# Global perspectives on dark matter simplified models

### Emanuele A. Bagnaschi (DESY Hamburg)



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Based on Bagnaschi E. et al [1809.xxxx]

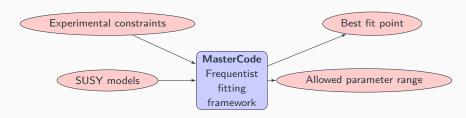


# Introduction

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#### Global likelihood analyses

- Global-fit frameworks such as MasterCode, have been used for years to study BSM extensions of the SM. In our case, up to now we focused our attention to SUSY models (see the other MC talks on SUSY-DM and SUSY @ LHC later today).
- Dark Matter Simplified Models are not complete models, not meaningful to compute p-values. Use the framework with the aim of correlating different sectors and finding still allowed regions.



# The models

# Dark Matter Simplified Models

# The Lagrangians

- We consider DMSMs with a spin-1 (Y<sub>1</sub>) s-channel mediator.
- The dark matter candidate is a Dirac fermion  $(X_D)$ .
- We use the model files provided by the DMSIMP package for our implementation.

### Spin-1 mediator

- Interaction Lagrangian mediator-DM  $\mathcal{L}_{X_D}^{Y_1} = \bar{X}_D \gamma_{\mu} \left( g_{X_D}^V + g_{X_D}^A \gamma_5 \right) X_D Y_1^{\mu}.$
- Interaction Lagrangian mediator-quarks

$$\begin{split} \mathcal{L}_{quarks}^{Y_{1}} &= \sum_{i,j} \left[ \bar{d}_{i} \gamma_{\mu} \left( g_{d_{i,j}}^{V} + g_{d_{i,j}}^{A} \gamma_{5} \right) d_{j} \right. \\ &+ \left. \bar{u}_{i} \gamma_{\mu} \left( g_{u_{i,j}}^{V} + g_{u_{i,j}}^{A} \gamma_{5} \right) u_{j} \right] Y_{1}^{\mu} \end{split}$$

Interaction Lagrangian mediator-leptons  $\mathcal{L}_{leptons}^{Y_1} =$ 

$$\sum_{i,j} \left[ \overline{l}_i \gamma_{\mu} \left( g_{l_{i,j}}^V + g_{l_{i,j}}^A \gamma_5 \right) l_j \right] Y_1^{\mu}$$

#### **Scenarios**

- Leptophobic,  $g_{l_{i,j}}^V = g_{l_{i,j}}^A = 0$  (no constraints from dilepton searches).
- Flavor diagonal,  $g_{u/d_{i,j}}^{V/A} = 0$  if  $i \neq j$ .
- Flavor blind,  $g_{u_{i,j}}^{V/A} = g_{d_{i,j}}^{V/A}$ .
- 1.  $g_{X_D}^V \equiv g_{DM}$   $g_{X_D}^V = 0$   $g_{u/d}^U \equiv g_{SM}$   $g_{u/d}^A = 0$ , pure vector.
- $\begin{aligned} & 2. & \ g_{X_D}^V = 0 & \ \ g_{X_D}^V \equiv g_{DM} \\ & \ \ g_{u/d}^V = 0 & \ \ g_{u/d}^A = g_{SM}, \end{aligned}$  pure axial-vector.

# The framework

#### The framework

- Frequentist fitting framework written in Python/Cython and C++.
- The Multinest algorithm is used to sample the parameter space.
- udocker used for deployment.

Parameter	Range	Number of
		segments
$M_{Y_1}$	(-4 , 4 ) TeV	6
$M_{DM}$	(0,1.2) TeV	2
g <sub>SM</sub>	$(0, \sqrt{4\pi})$	1
gDМ	$(1, \sqrt{4\pi})$	1
Total number of boxes		144

#### Codes

#### Mediator properties

Analytic/MadGraph\_aMC@NLO

#### Collider

MadGraph\_aMC@NLO (DMSIMP models)

# Dark matter

MicrOMEGAs

Sampled a total of xx points.

#### Docker and udocker

- As suggested by the name, udocker uses docker containers.
- Docker: software framework to automatize the deployment of application inside Linux containers.
- Other options, such as LXC, are available to use the Linux container infrastructure.







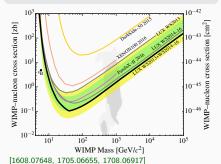
 Middleware suite developed in the context of the INDIGO data-cloud project to run docker containers in userspace, without requiring root privileges (both for installation and execution).

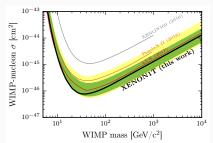
# The constraints

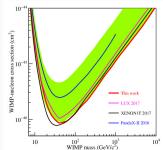
#### Non-LHC constraints

#### Dark matter

- Relic density constraints from Planck.
- Direct detection constraints on σ<sup>SI</sup><sub>p</sub> from LUX, XENON1T and PANDAX.
- Direct detection constraints on  $\sigma_p^{SD}$  from PIC060.







# **Validation**

### **Collider constraints**

# Monojet

# **Validation**

# Dijet

# **Validation**

# **Results**

### Pure vector

# Backup slides