

# Negative mode problem of false vacuum decay revisited

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We study negative modes around the Coleman–de Luccia bounce solution. The conditions for the appearance of an infinite number of negative modes do not coincide between Lagrangian and Hamiltonian formulations in the literature, and we discuss the origin of this difference in detail. We show how different choices of the variable for the fluctuation around the bounce solution affect the negative mode condition in the Hamiltonian approach, and point out that there exists a choice that gives the same negative mode condition as the Lagrangian approach.

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## I. INTRODUCTION

False vacua appear in a wide class of models in particle physics. Such vacua decay into the true vacuum via quantum tunneling within a finite lifetime. The lifetime of the false vacua gives important implications for model building and cosmology.

Coleman and Callan proposed calculating the decay width of the false vacua by utilizing the Euclidean path integral [1,2]. In this formalism, the “imaginary part” of the energy eigenvalue of the vacuum is interpreted as the decay width. The dominant contribution for the decay width in this path integral comes from the bounce solution, which gives the least action among nontrivial solutions of the equation of motion for the scalar fields. Since the bounce solution is a saddle point of the action, there exists a mode of fluctuation that makes the action smaller. Such a mode is called a *negative mode*. The existence of one and only one negative mode was proved in Ref. [3] in a wide class of models. This one negative mode plays an essential role in giving the imaginary part of the path integral.

Coleman and De Luccia (CDL) [4] extended this study to the case with gravity. They claimed that a saddle-point configuration of the Euclidean gravity action should give the dominant contribution to the decay width. However, the CDL bounce solution has a puzzling property. In the CDL case, the existence of only one negative mode is not guaranteed, and an infinite number of negative modes can exist in some situations [5]. This problem has been known as

the negative mode problem and it is not clear how we should interpret the existence of these modes. Despite quite a few studies [6–14], there is no conclusive answer to this problem yet. (For other problems in the Euclidean path integral with gravity, see, e.g., Ref. [15]. In the context of axion wormholes, see, e.g., Ref. [16].)

There has been confusion about the condition under which this infinite number of negative modes appear. One formulation to derive the condition is based on the expansion of the action around the bounce solution by Lavrelashvili, Rubakov, and Tinyakov (LRT) [5] and Lee and Weinberg (LW) [6] (see also approach I (L-I) in Ref. [7]). In addition to this, there are other formulations based on the Hamiltonian of the fluctuation by Khvedelidze, Lavrelashvili, and Tanaka (KLT) [8] (see also Gratton and Turok (GT) [9] and approach III (L-III) in Ref. [7]), and by Garriga, Montes, Sasaki, and Tanaka (GMST) [17] (see also approach II (L-II) in Ref. [7]), and these formulations give different conditions from Refs. [5,6]. Thus, the problems regarding the false vacuum decay with gravity are twofold: one is that an appropriate treatment of an infinite number of negative modes is not known (this is exactly the negative mode problem), and the other is that different conditions on the appearance of negative modes are presented in the literature. The ultimate goal is solving the former problem and this automatically includes the solution to the latter problem. Before going to this stage, we tackle the latter problem, i.e., the goal of this paper is to clarify the origin of the different conditions in the Lagrangian and Hamiltonian formulations. As we will see, we can mix a field and its conjugate momentum by a canonical coordinate transformation in the Hamiltonian formulation, and this is the reason why the literature gives different negative mode conditions.

The rest of this paper is organized as follows. In Sec. II we derive the negative mode condition using the Lagrangian formulation. In Sec. III we derive the condition using the Hamiltonian formulation. In Sec. IV we provide a summary

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and discussions. In the Appendix A we discuss the negative mode conditions in Refs. [5–9] in detail.

## II. LAGRANGIAN FORMULATION

We follow the notation in Ref. [7]. The Euclidean action for a real scalar field with gravity is given as

$$S = \int d^4x \sqrt{g} \left[ -\frac{1}{2\kappa} R + \frac{1}{2} (\partial_\mu \phi)(\partial^\mu \phi) + V(\phi) \right], \quad (2.1)$$

where  $\kappa = 8\pi G$ . In order to discuss the  $O(4)$ -symmetric background<sup>1</sup> and fluctuations around it, we parametrize the metric and the scalar field as

$$ds^2 = a(\tau)^2 [(1 + 2A(\tau))d\tau^2 + \gamma_{ij}(1 - 2\Psi(\tau))dx^i dx^j], \quad (2.2)$$

$$\phi = \varphi(\tau) + \Phi(\tau). \quad (2.3)$$

Here  $A$ ,  $\Psi$ , and  $\Phi$  are perturbations around the background, and  $\gamma_{ij}$  is the three-dimensional metric with a constant spatial curvature. We assume the latter to be positive definite in this paper. For the background field  $a$  and  $\varphi$ , we obtain the following field equations:

$$\mathcal{H}^2 - \mathcal{H}' - \mathcal{K} = \frac{\kappa}{2} \varphi'^2, \quad (2.4)$$

$$2\mathcal{H}' + \mathcal{H}^2 - \mathcal{K} = -\frac{\kappa}{2} (\varphi'^2 + 2a^2 V), \quad (2.5)$$

$$\varphi'' + 2\mathcal{H}\varphi' - a^2 \frac{\partial V}{\partial \varphi} = 0. \quad (2.6)$$

Here,  $X' \equiv dX/d\tau$  and  $\mathcal{H} \equiv a'/a$ . The curvature values  $\mathcal{K} = +1, 0, -1$  are for closed, flat, and open universes, respectively (with an appropriate unit). We take  $\mathcal{K} = 1$  for the perturbation around the CDL bounce solution. The case with  $\mathcal{K} = 0$  tells us about the zero mode of the inflaton in a flat universe. For details, see Appendix A 1.

Let us expand the Lagrangian around the background field as

$$\mathcal{L} = \mathcal{L}^{(0)} + \mathcal{L}^{(2)} + \dots, \quad (2.7)$$

where  $\mathcal{L}^{(0)}$  does not include  $A$ ,  $\Psi$ , and  $\Phi$ , and  $\mathcal{L}^{(2)}$  consists of quadratic term of them

$$\begin{aligned} \mathcal{L}^{(2)} = & \frac{a^2 \sqrt{\gamma}}{2\kappa} \left[ -6\Psi'^2 + 6\mathcal{K}\Psi^2 + \kappa \left( \Phi'^2 + a^2 \frac{\partial^2 V}{\partial \varphi^2} \Phi^2 + 6\varphi' \Psi' \Phi \right) \right. \\ & \left. - \left( 12\mathcal{H}\Psi' + 12\mathcal{K}\Psi + 2\kappa\varphi' \Phi' - 2\kappa a^2 \frac{\partial V}{\partial \varphi} \Phi \right) A - 2(\mathcal{H}' + 2\mathcal{H}^2 + \mathcal{K}) A^2 \right]. \end{aligned} \quad (2.8)$$

The action (2.1) is invariant under a general coordinate transformation. With the parametrization as in Eqs. (2.2) and (2.3), the action is invariant under a reparametrization of  $\tau$ :  $\tau \rightarrow \tau'(\tau)$ . This invariance is inherited as a gauge invariance of  $\mathcal{L}^{(2)}$ . Infinitesimal gauge transformations for  $A$ ,  $\Psi$ , and  $\Phi$  are given as

$$\delta A = \frac{\epsilon'}{a}, \quad \delta \Psi = -\frac{\epsilon}{a} \mathcal{H}, \quad \delta \Phi = \frac{\epsilon}{a} \varphi'. \quad (2.9)$$

Indeed, we can write down the action (2.8) in a gauge-invariant form:

$$\begin{aligned} \mathcal{L}^{(2)} = & a^2 \sqrt{\gamma} \left[ \frac{\varphi'^2}{2\mathcal{H}^2} \mathcal{R}'^2 - \frac{1}{2\mathcal{H}^2} \left( \frac{6\mathcal{K}^2}{\kappa} + 3\mathcal{K}\varphi'^2 \right) \mathcal{R}^2 \right. \\ & \left. - \frac{\varphi'^2}{\mathcal{H}} \mathcal{R}' \mathcal{A} - \frac{6\mathcal{K}}{\kappa} \mathcal{R} \mathcal{A} - \frac{3\mathcal{H}^2}{\kappa} \mathcal{A}^2 + \frac{1}{2} \mathcal{A}^2 \varphi'^2 \right]. \end{aligned} \quad (2.10)$$

Here  $\mathcal{R}$  and  $\mathcal{A}$  are gauge-invariant combinations:

$$\mathcal{R} = \Psi + \frac{\mathcal{H}}{\varphi'} \Phi, \quad (2.11)$$

$$\mathcal{A} = A + \frac{1}{\mathcal{H}} \Psi' + \left( \frac{\mathcal{H}'}{\mathcal{H}\varphi'} - \frac{\mathcal{H}}{\varphi'} \right) \Phi. \quad (2.12)$$

Integrating  $\mathcal{A}$  out and performing integration by parts, we get

$$\begin{aligned} \mathcal{L}^{(2)} = & a^2 \sqrt{\gamma} \left[ \frac{3\varphi'^2}{6\mathcal{H}^2 - \kappa\varphi'^2} \mathcal{R}'^2 + \frac{6\mathcal{K}}{(6\mathcal{H}^2 - \kappa\varphi'^2)^2} \right. \\ & \left. \times \left( 6\mathcal{H}^2 \varphi'^2 - 6\mathcal{K}\varphi'^2 - \kappa\varphi'^4 - 6\mathcal{H}\varphi' a^2 \frac{\partial V}{\partial \varphi} \right) \mathcal{R}^2 \right]. \end{aligned} \quad (2.13)$$

Therefore, we can directly read off the coefficient of the kinetic term of  $\mathcal{R}$  to obtain the negative mode condition:

$$\mathcal{Q} \equiv \mathcal{H}^2 - \frac{\kappa}{6} \varphi'^2. \quad (2.14)$$

<sup>1</sup>A proof of the  $O(4)$  symmetry of the bounce solution with gravity is not known, though it is widely believed. For the case without gravity, see Refs. [18–21].

The coefficient of the kinetic term of  $\mathcal{R}$  becomes negative if  $\mathcal{Q} < 0$ . This result is consistent with LRT [5], LW [6], and L-I [7].

### III. HAMILTONIAN FORMULATION

In this section we discuss the Hamiltonian formulation for the perturbations around the background field. Let us start from the Lagrangian given in Eq. (2.8). Conjugate momenta for  $\Psi$ ,  $\Phi$ , and  $A$  are defined as

$$\Pi_A \equiv \frac{\partial \mathcal{L}^{(2)}}{\partial A'} = 0, \quad (3.1)$$

$$\Pi_\Psi \equiv \frac{\partial \mathcal{L}^{(2)}}{\partial \Psi'} = \frac{6a^2\sqrt{\gamma}}{\kappa} \left( -\Psi' + \frac{\kappa}{2}\varphi'\Phi - \mathcal{H}A \right), \quad (3.2)$$

$$\Pi_\Phi \equiv \frac{\partial \mathcal{L}^{(2)}}{\partial \Phi'} = a^2\sqrt{\gamma}(\Phi' - \varphi'A). \quad (3.3)$$

Thus, the Lagrangian  $\mathcal{L}^{(2)}$  is singular and we have a primary constraint,

$$C_1 \equiv \Pi_A = 0. \quad (3.4)$$

The total Hamiltonian [22] is given as

$$H_T = H_C + v_1(\tau)C_1, \quad (3.5)$$

$$\begin{aligned} H_C = & -\frac{\kappa}{12a^2\sqrt{\gamma}}\Pi_\Psi^2 + \frac{1}{2a^2\sqrt{\gamma}}\Pi_\Phi^2 + \frac{1}{2}\kappa\varphi'\Pi_\Psi\Phi \\ & - a^2\sqrt{\gamma}\left[\frac{3\mathcal{K}}{\kappa}\Psi^2 + \frac{1}{2}\left(a^2\frac{\partial^2 V}{\partial\varphi^2} + \frac{3}{2}\kappa\varphi'^2\right)\Phi^2\right] \\ & + A\left[\varphi'\Pi_\Phi - \mathcal{H}\Pi_\Psi + a^2\sqrt{\gamma}\left(\left(3\mathcal{H}\varphi' - a^2\frac{\partial V}{\partial\varphi}\right)\Phi\right.\right. \\ & \left.\left.+ \frac{6\mathcal{K}}{\kappa}\Psi\right)\right], \end{aligned} \quad (3.6)$$

where  $v_1$  is an arbitrary function of  $\tau$ . An explicit form of  $v_1$  can be determined by a gauge fixing. The time evolution of an arbitrary variable  $f$  is given as

$$f' = \frac{\partial f}{\partial \tau} + [f, H_T], \quad (3.7)$$

where the Poisson bracket is defined as

$$[f, g] \equiv \sum_{q=A, \Psi, \Phi} \left( \frac{\partial f}{\partial q} \frac{\partial g}{\partial \Pi_q} - \frac{\partial f}{\partial \Pi_q} \frac{\partial g}{\partial q} \right). \quad (3.8)$$

Note that the partial derivative acts only on background quantities in Eq. (3.7). Since  $C'_1 = 0$  should be satisfied for consistency with  $C_1 = 0$ , we obtain a secondary constraint as

$$\begin{aligned} C_2 \equiv C'_1 = & \frac{\partial C_1}{\partial \tau} + [C_1, H_T] = -\varphi'\Pi_\Phi + \mathcal{H}\Pi_\Psi \\ & + a^2\sqrt{\gamma}\left[\left(-3\mathcal{H}\varphi' + a^2\frac{\partial V}{\partial\varphi}\right)\Phi - \frac{6\mathcal{K}}{\kappa}\Psi\right]. \end{aligned} \quad (3.9)$$

The time evolution of  $C_2$  is written in terms of  $C_2$  itself,

$$C'_2 = \frac{\partial C_2}{\partial \tau} + [C_2, H_T] = \mathcal{H}C_2, \quad (3.10)$$

and thus no tertiary constraint arises.  $C_1$  and  $C_2$  are first-class constraints because  $[C_i, C_j] = 0$  ( $i, j = 1, 2$ ) and the matrix  $[C_i, C_j]$  is not invertible.

As we have seen, the Lagrangian  $\mathcal{L}^{(2)}$  is invariant up to total derivatives under the gauge transformation (2.9). Let us discuss this gauge invariance in light of the Hamiltonian. Although Dirac conjectured that all of the first-class constraints generate gauge symmetry [22], not all of them are generators of gauge symmetry in reality [23]. The number of generators of gauge symmetry is equal to the number of primary first-class constraints [24,25], though the generators are in general linear combinations of primary and nonprimary first-class constraints. Indeed, in the present case we have only one degree of freedom  $\epsilon$  for the gauge transformation [see Eq. (2.9)].

According to Refs. [24,25], a generator of gauge symmetry  $G_\epsilon$  with an arbitrary function  $\epsilon(\tau)$  should satisfy

$$\frac{\partial G_\epsilon}{\partial \tau} + [G_\epsilon, H_T] = (\text{primary constraints}). \quad (3.11)$$

By using Eqs. (3.9) and (3.10), we can determine the solution of the above equation up to an overall factor:

$$G_\epsilon = -\frac{\epsilon}{a}C_2 + \frac{\epsilon'}{a}C_1. \quad (3.12)$$

By using this generator, we obtain the gauge transformation for perturbations as

$$[\Psi, G_\epsilon] = -\frac{\epsilon}{a}\mathcal{H}, \quad (3.13)$$

$$[\Phi, G_\epsilon] = \frac{\epsilon}{a}\varphi', \quad (3.14)$$

$$[A, G_\epsilon] = \frac{\epsilon'}{a}, \quad (3.15)$$

$$[\Pi_\Psi, G_\epsilon] = -\epsilon a\sqrt{\gamma}\frac{6\mathcal{K}}{\kappa}, \quad (3.16)$$

$$[\Pi_\Phi, G_\epsilon] = \epsilon a\sqrt{\gamma}\left(a^2\frac{\partial V}{\partial\varphi} - 3\mathcal{H}\varphi'\right) = \epsilon a\sqrt{\gamma}(\varphi'' - \varphi'\mathcal{H}), \quad (3.17)$$

$$[\Pi_A, G_\epsilon] = 0. \quad (3.18)$$

The above transformation law is consistent with Eq. (2.9) and Refs. [7,17].

As we will discuss later, gauge fixing is required when we use the Dirac bracket. Let us briefly comment on the gauge-fixing conditions in the literature. Reference [8] imposed  $\Pi_\Psi = 0$  and  $A - \Psi = 0$ , and Ref. [7] imposed  $\Pi_\Psi = 0$  and  $A = 0$  in their approach III. The time evolution of  $\Pi_\Psi$  is  $[\Pi_\Psi, H_T] = -6\sqrt{\gamma}\mathcal{K}a^2(A - \Psi)/\kappa$ . Thus, the gauge-fixing condition in Ref. [8] ( $\Pi_\Psi = 0$  and  $A - \Psi = 0$ ) is consistent, but the condition in Ref. [7] ( $\Pi_\Psi = 0$  and  $A = 0$ ) is inconsistent.

### A. Canonical transformation approach

So far, all of the variables have been gauge variant in the current coordinate. Let us move to a new coordinate system with gauge-invariant variables. We can see that there exist three independent gauge-invariant linear combinations of  $\Psi$ ,  $\Phi$ ,  $\Pi_\Psi$ , and  $\Pi_\Phi$ .  $C_2$  is one of such linear combinations, and it is convenient to adopt constraints as coordinates, as discussed in Ref. [26]. Another convenient gauge-invariant combination is

$$\mathcal{R} \equiv \Psi + \frac{\mathcal{H}}{\varphi'}\Phi. \quad (3.19)$$

In addition to these,  $\mathcal{R}' = \partial\mathcal{R}/\partial\tau + [\mathcal{R}, H_T]$  is also a gauge-invariant variable, which is given by the original variables as

$$\mathcal{R}' = \left( \frac{3\mathcal{H}^2 - \mathcal{K}}{\varphi'} - \frac{a^2\mathcal{H}}{\varphi'^2} \frac{\partial V}{\partial\varphi} \right) \Phi + \frac{1}{\varphi'} \frac{\mathcal{H}}{a^2\sqrt{\gamma}} \Pi_\Phi - \frac{\kappa}{6a^2\sqrt{\gamma}} \Pi_\Psi. \quad (3.20)$$

Let us move to a coordinate system that has  $\mathcal{R}$  and  $C_2$  as canonical coordinates. We define their conjugate momenta as

$$\Pi_{\mathcal{R}} \equiv \frac{a^2\sqrt{\gamma}\varphi'^2}{Q} \mathcal{R}', \quad (3.21)$$

$$\Pi_{C_2} \equiv -\frac{1}{Q} \left( \mathcal{H}\Psi + \frac{\kappa}{6}\varphi'\Phi \right) + f(\tau)C_2. \quad (3.22)$$

Here  $f$  is an arbitrary function of  $\tau$ , and  $Q$  is defined in Eq. (2.14). By using the above definition, we can check that  $\mathcal{R}$ ,  $\Pi_{\mathcal{R}}$ ,  $C_2$ , and  $\Pi_{C_2}$  satisfy

$$[\mathcal{R}, \Pi_{\mathcal{R}}] = [C_2, \Pi_{C_2}] = 1, \quad (3.23)$$

$$[\mathcal{R}, C_2] = [\Pi_{\mathcal{R}}, C_2] = [\mathcal{R}, \Pi_{C_2}] = [\Pi_{\mathcal{R}}, \Pi_{C_2}] = 0, \quad (3.24)$$

$$[\mathcal{R}, A] = [\Pi_{\mathcal{R}}, A] = [C_2, A] = [\Pi_{C_2}, A] = 0, \quad (3.25)$$

$$[\mathcal{R}, \Pi_A] = [\Pi_{\mathcal{R}}, \Pi_A] = [C_2, \Pi_A] = [\Pi_{C_2}, \Pi_A] = 0. \quad (3.26)$$

Thus, we can use  $(\mathcal{R}, C_2, A, \Pi_{\mathcal{R}}, \Pi_{C_2}, \Pi_A)$  as canonical coordinates. The Hamiltonian in the new coordinates can be derived from a (type-three) generating function  $W = W(\Pi_\Psi, \Pi_\Phi, \mathcal{R}, C_2)$ . Taking the generating function as

$$W = \frac{1}{a^2\sqrt{\gamma}(3\kappa\mathcal{H}\varphi'^2 - 6\mathcal{K}\mathcal{H} - a^2\kappa\varphi'(\partial V/\partial\varphi))} \times \left[ \frac{1}{2}\kappa\mathcal{H}^2\Pi_\Psi^2 + \frac{1}{2}\kappa\varphi'^2\Pi_\Phi^2 - \kappa\mathcal{H}\varphi'\Pi_\Phi\Pi_\Psi + (6a^2\sqrt{\gamma}\mathcal{K}\varphi'\mathcal{R} + \kappa\varphi'C_2)\Pi_\Phi + (a^2\sqrt{\gamma}\kappa(a^2\varphi'(\partial V/\partial\varphi) - 3\mathcal{H}\varphi'^2)\mathcal{R} - \kappa\mathcal{H}C_2)\Pi_\Psi + \frac{1}{2}\kappa C_2^2 + \frac{a^2\sqrt{\gamma}[(3\mathcal{H}^2 - \mathcal{K})\varphi'^2 - a^2\mathcal{H}\varphi'(\partial V/\partial\varphi)]}{Q} (3a^2\sqrt{\gamma}\mathcal{K}\mathcal{R}^2 + \kappa\mathcal{R}C_2) \right] - \frac{1}{2}fC_2^2, \quad (3.27)$$

we can reproduce the transformations via

$$\Psi = -\frac{\partial W}{\partial\Pi_\Psi}, \quad \Phi = -\frac{\partial W}{\partial\Pi_\Phi}, \quad \Pi_{\mathcal{R}} = -\frac{\partial W}{\partial\mathcal{R}}, \quad \Pi_{C_2} = -\frac{\partial W}{\partial C_2}. \quad (3.28)$$

The new (total) Hamiltonian becomes

$$\begin{aligned} K_T &= H_T + \frac{\partial W}{\partial\tau} \Big|_{\Pi_\Psi=\Pi_\Psi(\mathcal{R}, C_2, \Pi_{\mathcal{R}}, \Pi_{C_2}), \Pi_\Phi=\Pi_\Phi(\mathcal{R}, C_2, \Pi_{\mathcal{R}}, \Pi_{C_2})} \\ &= \frac{1}{2} \frac{Q}{a^2\sqrt{\gamma}\varphi'^2} \Pi_{\mathcal{R}}^2 + a^2\sqrt{\gamma}\mathcal{K} \left( -\frac{\varphi'^2}{Q} + \frac{\mathcal{K}\varphi'^2 + a^2\mathcal{H}\varphi'(\partial V/\partial\varphi)}{Q^2} \right) \mathcal{R}^2 \\ &\quad + \mathcal{H}\Pi_{C_2}C_2 - \left( \frac{\kappa}{12a^2\sqrt{\gamma}Q} + f\mathcal{H} + \frac{1}{2}f' \right) C_2^2 + \left( -\frac{\kappa\varphi'^2}{2Q} + \frac{\kappa\mathcal{K}\varphi'^2 + a^2\mathcal{H}\varphi'(\partial V/\partial\varphi)}{3Q^2} \right) \mathcal{R}C_2 \\ &\quad - AC_2 + v_1\Pi_A. \end{aligned} \quad (3.29)$$

The Hamiltonian  $K_T$  shows that the  $\mathcal{R}$  sector and the  $C_2 - A$  sector are coupled only through the  $\mathcal{R}C_2$  term. Indeed, the Hamilton equations are

$$\mathcal{R}' = \frac{\partial K_T}{\partial \Pi_{\mathcal{R}}} = \frac{Q}{a^2 \sqrt{\gamma} \varphi'^2} \Pi_{\mathcal{R}}, \quad (3.30)$$

$$C_2' = \frac{\partial K_T}{\partial \Pi_{C_2}} = \mathcal{H} C_2, \quad (3.31)$$

$$A' = \frac{\partial K_T}{\partial \Pi_A} = v_1, \quad (3.32)$$

$$\begin{aligned} \Pi_{\mathcal{R}}' &= -\frac{\partial K_T}{\partial \mathcal{R}} \\ &= -2\mathcal{K}a^2 \sqrt{\gamma} \left( -\frac{\varphi'^2}{Q} + \frac{\mathcal{K}\varphi'^2 + a^2 \mathcal{H}\varphi'(\partial V/\partial \varphi)}{Q^2} \right) \mathcal{R}, \end{aligned} \quad (3.33)$$

$$\begin{aligned} \Pi_{C_2}' &= -\frac{\partial K_T}{\partial C_2} = -\mathcal{H}\Pi_{C_2} + \left( \frac{\kappa}{6a^2 \sqrt{\gamma}} \frac{1}{Q} + 2f\mathcal{H} + f' \right) C_2 \\ &\quad - \left( -\frac{\kappa \varphi'^2}{2Q} + \frac{\kappa \mathcal{K}\varphi'^2 + a^2 \mathcal{H}\varphi'(\partial V/\partial \varphi)}{3Q^2} \right) \mathcal{R} + A, \end{aligned} \quad (3.34)$$

$$\Pi_A' = -\frac{\partial K_T}{\partial A} = C_2. \quad (3.35)$$

As long as the constraints are satisfied ( $\Pi_A = C_2 = 0$ , which we choose as the boundary condition), the Hamilton equations for  $\mathcal{R}$  and  $\Pi_{\mathcal{R}}$  are not affected. Thus, we can define the following reduced Hamiltonian:

$$\begin{aligned} H_R &= \frac{1}{2} \frac{Q}{a^2 \sqrt{\gamma} \varphi'^2} \Pi_{\mathcal{R}}^2 + \sqrt{\gamma} \mathcal{K} a^2 \\ &\quad \times \left( -\frac{\varphi'^2}{Q} + \frac{\mathcal{K}\varphi'^2 + a^2 \mathcal{H}\varphi'(\partial V/\partial \varphi)}{Q^2} \right) \mathcal{R}^2, \end{aligned} \quad (3.36)$$

with the subscript  $R$  denoting “reduced.” The above Hamiltonian shows that the sign of the kinetic term is determined by the sign of  $Q$ , and thus we obtain the same condition for the negative mode condition as in Sec. II. This Hamiltonian is the same as the one directly constructed from the gauge-invariant Lagrangian (2.13).

### B. Dirac bracket approach

Let us take another approach which utilizes the Dirac bracket [22]. We impose a gauge-fixing condition  $\chi_1 = 0$  such that  $[\chi_1, C_1] = 0$  and  $[\chi_1, C_2] \neq 0$  are satisfied [27]. For consistency, we have to impose  $\chi_2 \equiv \partial \chi_1 / \partial \tau + [\chi_1, H_T] = 0$ . No other conditions appear because  $\partial \chi_2 / \partial \tau + [\chi_2, H_T] = 0$  can be maintained by choosing an appropriate  $v_1(\tau)$  in  $H_T$ . The Dirac bracket is defined as

$$\begin{aligned} [\alpha, \beta]_D &\equiv [\alpha, \beta] - \sum_{i,j=1}^4 [\alpha, \phi_i] M_{ij}^{-1} [\phi_j, \beta], \\ M_{ij} &= [\phi_i, \phi_j], \end{aligned} \quad (3.37)$$

where  $(\phi_1, \phi_2, \phi_3, \phi_4) = (C_1, C_2, \chi_1, \chi_2)$ . The time evolution of an arbitrary variable  $\alpha$  (see, e.g., Ref. [28]) is

$$\alpha' = \frac{\partial \alpha}{\partial \tau} - \sum_{i,j=1}^4 \frac{\partial \phi_i}{\partial \tau} M_{ij}^{-1} [\phi_j, \alpha] + [\alpha, H_T]_D. \quad (3.38)$$

As we have discussed so far, an infinite number of negative modes appear when the kinetic term has the wrong sign. Let us briefly discuss how we can extract the coefficient of the kinetic term, using the following Hamiltonian with time-dependent coefficients  $m(t)$ ,  $\omega(t)$ , and  $a(t)$ :

$$H = \frac{p^2}{2m} + \frac{1}{2} m \omega^2 q^2 + a p q. \quad (3.39)$$

The coefficient of the kinetic term can be extracted using the Poisson bracket  $[q, dq/dt]$  for the coordinate  $q$ :

$$\left[ q, \frac{dq}{dt} \right] = \left[ q, \frac{p}{m} + a q \right] = \frac{1}{m}. \quad (3.40)$$

Note that this is essentially the inverse of the coefficient of  $\dot{q}^2$  when we Legendre transform back to the Lagrangian formulation,

$$\begin{aligned} L &= [p\dot{q} - H]_{p=p(q,\dot{q})} = \frac{m}{2} \dot{q}^2 - m a q \dot{q} \\ &\quad + \frac{m}{2} (a^2 - \omega^2) q^2, \end{aligned} \quad (3.41)$$

and thus we can discuss the negative mode condition using  $[q, dq/dt]$ .

Note that the coefficient of the kinetic term changes if we use a different coordinate. Indeed, the Poisson bracket for a general linear combination  $Q \equiv c_q q + c_p p$  is given as

$$\begin{aligned} \left[ Q, \frac{dQ}{dt} \right] &= \left[ Q, \frac{\partial Q}{\partial t} + [Q, H] \right] = \frac{c_q^2}{m} + c_p^2 m \omega^2 \\ &\quad + c_q \frac{dc_p}{dt} - \frac{dc_q}{dt} c_p. \end{aligned} \quad (3.42)$$

This indicates that, once we canonically transform from  $(p, q)$  to  $(P, Q)$  and then Legendre transform back to the Lagrangian formulation, the coefficient of the kinetic term takes a different form. Thus, the choice of the coordinate  $q$  is essential when we discuss the sign of the kinetic term. The same argument also holds for the Dirac bracket. Therefore, we can calculate the coefficient of the kinetic term for the fluctuation using the Dirac bracket  $[q, q']_D$  once we specify



the fluctuation  $q$ . We discuss explicit examples related to this point in the following subsections and in Appendix A.

### 1. Comparison with LRT, LW, and L-I

Inspired by the Lagrangian approach, we discuss the kinetic term of  $\xi$  defined as

$$\xi \equiv g(\tau)(\varphi'\Psi + \mathcal{H}(\Phi)), \quad (3.43)$$

where  $g(\tau)$  is an arbitrary function of  $\tau$ . Here we take a generic gauge-fixing condition

$$\chi_1 = c_\Phi(\tau)\Phi + c_\Psi(\tau)\Psi + c_{\Pi_\Phi}(\tau)\Pi_\Phi + c_{\Pi_\Psi}(\tau)\Pi_\Psi. \quad (3.44)$$

For consistency, we also need to impose  $\chi_2 \equiv \partial\chi_1/\partial\tau + [\chi_1, H_T] = 0$ .  $\chi_2$  is given as

$$\begin{aligned} \chi_2 = & c_\Phi \left( \frac{\Pi_\Phi}{a^2\sqrt{\gamma}} + \varphi'A \right) + c_\Psi \left( -\frac{\kappa}{6a^2\sqrt{\gamma}}\Pi_\Psi + \frac{1}{2}\kappa\varphi'\Phi - \mathcal{H}A \right) \\ & + c_{\Pi_\Phi} \left[ -\frac{1}{2}\kappa\varphi'\Pi_\Psi + a^2\sqrt{\gamma} \left( a^2 \frac{\partial^2 V}{\partial\varphi^2} + \frac{3}{2}\kappa\varphi'^2 \right) \Phi - a^2\sqrt{\gamma} \left( 3\varphi'\mathcal{H} - a^2 \frac{\partial V}{\partial\varphi} \right) A \right] \\ & + c_{\Pi_\Psi} \frac{6\mathcal{K}}{\kappa} a^2\sqrt{\gamma}(\Psi - A) + c'_\Phi\Phi + c'_\Psi\Psi + c'_{\Pi_\Phi}\Pi_\Phi + c'_{\Pi_\Psi}\Pi_\Psi. \end{aligned} \quad (3.45)$$

By using Eqs. (3.37) and (3.38), a straightforward calculation gives

$$[\xi, \xi']_D = \frac{g^2 Q}{a^2\sqrt{\gamma}}. \quad (3.46)$$

An infinite number of negative modes of  $\xi$  appear if the above expression is negative, and thus we obtain the same condition for the negative mode condition as in Sec. II, LRT [5], LW [6], and L-I [7].

### 2. Comparison with KLT, GT, and L-III

Next we discuss the result from KLT [8] and GT [9]. Both papers discussed the kinetic term of  $\Phi$  in  $\Pi_\Psi = 0$  gauge. Thus, we take  $\chi_1 = \Pi_\Psi$  and impose  $\chi_2 = A - \Psi$  for consistency. Again, by using Eqs. (3.37) and (3.38), we obtain

$$[\Phi, \Phi']_D = \frac{1}{a^2\sqrt{\gamma}} \left( 1 - \frac{\kappa\varphi'^2}{6\mathcal{K}} \right). \quad (3.47)$$

This result is consistent with the result obtained in KLT and GT (see also L-III [7]).

The above discussion relies on the specific choice of the gauge-fixing condition. Let us discuss this result in a general gauge. The gauge-invariant combination equivalent to  $\Phi$  in  $\Pi_\Psi = 0$  gauge is

$$\hat{\Phi} \equiv \Phi + \frac{\kappa\varphi'}{6\sqrt{\gamma}a^2\mathcal{K}}\Pi_\Psi. \quad (3.48)$$

For a generic gauge condition  $\chi_1 = c_\Phi(\tau)\Phi + c_\Psi(\tau)\Psi + c_{\Pi_\Phi}(\tau)\Pi_\Phi + c_{\Pi_\Psi}(\tau)\Pi_\Psi$ , with  $\chi_2$  given in Eq. (3.45), a straightforward calculation shows that

$$[\hat{\Phi}, \hat{\Phi}']_D = \frac{1}{a^2\sqrt{\gamma}} \left( 1 - \frac{\kappa\varphi'^2}{6\mathcal{K}} \right). \quad (3.49)$$

The signs of the kinetic terms of  $\xi$  and  $\hat{\Phi}$  do not coincide in general. We have three independent gauge-invariant variables:  $\mathcal{R}$ ,  $\Pi_{\mathcal{R}}$ , and  $C_2$ . Thus,  $\hat{\Phi}$  can be written as a linear combination of them. By using Eqs. (3.19), (3.20), (3.21), and (3.9), it is straightforward to show the following relation:

$$\hat{\Phi} = \frac{\mathcal{H}\varphi'}{Q}\mathcal{R} + \frac{\kappa\varphi'}{6\sqrt{\gamma}a^2\mathcal{K}}\Pi_{\mathcal{R}} + \frac{\kappa\varphi'\mathcal{H}}{6\sqrt{\gamma}a^2\mathcal{K}Q}C_2. \quad (3.50)$$

Also,  $\xi$  discussed in Sec. III B 1 is proportional to  $\mathcal{R}$ . Thus,  $\Phi$  in KLT can be understood as a mixture of a gauge-invariant fluctuation and its conjugate momentum in LRT and LW. This explains the difference between  $[\xi, \xi']_D$  given in Eq. (3.46) and  $[\hat{\Phi}, \hat{\Phi}']_D$  given in Eq. (3.49). As we have discussed in Eq. (3.42), a canonical coordinate transformation that mixes a coordinate and its conjugate momentum changes the coefficient of the kinetic term. We can obtain the Hamiltonian for  $\hat{\Phi}$  and its conjugate momentum from the Hamiltonian given in Eq. (3.36) by a canonical coordinate transformation from  $\mathcal{R}$  and  $\Pi_{\mathcal{R}}$ . We can see that the sign of the coefficient of the kinetic term in the new coordinate is the same as the sign of Eq. (3.47). This is the reason why the negative mode condition in KLT is different from that in LRT and LW.

### 3. Comparison with GMST and L-II

Finally, we discuss L-II [7], which relies on the results given in Appendix B of Refs. [17,29], although the negative mode problem was not directly discussed there. The negative mode condition which can be read off from this literature can be understood as the kinetic term of  $\Psi$  in  $\Phi = 0$  gauge (although the authors used gauge-invariant counterparts  $\Psi$

and  $\Pi_\Psi$ ; see Appendix A 3). Therefore, we take  $\chi_1 = \Phi$  and  $\chi_2 = \chi'_1 = \Pi_\Phi/a^2\sqrt{\gamma} + \varphi'A$ . By using the Dirac bracket given in Eq. (3.37), we obtain

$$[\Psi, \Pi_\Psi]_D = 1, \quad (3.51)$$

and

$$[\Psi, [\Psi, H_T]_D]_D = \frac{\mathcal{H}^2}{a^2\sqrt{\gamma}\varphi'^2} \left(1 - \frac{\kappa\varphi'^2}{6\mathcal{H}^2}\right), \quad (3.52)$$

$$[\Psi, [\Pi_\Psi, H_T]_D]_D = \frac{6\mathcal{K}\mathcal{H}}{\kappa\varphi'^2}, \quad (3.53)$$

$$[\Pi_\Psi, [\Pi_\Psi, H_T]_D]_D = \frac{6a^2\sqrt{\gamma}\mathcal{K}}{\kappa} \frac{6\mathcal{K} - \kappa\varphi'^2}{\kappa\varphi'^2}, \quad (3.54)$$

thus reproducing the coefficients of their Hamiltonian [see Eq. (A30)]. Note that we started with the Euclidean action (2.1), while Appendix B of Ref. [17] started with the Lorentzian action. Note also that here we consider the  $O(4)$ -invariant mode only. The coefficient of  $\Pi_\Psi^2$  can be read off as

$$[\Psi, \Psi']_D = \frac{\mathcal{H}^2}{a^2\sqrt{\gamma}\varphi'^2} \left(1 - \frac{\kappa\varphi'^2}{6\mathcal{H}^2}\right), \quad (3.55)$$

which is consistent with Eq. (B12) of Ref. [17]. The authors further transformed the Hamiltonian using the variables  $\tilde{p}$  and  $\tilde{q}$  defined as

$$\Psi = \frac{\kappa\varphi'}{4}\tilde{q} + \frac{\mathcal{H}}{3a^2\sqrt{\gamma}\mathcal{K}\varphi'}\tilde{p}, \quad (3.56)$$

$$\Pi_\Psi = -\frac{3a^2\sqrt{\gamma}\mathcal{K}\varphi'}{2\mathcal{H}}\tilde{q} + \frac{2}{\kappa\varphi'}\tilde{p}. \quad (3.57)$$

The coefficient of  $\tilde{p}^2$  can be read off from the Dirac bracket with the same gauge-fixing condition as

$$[\tilde{q}, \tilde{q}']_D = \frac{1}{3a^2\sqrt{\gamma}\mathcal{K}}. \quad (3.58)$$

This is consistent with Eq. (B18) of Ref. [17] [see also Eq. (A35)]. In this case, the sign of the kinetic term for the  $O(4)$ -symmetric fluctuation is positive definite and a negative mode does not appear. Note that the Euclidean action given in Eq. (12) in Ref. [30] and Eq. (38) in Ref. [7] has the opposite sign for  $\mathcal{K} = 1$ . For details, see Appendix A 3. Reference [7] used the analytic continuation  $\tilde{q} \rightarrow -i\tilde{q}$  to obtain a positive kinetic term.

#### IV. SUMMARY AND DISCUSSIONS

It has been known that an infinite number of negative modes can appear in the fluctuations around the CDL bounce solution, and this negative mode problem [5] obscures physical interpretations of the CDL bounce solution. In this paper, we discussed the negative mode condition using different approaches.

In the Lagrangian formalism discussed in Sec. II, there is a clear difference between a field and its first derivative. We can see that  $\mathcal{R} = \Psi + (\mathcal{H}/\varphi')\Phi$  is the only choice for the gauge-invariant dynamical variable for  $O(4)$ -symmetric fluctuations, and its kinetic term becomes negative if

$$\mathcal{H}^2 - \frac{\kappa}{6}\varphi'^2 < 0, \quad (4.1)$$

as derived in Refs. [5,6].

On the other hand, in the Hamiltonian formalism discussed in Sec. III, there is no clear difference between a field and its conjugate momentum because we can mix them by a canonical coordinate transformation. Hence, we need to specify a fluctuation variable to discuss the sign of its kinetic term. If we choose  $\mathcal{R}$ , we obtain a result consistent with that in Refs. [5,6]. The negative mode condition is given by Eq. (4.1). However, if we choose a different variable, we obtain a different negative mode condition. By choosing  $\hat{\Phi} = \Phi + (\kappa\varphi'/6\sqrt{\gamma}a^2\mathcal{K})\Pi_\Psi$ , we obtain the same result as in Ref. [8]. The kinetic term of  $\hat{\Phi}$  becomes negative if

$$1 - \frac{\kappa}{6\mathcal{K}}\varphi'^2 < 0. \quad (4.2)$$

References [17,30], used  $\tilde{q}$  (and  $\tilde{p}$ ) defined in Eqs. (3.56) and (3.57), and discussed the sign of the kinetic term. As shown in Ref. [30], when we take  $\tilde{q}$  as the fluctuation variable, the sign of the kinetic term is positive definite [17]. However, the signs of the kinetic terms for an  $O(4)$ -symmetric mode and a high- $\ell$  mode are opposite [17,30], and this requires analytic continuation such as  $\tilde{q} \rightarrow -i\tilde{q}$  only for a specific partial wave(s).

We cannot give a conclusive answer to the question of which condition to use for the negative mode problem. This unclear situation stems from the fact that we do not have a concrete formalism to treat quantum gravity. The action of the Euclidean gravity is unbounded below [31] and this is a serious obstacle when we discuss non-perturbative property of the Euclidean gravity. The negative mode problem could be regarded as a part of this problem. As discussed in the introduction of Ref. [9] (see also Ref. [10]), the negative mode problem could be (partly) solved by taking a new Lagrangian obtained from the Hamiltonian after a canonical coordinate transformation. Although the original and new Lagrangians give consistent equations of motion, their value is different at the point in the phase space which does not satisfy the equations of motion. A flaw of this treatment is the fact that there is no clear criteria for the choice of the

fluctuation variable. In order to avoid these subtle theoretical issues, the Wheeler-DeWitt equation would be another interesting tool to analyze vacuum decay with gravity [32].

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### APPENDIX A: COMPARISON WITH THE LITERATURE

In this Appendix we review results derived in the literature. We also give a comparison with the inflationary quadratic action. In the main text we discuss the  $O(4)$ -symmetric case only, but in this Appendix A 1, we do not assume  $O(4)$  symmetry of fluctuation in order to compare its action with inflationary quadratic action.

#### 1. Comparison with inflationary quadratic action

Although our main interest is the negative modes around the bounce solution, we can cross-check our action with the

single-field inflationary quadratic action since both actions involve one scalar field and Einstein-Hilbert gravity. We expand the action

$$S = \int d^4x \sqrt{\mp g} \left[ -\frac{1}{2\kappa} R + \frac{1}{2} (\partial_\mu \phi)(\partial^\mu \phi) \mp V(\phi) \right], \quad (\text{A1})$$

where the upper (lower) signs are for Lorenzian (Euclidean) metric, with respect to the following perturbations:

$$ds^2 = a^2 [\mp (1 + 2A) d\tau^2 + 2B_{;i} d\tau dx^i + (\gamma_{ij}(1 - 2\Psi) + 2E_{;ij}) dx^i dx^j], \quad (\text{A2})$$

$$\phi = \varphi + \Phi. \quad (\text{A3})$$

Here the semicolon denotes the covariant derivative with respect to the spatial metric  $\gamma_{ij}$  with constant curvature. We assume  $\gamma_{ij}$  to be positive definite. The resulting quadratic action is [17]

$$S^{(2)} = \int d\tau d^3x \mathcal{L}^{(2)}, \quad (\text{A4})$$

with

$$\begin{aligned} \mathcal{L}^{(2)} = & \frac{a^2 \sqrt{\gamma}}{2\kappa} \left[ -6\Psi'^2 \mp 6\mathcal{K}\Psi^2 \mp 2\Psi\Delta\Psi + \kappa \left( \Phi'^2 \mp a^2 \frac{\partial^2 V}{\partial \varphi^2} \Phi^2 + 6\varphi' \Psi' \Phi \pm \Phi \Delta \Phi \right) \right. \\ & - \left( 12\mathcal{H}\Psi' \mp 12\mathcal{K}\Psi + 2\kappa\varphi' \Phi' \pm 2\kappa a^2 \frac{\partial V}{\partial \varphi} \Phi \mp 4\Delta\Psi \right) A \\ & - (2\mathcal{H}' + 4\mathcal{H}^2 \mp 2\mathcal{K}) A^2 \\ & \left. + \Delta(B - E')(2\kappa\varphi' \Phi - 4\Psi') + 2\mathcal{K}(B - E')\Delta(B - E') - 4\mathcal{H}A\Delta(B - E') \right]. \end{aligned} \quad (\text{A5})$$

The background equations of motion are given by

$$\mathcal{H}^2 - \mathcal{H}' \pm \mathcal{K} = \frac{\kappa}{2} \varphi'^2, \quad (\text{A6})$$

$$2\mathcal{H}' + \mathcal{H}^2 \pm \mathcal{K} = \frac{\kappa}{2} (-\varphi'^2 \pm 2a^2 V), \quad (\text{A7})$$

$$\varphi'' + 2\mathcal{H}\varphi' \pm a^2 \frac{\partial V}{\partial \varphi} = 0. \quad (\text{A8})$$

Below we eliminate the auxiliary fields  $A$  and  $B - E'$ . However, before doing so, we rewrite the action in a gauge-invariant form:

$$\begin{aligned} \mathcal{L}^{(2)} = & a^2 \sqrt{\gamma} \left[ \frac{\varphi'^2}{2\mathcal{H}^2} \mathcal{R}'^2 + \frac{1}{2\mathcal{H}^2} \left( -\frac{6\mathcal{K}^2}{\kappa} \pm (3\mathcal{K} + \Delta)\varphi'^2 \right) \mathcal{R}^2 \right. \\ & - \frac{\varphi'^2}{\mathcal{H}} \mathcal{R}' \mathcal{A} \pm \frac{6\mathcal{K}}{\kappa} \mathcal{R} \mathcal{A} \pm \frac{2\mathcal{K}\Delta}{\kappa\mathcal{H}} \mathcal{R} \mathcal{B} - \frac{3\mathcal{H}^2}{\kappa} \mathcal{A}^2 \\ & \left. + \frac{\varphi'^2}{2} \mathcal{A}^2 - \frac{2\mathcal{H}\Delta}{\kappa} \mathcal{A} \mathcal{B} + \frac{\mathcal{K}\Delta}{\kappa} \mathcal{B}^2 \right]. \end{aligned} \quad (\text{A9})$$

Here  $\mathcal{R}$ ,  $\mathcal{A}$ , and  $\mathcal{B}$  are the following gauge-invariant combinations:

$$\mathcal{R} = \Psi + \frac{\mathcal{H}}{\varphi'} \Phi, \quad (\text{A10})$$

$$\mathcal{A} = A + \frac{1}{\mathcal{H}} \Psi' + \left( \frac{\mathcal{H}'}{\mathcal{H}\varphi'} - \frac{\mathcal{H}}{\varphi'} \right) \Phi, \quad (\text{A11})$$



$$\mathcal{B} = B - E' \mp \frac{1}{\mathcal{H}}\Psi. \quad (\text{A12})$$

Also, in the above Lagrangian we wrote  $\Delta\mathcal{R}^2 = \mathcal{R}\Delta\mathcal{R}$ ,  $\Delta\mathcal{A}\mathcal{B} = \mathcal{A}\Delta\mathcal{B}$ , and so on for notational simplicity. We can integrate out  $\mathcal{A}$  and  $\mathcal{B}$  from the Lagrangian (A9). The minimization condition gives

$$\mathcal{A} = \frac{\pm(6\mathcal{K}^2\mathcal{H} + 2\mathcal{K}\mathcal{H}\Delta)\mathcal{R} - \kappa\mathcal{K}\Phi'^2\mathcal{R}'}{\mathcal{H}(2(3\mathcal{K} + \Delta)\mathcal{H}^2 - \kappa\mathcal{K}\Phi'^2)}, \quad (\text{A13})$$

$$\mathcal{B} = \frac{\pm\kappa\mathcal{K}\Phi'^2\mathcal{R} - \kappa\mathcal{H}\Phi'^2\mathcal{R}'}{\mathcal{H}(2(3\mathcal{K} + \Delta)\mathcal{H}^2 - \kappa\mathcal{K}\Phi'^2)}, \quad (\text{A14})$$

and the resulting Lagrangian becomes

$$\begin{aligned} \mathcal{L}^{(2)} = & a^2\sqrt{\gamma} \left[ \frac{(3\mathcal{K} + \Delta)\kappa\varphi'^2\mathcal{R}'^2}{2(3\mathcal{K} + \Delta)\mathcal{H}^2 - \kappa\mathcal{K}\varphi'^2} \right. \\ & + \frac{(3\mathcal{K} + \Delta)\mathcal{R}^2}{(2(3\mathcal{K} + \Delta)\mathcal{H}^2 - \kappa\mathcal{K}\varphi'^2)^2} \\ & \times \left( (3\mathcal{K} + \Delta) \left( \mp 2\mathcal{H}^2\varphi'^2(2\mathcal{K} - \Delta) - 4\mathcal{K}^2\varphi'^2 \right. \right. \\ & \left. \left. - 4\mathcal{K}\mathcal{H}\varphi'a^2\frac{\partial V}{\partial\varphi} \right) \pm \kappa\mathcal{K}\varphi'^4(2\mathcal{K} + \Delta) \right) \Big]. \quad (\text{A15}) \end{aligned}$$

The same action can be obtained if we directly integrate out  $A$  and  $B - E'$  from the action (A5). This action is reduced to the single-field inflationary quadratic action in the zero-curvature limit  $\mathcal{K} \rightarrow 0$  and  $\gamma_{ij} \rightarrow \delta_{ij}$  (see, e.g., Ref. [33]):

$$\mathcal{L}^{(2)} \xrightarrow{\mathcal{K} \rightarrow 0} \frac{a^2\varphi'^2}{2\mathcal{H}^2} (\mathcal{R}'^2 \pm \mathcal{R}\Delta\mathcal{R}). \quad (\text{A16})$$

The other limit is the zero mode  $\Delta \rightarrow 0$ . We get

$$\begin{aligned} \mathcal{L}^{(2)} \xrightarrow{\Delta \rightarrow 0} & a^2\sqrt{\gamma} \left[ \frac{3\varphi'^2}{6\mathcal{H}^2 - \kappa\varphi'^2} \mathcal{R}'^2 + \frac{6\mathcal{K}}{(6\mathcal{H}^2 - \kappa\varphi'^2)^2} \right. \\ & \times \left( \mp 6\mathcal{H}^2\varphi'^2 - 6\mathcal{K}\varphi'^2 \pm \kappa\varphi'^4 - 6\mathcal{H}\varphi'a^2\frac{\partial V}{\partial\varphi} \right) \mathcal{R}^2 \Big]. \quad (\text{A17}) \end{aligned}$$

In Sec. II we discuss the negative mode condition based on this Lagrangian.

## 2. Comparison with L-I

The authors started with the Lagrangian (2.8). They took the gauge

$$\Psi = 0 \quad (\text{A18})$$

and eliminated  $A$  as an auxiliary field. This gauge condition corresponds to taking  $\epsilon = a\Psi/\varphi'$  in Eq. (2.9). The resulting action is<sup>2</sup>

$$\begin{aligned} S^{(2)} = & \int d\tau d^3x \frac{a^2\sqrt{\gamma}}{2\kappa} \\ & \times \left[ \kappa \left( \Phi'^2 + a^2 \frac{\partial^2 V}{\partial\varphi^2} \Phi^2 \right) - \left( 2\kappa\varphi'\Phi' - 2\kappa a^2 \frac{\partial V}{\partial\varphi} \Phi \right) A \right. \\ & \left. - 2(\mathcal{H}' + 2\mathcal{H}^2 + \mathcal{K})A^2 \right] \\ = & \int d\tau d^3x \frac{a^2\sqrt{\gamma}\mathcal{H}^2}{2Q} \left[ \Phi'^2 - \kappa a^2 \frac{\partial V}{\partial\varphi} \frac{\varphi'}{3\mathcal{H}^2} \Phi'\Phi \right. \\ & \left. + \left( \frac{\kappa}{6\mathcal{H}^2} \left( a^2 \frac{\partial V}{\partial\varphi} \right)^2 + a^2 \frac{Q}{\mathcal{H}^2} \frac{\partial^2 V}{\partial\varphi^2} \right) \Phi^2 \right]. \quad (\text{A19}) \end{aligned}$$

They read off the negative mode condition as

$$Q = \mathcal{H}^2 - \frac{\kappa\varphi'^2}{6}. \quad (\text{A20})$$

## 3. Comparison with GMST and L-II

We follow Appendix B of Ref. [17], focusing on the  $O(4)$ -invariant mode only. Approach II of Ref. [7] uses equations derived in Ref. [17]. Reference [17] started from the Lorentzian action, while we discuss the Euclidean action in Sec. III. Here we keep both conventions. We start with the Lagrangian (A5) with  $\Delta \rightarrow 0$  and define the canonical momenta as

$$\Pi_\Psi = \frac{2a^2\sqrt{\gamma}}{\kappa} \left( -3\Psi' + \frac{3\kappa}{2}\varphi'\Phi - 3\mathcal{H}A \right), \quad (\text{A21})$$

$$\Pi_\Phi = a^2\sqrt{\gamma}(\Phi' - \varphi'A). \quad (\text{A22})$$

The constrained Hamiltonian  $H_C$  is constructed from

$$\mathcal{L}^{(2)} = \Pi_\Psi\Psi' + \Pi_\Phi\Phi' - H_C, \quad (\text{A23})$$

and it becomes

$$\begin{aligned} H_C = & \frac{\kappa}{12a^2\sqrt{\gamma}\mathcal{K}} \left( -\mathcal{K}\Pi_\Psi^2 + \frac{6\mathcal{K}}{\kappa}\Pi_\Phi^2 \right) + \frac{\kappa\varphi'}{2}\Pi_\Psi\Phi + a^2\sqrt{\gamma} \\ & \times \left[ \pm \frac{3\mathcal{K}}{\kappa}\Psi^2 + \frac{1}{2} \left( \mp 3\mathcal{K} + \mathcal{H}^2 + \mathcal{H}' - \frac{\varphi'''}{\varphi'} \right) \Phi^2 \right] \\ & - AC_2. \quad (\text{A24}) \end{aligned}$$

<sup>2</sup>Note that the square for the term  $(\partial V/\partial\varphi)^2$  is missing in Eq. (33) of Ref. [7].

where the upper (lower) signs are for Lorentzian (Euclidean) metric, and  $C_2$  is given in Eq. (3.9),

$$\begin{aligned} C_2 &= -\varphi' \Pi_\Phi + \mathcal{H} \Pi_\Psi + a^2 \sqrt{\gamma} \\ &\times \left[ \pm \frac{6\mathcal{K}}{\kappa} \Psi + (\varphi'' - \mathcal{H}\varphi') \Phi \right] \\ &= -\varphi' \Pi_\Phi + \mathcal{H} \Pi_\Psi + a^2 \sqrt{\gamma} \\ &\times \left[ \pm \frac{6\mathcal{K}}{\kappa} \Psi + \left( -3\mathcal{H}\varphi' \mp a^2 \frac{\partial V}{\partial \varphi} \right) \Phi \right]. \end{aligned} \quad (\text{A25})$$

By using  $C_2$ ,  $\mathcal{L}^{(2)}$  can be rewritten as

$$\mathcal{L}^{(2)}[\Pi_\Phi, \Pi_\Psi, \Psi, \Phi] = \mathcal{L}^{(2)*}[\Pi_\Psi, \Psi, \Phi] + \frac{C_2^2}{2a^2 \sqrt{\gamma} \varphi'^2}. \quad (\text{A26})$$

Since the rhs does not contain any linear terms in  $C_2$ , the equations of motion from  $\mathcal{L}^{(2)*}$  are consistent with those from  $\mathcal{L}^{(2)}$  as long as  $C_2 = 0$  is satisfied. Thus, we can eliminate  $\Pi_\Phi$  by using  $C_2 = 0$  and use  $\mathcal{L}^{(2)*}$  as the Lagrangian. The Lagrangian  $\mathcal{L}^{(2)*}$  can be simplified by adopting gauge-invariant variables,

$$\Psi \equiv \Psi + \frac{\mathcal{H}}{\varphi'} \Phi, \quad (\text{A27})$$

$$\Pi_\Psi \equiv \Pi_\Psi \mp \frac{6a^2 \sqrt{\gamma} \mathcal{K}}{\kappa \varphi'} \Phi. \quad (\text{A28})$$

In Sec. III B we fix the gauge to be  $\Phi = 0$  rather than use these gauge-invariant variables  $\Psi$  and  $\Pi_\Psi$ . Here we proceed with  $\Psi$  and  $\Pi_\Psi$ . By using  $\Psi$  and  $\Pi_\Psi$ , the reduced Lagrangian (denoted by the superscript  $*$ ) becomes

$$\mathcal{L}^{(2)*} = \Pi_\Psi \Psi' - H_C^*, \quad (\text{A29})$$

where the constrained Hamiltonian is

$$\begin{aligned} H_C^* &= \frac{2a^2 \sqrt{\gamma}}{\varphi'^2} \left( \pm \frac{3\mathcal{K}}{\kappa} \Psi + \frac{\mathcal{H}}{2a^2 \sqrt{\gamma}} \Pi_\Psi \right)^2 \\ &\pm \frac{3a^2 \sqrt{\gamma}}{\kappa} \mathcal{K} \Psi^2 - \frac{\kappa}{12a^2 \sqrt{\gamma}} \Pi_\Psi^2 \\ &= a^2 \sqrt{\gamma} \left[ \frac{6\mathcal{H}^2 - \kappa \varphi'^2}{12\varphi'^2} \left( \frac{\Pi_\Psi}{a^2 \sqrt{\gamma}} \right)^2 \right. \\ &\left. \pm \frac{6\mathcal{K}\mathcal{H}}{\kappa \varphi'^2} \frac{\Pi_\Psi}{a^2 \sqrt{\gamma}} \Psi + \frac{3\mathcal{K}}{\kappa} \frac{6\mathcal{K} \pm \kappa \varphi'^2}{\kappa \varphi'^2} \Psi^2 \right]. \end{aligned} \quad (\text{A30})$$

We next perform a canonical transformation using a (type-three) generating function

$$W = \mp \frac{3a^2 \sqrt{\gamma} \mathcal{K}}{\mathcal{H}} \left( \frac{1}{\kappa} \Psi^2 - \varphi' \Psi \tilde{\mathbf{q}} + \frac{\kappa \varphi'^2}{8} \tilde{\mathbf{q}}^2 \right). \quad (\text{A31})$$

The relation between the original variables  $(\Psi, \Pi_\Psi)$  and new ones  $(\tilde{\mathbf{q}}, \tilde{\mathbf{p}})$  is obtained from  $\Pi_\Psi = \partial W / \partial \Psi$  and  $\tilde{\mathbf{p}} = -\partial W / \partial \tilde{\mathbf{q}}$ , which gives<sup>3</sup>

$$\Psi = \frac{\kappa \varphi'}{4} \tilde{\mathbf{q}} \mp \frac{\mathcal{H}}{3a^2 \sqrt{\gamma} \mathcal{K} \varphi'} \tilde{\mathbf{p}}, \quad (\text{A32})$$

$$\Pi_\Psi = \pm \frac{3a^2 \sqrt{\gamma} \mathcal{K} \varphi'}{2\mathcal{H}} \tilde{\mathbf{q}} + \frac{2}{\kappa \varphi'} \tilde{\mathbf{p}}. \quad (\text{A33})$$

The new Hamiltonian  $K_C^*$  becomes

$$\begin{aligned} K_C^* &= H_C^* + \frac{\partial W}{\partial \tau} \\ &= \mp \frac{1}{6a^2 \sqrt{\gamma} \mathcal{K}} \tilde{\mathbf{p}}^2 + \left( \mathcal{H} - \frac{\kappa \varphi'^2}{4\mathcal{H}} + \frac{a^2 dV}{\varphi' d\varphi} \right) \tilde{\mathbf{p}} \tilde{\mathbf{q}} \\ &\quad + a^2 \sqrt{\gamma} \mathcal{K} \left[ \mp \frac{9\mathcal{H}^2}{2} \pm \frac{3\kappa \varphi'^2}{2} \pm \frac{\kappa^2 \varphi'^4}{32\mathcal{H}^2} \right. \\ &\quad \left. + \kappa a^2 \left( \frac{3}{2} - \frac{\kappa \varphi'^2}{8\mathcal{H}^2} \right) V \right] \tilde{\mathbf{q}}^2. \end{aligned} \quad (\text{A34})$$

After a Legendre transformation, we get a new Lagrangian,

$$\mathcal{L}^{(2),\text{new}} = \mp \frac{3a^2 \sqrt{\gamma} \mathcal{K}}{2} (\tilde{\mathbf{q}}'^2 \mp a^2 m^2 \tilde{\mathbf{q}}^2), \quad (\text{A35})$$

with  $m^2$  being

$$\begin{aligned} m^2 &= -\frac{1}{a^2} \left[ 4\mathcal{K} \mp 2\mathcal{H}' \pm \varphi' \left( \frac{1}{\varphi'} \right)' \right] \\ &= \mp \frac{2\kappa \varphi'^2}{3a^2} - \left( \frac{4\kappa}{3} V + \frac{8\mathcal{H}}{\varphi'} \frac{dV}{d\varphi} + \frac{d^2 V}{d\varphi^2} \right) \mp \frac{2a^2}{\varphi'^2} \left( \frac{dV}{d\varphi} \right)^2. \end{aligned} \quad (\text{A36})$$

In Sec. III B we discuss the coefficient  $3a^2 \sqrt{\gamma} \mathcal{K} / 2$  (starting from the Euclidean action) using the Dirac bracket. Equation (A35) is consistent with Eq. (23) in Ref. [34] and Eq. (B18) in Ref. [17]. See also Eq. (B19) in Ref. [17] and its errata [29].

Equation (10) in Ref. [30] is consistent with the Lorentzian action given in Eq. (A35) for a closed universe,  $\mathcal{K} = 1$ . On the other hand, the Euclidean actions given in Eq. (12) in Ref. [30] and Eq. (38) in Ref. [7] have the opposite sign. Reference [30] derived the Euclidean action from the Lorentzian action by using analytic continuation. As shown in Eqs. (A32) and (A33), the relation between  $\tilde{\mathbf{q}}$

<sup>3</sup>As discussed in Eq. (B20) of Ref. [17], the variable  $\tilde{\mathbf{q}}$  is equivalent to the Bardeen potential after substituting the minimization condition of the auxiliary fields.

and  $\Psi$  includes  $\varphi'$ , and the sign of the kinetic term of  $\tilde{\mathbf{q}}$  is flipped if we replace  $\varphi'$  to  $i\varphi'$ . Reference [30] did not adopt this replacement and the signs of the kinetic terms of  $\tilde{\mathbf{q}}$  are the same for both the Lorentzian and Euclidean actions.

#### 4. Comparison with L-III and KLT

Again, let us start from the Hamiltonian  $H_T$  given in Eq. (3.5). In Ref. [8], the authors set two gauge-fixing conditions,

$$A - \Psi = 0, \quad \Pi_\Psi = 0, \quad (\text{A37})$$

and eliminated  $A$  and  $\Pi_\Psi$  from the Hamiltonian. The reduced Hamiltonian is obtained as

$$\begin{aligned} H_{\text{KLT}} = & \frac{1}{2a^2\sqrt{\gamma}} \left( 1 - \frac{\kappa\varphi'^2}{6\mathcal{K}} \right) \Pi_\Phi^2 + \frac{\kappa\varphi'}{6\mathcal{K}} \left( a^2 \frac{\partial V}{\partial \varphi} - 3\mathcal{H}\varphi' \right) \Pi_\Phi \Phi \\ & - \frac{1}{2} a^2 \sqrt{\gamma} \left[ \frac{\kappa}{6\mathcal{K}} \left( a^2 \frac{\partial V}{\partial \varphi} - 3\varphi' \mathcal{H} \right)^2 \right. \\ & \left. + a^2 \frac{\partial^2 V}{\partial \varphi^2} + \frac{3}{2} \kappa \varphi'^2 \right] \Phi^2. \end{aligned} \quad (\text{A38})$$

Reference [7] derived the same effective Hamiltonian in approach III, even though their gauge-fixing condition ( $A = 0$ ,  $\Pi_\Psi = 0$ ) is inconsistent with the time evolution. Reference [7] eliminated  $\Psi$ ,  $\Pi_\Psi$ ,  $A$ , and  $\Pi_A$  by using constraints and gauge-fixing conditions. We can justify this elimination of variables, as discussed in Appendix B. The discussion of L-III and KLT is the same as the discussion utilizing the Dirac bracket in the main text, and we obtain consistent results. For details, see Sec. III B.

#### 5. Comparison with GT

Reference [9] started from the action given in Eq. (18) in Ref. [34]. In this section we follow the notation of Refs. [9,34], keeping signs for both Lorentzian and Euclidean setups. We rewrite Eq. (A5) for  $O(4)$ -symmetric fluctuations in the notation of Refs. [9,34] as

$$\begin{aligned} S^{(2)} = & \int d\tau d^3x \frac{a^2\sqrt{\gamma}}{2\kappa} [-6\psi'^2 - 12\mathcal{H}A\psi' - 2(\mathcal{H}' + 2\mathcal{H}^2)A^2 \\ & + \kappa(\delta\phi'^2 \mp a^2 V_{,\phi\phi} \delta\phi^2) \\ & + 2\kappa(3\phi'_0\psi'\delta\phi - \phi'_0\delta\phi'A \mp a^2 V_{,\phi} A\delta\phi) \\ & \pm \mathcal{K}(-6\psi'^2 + 2A^2 + 12\psi A)], \end{aligned} \quad (\text{A39})$$

where the upper (lower) signs are for Lorentzian (Euclidean) metric. For  $O(4)$ -nonsymmetric fluctuations, see Eq. (18) in Ref. [34]. Conjugate momenta for  $\psi$  and  $\delta\phi$  are introduced as

$$\begin{aligned} \Pi_\psi & \equiv \frac{2a^2\sqrt{\gamma}}{\kappa} \left( -3\psi' + \frac{3}{2}\kappa\phi'_0\delta\phi - 3\mathcal{H}A \right), \\ \Pi_{\delta\phi} & \equiv a^2\sqrt{\gamma}(\delta\phi' - \phi'_0 A), \end{aligned} \quad (\text{A40})$$

and the action can be rewritten as

$$\begin{aligned} S^{(2)} = & \int d\tau d^3x \left[ \Pi_\psi \psi' + \Pi_{\delta\phi} \delta\phi' - \frac{\kappa}{12a^2\sqrt{\gamma}\mathcal{K}} \right. \\ & \times \left( -\kappa\Pi_\psi^2 + \frac{6\mathcal{K}}{\kappa}\Pi_{\delta\phi}^2 \right) - \frac{\kappa\phi'_0}{2}\Pi_\psi\delta\phi \\ & - a^2\sqrt{\gamma} \left\{ \pm \frac{3\mathcal{K}}{\kappa}\psi'^2 + \frac{1}{2} \left( \mp 3\mathcal{K} + \mathcal{H}^2 \right. \right. \\ & \left. \left. + \mathcal{H}' - \frac{\phi_0'''}{\phi_0'} \right) \delta\phi^2 \right\} + AC_2 \Big]. \end{aligned} \quad (\text{A41})$$

Here  $C_2$  is the same constraint as in Appendix A 3. In the notation of Refs. [9,34],

$$C_2 \equiv -\phi'_0\Pi_{\delta\phi} + \mathcal{H}\Pi_\psi + a^2\sqrt{\gamma} \left[ \pm \frac{6\mathcal{K}}{\kappa}\psi' + (\phi_0'' - \mathcal{H}\phi_0')\delta\phi \right]. \quad (\text{A42})$$

By using  $C_2$ , the action can be rewritten as

$$\begin{aligned} S^{(2)} = & \int d\tau d^3x \frac{a^2\sqrt{\gamma}}{3\mathcal{K}} \left[ \pm \frac{2}{\kappa\phi_0'} \Psi_l \delta\phi'_l \pm \frac{2(\mathcal{H}\phi_0' - \phi_0'')}{\kappa\phi_0'^2} \Psi_l \delta\phi_l \right. \\ & \left. \pm \frac{1}{2} \delta\phi_l^2 \mp \frac{1}{\kappa} \left( 1 \pm \frac{6\mathcal{K}}{\kappa\phi_0'^2} \right) \Psi_l^2 \right] + \frac{C_2^2}{2a^2\sqrt{\gamma}\phi_0'^2}. \end{aligned} \quad (\text{A43})$$

Here  $\Psi_l$  and  $\delta\phi_l$  are gauge-invariant variables defined as

$$\Psi_l = 3\mathcal{K}\psi \pm \frac{\kappa\mathcal{H}}{2a^2\sqrt{\gamma}}\Pi_\psi, \quad (\text{A44})$$

$$\delta\phi_l = 3\mathcal{K}\delta\phi \mp \frac{\kappa\phi'}{2a^2\sqrt{\gamma}}\Pi_\psi. \quad (\text{A45})$$

Since the rhs of Eq. (A43) does not contain linear terms in  $C_2$ , the equations of motion are not affected by dropping the  $C_2^2$  term from the Lagrangian as long as  $C_2 = 0$  is satisfied. In Eq. (A43),  $\Pi_{\delta\phi}$  is contained only in the  $C_2^2$  term. Thus, dropping the  $C_2^2$  term in Eq. (A43) is equivalent to eliminating  $\Pi_{\delta\phi}$  by using  $C_2 = 0$ :

$$\begin{aligned} S^{(2)} = & \int d\tau d^3x \frac{a^2\sqrt{\gamma}}{3\mathcal{K}} \left[ \pm \frac{2}{\kappa\phi_0'} \Psi_l \delta\phi'_l \pm \frac{2(\mathcal{H}\phi_0' - \phi_0'')}{\kappa\phi_0'^2} \Psi_l \delta\phi_l \right. \\ & \left. \pm \frac{1}{2} \delta\phi_l^2 \mp \frac{1}{\kappa} \left( 1 \pm \frac{6\mathcal{K}}{\kappa\phi_0'^2} \right) \Psi_l^2 \right]. \end{aligned} \quad (\text{A46})$$

After integrating out  $\Psi_l$  we obtain a quadratic action for  $\delta\phi_l$ . The coefficient of  $\Psi_l^2$  essentially determines the negative

mode condition, and we discuss it using Dirac bracket in Sec. III B. In the Lorentzian setup (as discussed in Refs. [9,34]), the CDL bounce corresponds to the case with  $\mathcal{K} = 1$  (with an appropriate unit) and we should flip the sign of  $\phi_0'^2$ , and the negative modes appear if  $1 - \kappa\phi_0'^2/6$  becomes negative.

## APPENDIX B: COMMENTS ON ELIMINATION OF VARIABLES BY CONSTRAINTS

In Refs. [7,8] the Hamiltonian was reduced by eliminating variables using constraints. For details, see Appendix A 4. In this Appendix we discuss the validity of this procedure.

### 1. Example in which elimination does NOT work

First, we show that eliminating variables using constraints can give the wrong equations of motion in general. Let us take a Hamiltonian  $H(q, p)$  with a constraint  $C(q, p)$ .  $C(q, p) = 0$  can be solved for  $p_n$ , and let us denote the solution as

$$p_n = f(q_1, \dots, q_n, p_1, \dots, p_{n-1}). \quad (\text{B1})$$

Then, we define the reduced Hamiltonian  $\tilde{H}$  as

$$\begin{aligned} \tilde{H}(q_1, \dots, q_n, p_1, \dots, p_{n-1}) \\ \equiv H(q_1, \dots, q_n, p_1, \dots, p_{n-1}, f(q_1, \dots, p_{n-1})). \end{aligned} \quad (\text{B2})$$

The derivatives of  $\tilde{H}$  are given as

$$\frac{\partial \tilde{H}}{\partial q_i} = \frac{\partial H}{\partial q_i} + \frac{\partial H}{\partial p_n} \frac{\partial f}{\partial q_i}, \quad \frac{\partial \tilde{H}}{\partial p_i} = \frac{\partial H}{\partial p_i} + \frac{\partial H}{\partial p_n} \frac{\partial f}{\partial p_i}. \quad (\text{B3})$$

Thus, in general,  $\tilde{H}$  does not give a Hamiltonian equation that is consistent with the original Hamiltonian.

Let us see this explicitly in the following simple example:

$$L = -\frac{3a\dot{a}^2}{N} + \frac{1}{2N}a^3\dot{\phi}^2 - a^3V(\phi). \quad (\text{B4})$$

This is the Lagrangian for the inflaton zero mode. The conjugate momenta for each variable are

$$\Pi_a = -\frac{6a\dot{a}}{N}, \quad \Pi_\phi = \frac{a^3\dot{\phi}}{N}, \quad \Pi_N = 0. \quad (\text{B5})$$

Thus, the Hamiltonian is

$$H = \Pi_N \dot{N} + N \left( -\frac{p_a^2}{12a} + \frac{1}{2a^3} p_\phi^2 + a^3 V(\phi) \right). \quad (\text{B6})$$

In this system,  $\Pi_N = 0$  is a primary constraint and we obtain a secondary constraint from  $\dot{\Pi}_N = \partial H / \partial N = 0$ :

$$-\frac{p_a^2}{12a} + \frac{1}{2a^3} p_\phi^2 + a^3 V = 0. \quad (\text{B7})$$

If we simplify the Hamiltonian by using this constraint, the new reduced Hamiltonian is  $\tilde{H} = \Pi_N \dot{N}$ .  $\tilde{H}$  does not depend on  $\phi$  and  $a$ , and apparently  $\tilde{H}$  does not provide the correct Hamilton equations.

### 2. Example in which elimination works

Next, we show an example in which elimination works. Let us discuss a system that has the following two constraints:

$$\phi_1 = \sum_{i=1}^n (a_i q_i + b_i p_i), \quad \phi_2 = c q_n + d p_n. \quad (\text{B8})$$

Here we assume that  $a_n d - b_n c$  is nonzero. By solving  $\phi_1 = \phi_2 = 0$  for  $q_n$  and  $p_n$ , we obtain

$$\begin{aligned} q_n &= -\frac{d}{a_n d - b_n c} \sum_{i=1}^{n-1} (a_i q_i + b_i p_i), \\ p_n &= \frac{c}{a_n d - b_n c} \sum_{i=1}^{n-1} (a_i q_i + b_i p_i). \end{aligned} \quad (\text{B9})$$

The Dirac bracket is defined as

$$[A, B]_D \equiv [A, B] + \frac{1}{[\phi_1, \phi_2]} ([A, \phi_1][\phi_2, B] - [A, \phi_2][\phi_1, B]). \quad (\text{B10})$$

Let us assume that  $A$  does not depend on  $q_n$  and  $p_n$ . In this case, the Dirac bracket is

$$\begin{aligned} [A, H]_D &= [A, H] + \frac{1}{[\phi_1, \phi_2]} [A, \phi_1][\phi_2, H] \\ &= \sum_{i=1}^{n-1} \left( \frac{\partial A}{\partial q_i} \left[ \frac{\partial H}{\partial p_i} + \frac{b_i}{a_n d - b_n c} \left( -\frac{\partial H}{\partial q_n} d + \frac{\partial H}{\partial p_n} c \right) \right] \right. \\ &\quad \left. - \sum_{i=1}^{n-1} \left( \frac{\partial A}{\partial p_i} \left[ \frac{\partial H}{\partial q_i} + \frac{a_i}{a_n d - b_n c} \left( -\frac{\partial H}{\partial q_n} d + \frac{\partial H}{\partial p_n} c \right) \right] \right) \right. \\ &\quad \left. - \sum_{i=1}^{n-1} \left( \frac{\partial A}{\partial q_i} \left[ \frac{\partial H}{\partial p_i} + \frac{\partial H}{\partial q_n} \frac{\partial q_n}{\partial p_i} + \frac{\partial H}{\partial p_n} \frac{\partial p_n}{\partial p_i} \right] \right. \right. \\ &\quad \left. \left. - \frac{\partial A}{\partial p_i} \left[ \frac{\partial H}{\partial q_i} + \frac{\partial H}{\partial q_n} \frac{\partial q_n}{\partial q_i} + \frac{\partial H}{\partial p_n} \frac{\partial p_n}{\partial q_i} \right] \right) \right). \end{aligned} \quad (\text{B11})$$

Here we have used Eq. (B9). Thus, the following relation is satisfied:

$$[A, H]_D = [A, \tilde{H}]. \quad (\text{B12})$$

where  $\tilde{H}$  is the reduced Hamiltonian,

$$\begin{aligned}
& \tilde{H}(q_1, \dots, q_{n-1}, p_1, \dots, p_{n-1}) \\
& \equiv H(q_1, \dots, q_{n-1}, q_n(q_1, \dots, p_{n-1}), \\
& \quad \times p_1, \dots, p_{n-1}, p_n(q_1, \dots, p_{n-1})). \quad (\text{B13})
\end{aligned}$$

Therefore, we can reduce the Hamiltonian for  $q_n$  and  $p_n$  at least when the constraints have the same form as in Eq. (B8). In order to derive the Hamiltonian in Eq. (17)

of Ref. [8], we have to utilize four constraints:  $\Pi_\Psi = 0$ ,  $A - \Psi = 0$ ,  $\Pi_A = 0$ , and  $\varphi'\Pi_\Phi - \mathcal{H}\Pi_\Psi + a^2\sqrt{\gamma}[(a^2(\partial V/\partial\varphi) + 3\varphi'\mathcal{H})\Phi - (6\mathcal{K}/\kappa)\Psi] = 0$ . From the above discussions, we can safely eliminate  $\Psi$ ,  $\Pi_\Psi$ , and  $A$  (and  $\Pi_A$ ) from the Hamiltonian by using these four constraints.

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- [1] S. R. Coleman, The fate of the false vacuum. 1. Semi-classical theory, *Phys. Rev. D* **15**, 2929 (1977).
  - [2] C. G. Callan, Jr. and S. R. Coleman, The fate of the false vacuum. 2. First quantum corrections, *Phys. Rev. D* **16**, 1762 (1977).
  - [3] S. R. Coleman, Quantum tunneling and negative eigenvalues, *Nucl. Phys.* **B298**, 178 (1988).
  - [4] S. R. Coleman and F. De Luccia, Gravitational effects on and of vacuum decay, *Phys. Rev. D* **21**, 3305 (1980).
  - [5] G. V. Lavrelashvili, V. A. Rubakov, and P. G. Tinyakov, Tunneling transitions with gravitation: Breaking of the Quasiclassical approximation, *Phys. Lett.* **161B**, 280 (1985).
  - [6] H. Lee and E. J. Weinberg, Negative modes of Coleman-De Luccia bounces, *Phys. Rev. D* **90**, 124002 (2014).
  - [7] G. V. Lavrelashvili, Negative mode problem in false vacuum decay with gravity, *Nucl. Phys. B, Proc. Suppl.* **88**, 75 (2000).
  - [8] A. Khvedelidze, G. V. Lavrelashvili, and T. Tanaka, On cosmological perturbations in closed FRW model with scalar field and false vacuum decay, *Phys. Rev. D* **62**, 083501 (2000).
  - [9] S. Gratton and N. Turok, Homogeneous modes of cosmological instantons, *Phys. Rev. D* **63**, 123514 (2001).
  - [10] T. Tanaka and M. Sasaki, False vacuum decay with gravity: Negative mode problem, *Prog. Theor. Phys.* **88**, 503 (1992).
  - [11] T. Tanaka, The no—negative mode theorem in false vacuum decay with gravity, *Nucl. Phys.* **B556**, 373 (1999).
  - [12] G. V. Dunne and Q.-h. Wang, Fluctuations about cosmological instantons, *Phys. Rev. D* **74**, 024018 (2006).
  - [13] M. Koehn, G. Lavrelashvili, and J.-L. Lehners, Towards a solution of the negative mode problem in quantum tunneling with gravity, *Phys. Rev. D* **92**, 023506 (2015).
  - [14] S. F. Bramberger, M. Chitishvili, and G. Lavrelashvili, Aspects of the negative mode problem in quantum tunneling with gravity, *Phys. Rev. D* **100**, 125006 (2019).
  - [15] A. Hebecker, T. Mikhail, and P. Soler, Euclidean wormholes, baby universes, and their impact on particle physics and cosmology, *Front. Astron. Space Sci.* **5**, 35 (2018).
  - [16] T. Hertog, B. Truijen, and T. Van Riet, Euclidean Axion Wormholes Have Multiple Negative Modes, *Phys. Rev. Lett.* **123**, 081302 (2019).
  - [17] J. Garriga, X. Montes, M. Sasaki, and T. Tanaka, Canonical quantization of cosmological perturbations in the one-bubble open Universe, *Nucl. Phys.* **B513**, 343 (1998).
  - [18] S. R. Coleman, V. Glaser, and A. Martin, Action minima among solutions to a class of euclidean scalar field equations, *Commun. Math. Phys.* **58**, 211 (1978).
  - [19] O. Lopes, Radial symmetry of minimizers for some translation and rotation invariant functionals, *J. Differ. Equ.* **124**, 378 (1996).
  - [20] J. Byeon, L. Jeanjean, and M. Mariş, Symmetry and monotonicity of least energy solutions, *Calc. Var. Partial Differ. Equ.* **36**, 481 (2009).
  - [21] K. Blum, M. Honda, R. Sato, M. Takimoto, and K. Tobioka,  $O(N)$  Invariance of the multi-field bounce, *J. High Energy Phys.* **05** (2017) 109; Erratum, **06** (2017) 060.
  - [22] P. A. M. Dirac, *Lectures on Quantum Mechanics* (Dover Publications, New York, 1964).
  - [23] R. Cawley, Determination of the Hamiltonian in the Presence of Constraints, *Phys. Rev. Lett.* **42**, 413 (1979).
  - [24] L. Castellani, Symmetries in constrained hamiltonian systems, *Ann. Phys. (N.Y.)* **143**, 357 (1982).
  - [25] R. Sugano, Poincare-Cartan invariant form and dynamical systems with constraints. II., *Prog. Theor. Phys.* **68**, 1377 (1982).
  - [26] T. Maskawa and H. Nakajima, Singular lagrangian and Dirac-Faddeev method: Existence theorems of constraints in standard forms, *Prog. Theor. Phys.* **56**, 1295 (1976).
  - [27] R. Sugano and T. Kimura, On the relation of first class constraints to gauge degrees of freedom, *Prog. Theor. Phys.* **69**, 252 (1983).
  - [28] M. de León, J. C. Marrero, and D. M. de Diego, Time-dependent constrained hamiltonian systems and dirac brackets, *J. Phys. A* **29**, 6843 (1996).
  - [29] J. Garriga, X. Montes, M. Sasaki, and T. Tanaka, Canonical quantization of cosmological perturbations in the one-bubble open universe, *Nucl. Phys.* **B513**, 343 (1998); **551B**, 511(E) (1999).
  - [30] G. V. Lavrelashvili, On the quadratic action of the Hawking-Turok instanton, *Phys. Rev. D* **58**, 063505 (1998).
  - [31] G. W. Gibbons, S. W. Hawking, and M. J. Perry, Path integrals and the indefiniteness of the gravitational action, *Nucl. Phys.* **B138**, 141 (1978).
  - [32] S. P. De Alwis, F. Muia, V. Pasquarella, and F. Quevedo, Quantum transitions between minkowski and de sitter spacetimes, *Fortschr. Phys.* **68**, 2000069 (2020).
  - [33] J. M. Maldacena, Non-Gaussian features of primordial fluctuations in single field inflationary models, *J. High Energy Phys.* **05** (2003) 013.
  - [34] S. Gratton and N. Turok, Cosmological perturbations from the no boundary Euclidean path integral, *Phys. Rev. D* **60**, 123507 (1999).