  ▷ Nesterov accelerated gradient  $v$  and index  $j$  of parameter update loop

// Regularization levels

**repeat**

    // Iterate in given level

**repeat**

$$\tilde{O}'_{k+1} = \text{gauss}_{\Omega_F^j}(\text{fade}(\text{mirr}(\tilde{O}_k)))$$

$$y_{k+1} = \tilde{O}'_k - \tau^j g_k$$

$$g'_{k+1} = \tau^j g_k + \eta^j \nabla \left[ \frac{1}{2} \left\| |\mathcal{D}_{\text{Fr}}(y_{k+1})| - \sqrt{U^j \mathcal{I}_{\text{det}}} \right\|_2^2 \right]$$

$$n_{k+1} = \mathcal{P}_{\Omega_P}(\tilde{O}'_k - g'_{k+1})$$

$$\tilde{O}_{k+1} = \text{cutfov}(n_{k+1})$$

**until** *Stopping criterion reached;*

$$\tilde{O}_{k+2} = U^{j+1} (U^j)^{-1} \tilde{O}_{k+1}$$

$$g_{k+2} = U^{j+1} (U^j)^{-1} g_{k+1}$$

$$j = j + 1$$

**until**  $j > j_{\max}$ ;

$$\tilde{O}^* = \tilde{O}_{k+2}$$

**return**  $\tilde{O}^*$

---

---

**Algorithm 5:** ASRM reconstruction of  $\tilde{O}$  created from a combination of Alg. 3 and Alg. 4

---

**Input:** Pre processed measurements  $\tilde{I}_{\text{det}}$ , update rate list  $\eta$  for the gradient step, Nesterov momentum operators  $\tau$ ,  $L_2$  regularization weight list  $\beta$ , filter coefficients  $\Omega_F$ , down-sampling operator list  $U$ , initial values for  $\tilde{O}_0$

**Output:** Approximated solution  $O^*$

$v_0 = 0, j = 0$  ▷ Nesterov accelerated gradient  $v$  and index  $j$  of parameter update loop

// Regularization levels

**repeat**

    // Iterate in given level

**repeat**

$$\tilde{O}'_{k+1} = \text{gauss}_{\Omega_F^j}(\text{fade}(\text{mirr}(\tilde{O}_k)))$$

$$y_{k+1} = \tilde{O}'_k - \tau^j g_k$$

$$g'_{k+1} = \tau^j g_k + \eta^j \nabla \left[ \frac{1}{2} \left\| |\mathcal{D}_{\text{Fr}}(y_{k+1})| - \sqrt{U^j \tilde{I}_{\text{det}}} \right\|_2^2 + \frac{\beta^j}{\eta^j} \|Im(y_{k+1})\|_2 \right]$$

$$n_{k+1} = \mathcal{P}_{\Omega_P}(\tilde{O}'_k - g'_{k+1})$$

$$\tilde{O}_{k+1} = \text{cutfov}(n_{k+1})$$

**until** *Stopping criterion reached;*

$$\tilde{O}_{k+2} = U^{j+1} (U^j)^{-1} O_{k+1}$$

$$g_{k+2} = U^{j+1} (U^j)^{-1} g_{k+1}$$

$$j = j + 1$$

**until**  $j > j_{\text{max}}$ ;

$$\tilde{O}^* = \tilde{O}_{k+2}$$

**return**  $\tilde{O}^*$

---

Sample	Energy	$z_{01}$	$z_{02}$	Fr	$t$
Spider hair	11.0 keV	79.95 mm	19.661 m	$7.790 \times 10^{-5}$	1.0 s
Cactus Needle	17.0 keV	28.54 mm	19.652 m	$4.347 \times 10^{-4}$	0.8 s
Tooth	17.0 keV	81.708 mm	19.652 m	$1.231 \times 10^{-4}$	0.8 s
Magnesium Wire	11.0 keV	43.70 mm	19.661 m	$4.337 \times 10^{-4}$	1.5 s

Table 2. Parameters for the setup shown in Fig. 1, the calculated effective Fr and the exposure time  $t$ .

	FWHM for $\sigma$ in $\tau$ at Downsampling				$A_0$
	16×	4×	2×	1×	
Spider hair	16	8	64	-	0.98
Cactus needle	16	8	64	-	0.98
Tooth	16	8	64	-	0.96
Magnesium wire	4	8	32	-	1.1

Table 3. Sample dependent parameters. Values of  $\sigma$  in  $\tau$  given as full width at half magnitude (FWHM) in pixels.  $\sigma$  can be derived by calculating  $\sigma = FWHM/2.35$ . Value  $A_0$  for the constant probe model.



Panel	Iterations	$U$ Downsampling	$\eta$ Update Rate	$\gamma$ Momentum	$\beta$ $L_2$ Weight	Filter $\tilde{O}$
Alg. 2	2000	-	1.1	0.99	-	2.0 / 8.0
Alg. 3	1500	-	1.1	0.99	10.0	2.0 / 8.0
	500	-	1.1	0.99	-	2.0 / 8.0
Alg. 4	700	16×	1.1	1.0	-	2.0 / 0.0
	300	4×	1.1	1.0	-	2.0 / 8.0
	500	2×	1.1	1.0	-	2.0 / 8.0
	500	-	1.1	1.0	-	2.0 / 8.0
Alg. 5	700	16×	1.1	1.0	10.0	2.0 / 0.0
	300	4×	1.1	1.0	1.0	2.0 / 8.0
	500	2×	1.1	1.0	0.1	2.0 / 8.0
	500	-	1.1	1.0	-	2.0 / 8.0

Table 4. Setup and sample independent reconstruction parameters for algorithms 2 to 5 that were used to generate results in figures 5 to 8. The filter  $\tilde{O}$  values are given in Full Width at Half Magnitude (FWHM) for real and imaginary parts in form of  $\text{FWHM}_{\text{real}}/\text{FWHM}_{\text{imag}}$ .

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