

# Naturalness and light higgsinos: why ILC is the right machine for SUSY discovery

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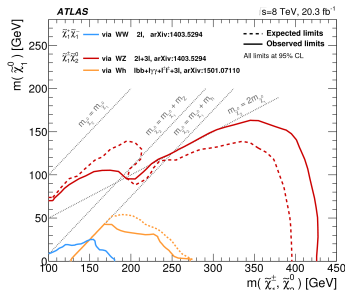
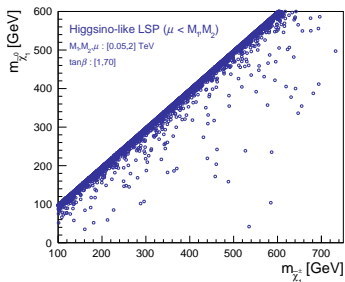
DER FORSCHUNG | DER LEHRE | DER BILDUNG

# Why study light higgsinos

- Naturalness and small fine tuning requires  $\mu$  parameter at the EW scale:

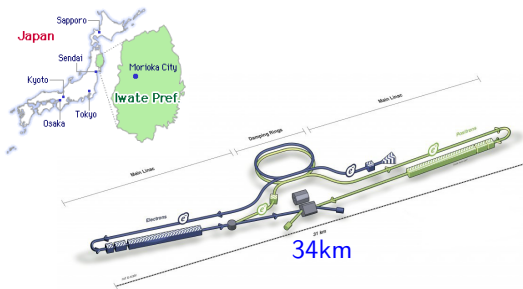
$$m_Z^2 = 2 \frac{m_{H_d}^2 - m_{H_u}^2 \tan^2 \beta}{\tan^2 \beta - 1} - 2\mu^2$$

- $\mu$  small  $\implies$  light higgsinos
- Typical mass difference 10 - 20 GeV  
 $\implies$  challenging for LHC if other particles are heavy



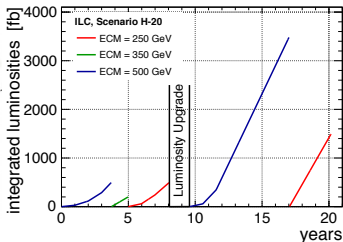
# International Linear Collider (ILC)

- Electron-positron collider at  $\sqrt{s} = 250 - 500\text{GeV}$  (1TeV)
- Polarisation of electrons 80%, positrons 30%
- Well-defined initial state: 4-momentum and spin config.
- Clean and completely reconstructable final state
- Under political consideration in Japan



## Typical 20yr running scenario

Integrated Luminosities [fb] arXiv:1506.07830



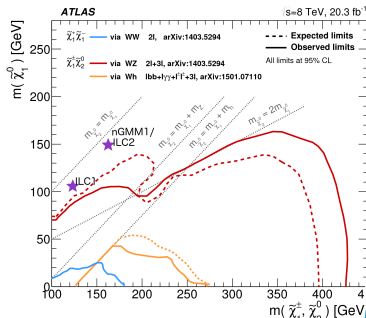
# Benchmarks studied

- $\tilde{\chi}_1^0, \tilde{\chi}_2^0, \tilde{\chi}_1^\pm$  observable,  $\tilde{\chi}_3^0$  accessible with a small cross section
- Other sparticles heavy
- Mass gaps  $\sim 10 - 20$  GeV  $\Rightarrow$  higgsinos decay via a virtual Z/W

## Three specific benchmarks

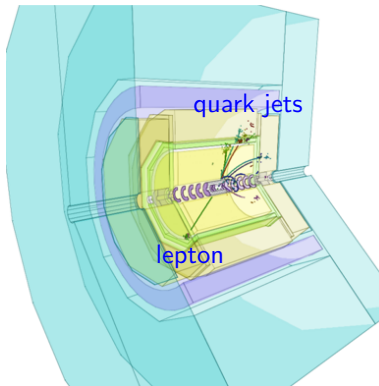
mass	ILC1	ILC2	nGMM1
$\tilde{\chi}_1^\pm - \tilde{\chi}_1^0$	14 GeV	10 GeV	7 GeV
$\tilde{\chi}_2^0 - \tilde{\chi}_1^0$	21 GeV	10 GeV	4 GeV
$\tilde{g}$	1560 GeV	2830 GeV	2860 GeV

Cross sections for production in  $e^+e^-$  at  $\sqrt{s} = 500$  GeV several hundred fb



# Detailed simulation: 500 GeV, 500 fb<sup>-1</sup>, $\mathcal{P}(\pm 80\%, \mp 30\%)$

$e^+e^- \rightarrow \tilde{\chi}_1^+ \tilde{\chi}_1^- \rightarrow \tilde{\chi}_1^0 q \bar{q} / \tilde{\chi}_1^0 e \nu_e$   
in the International Large Detector

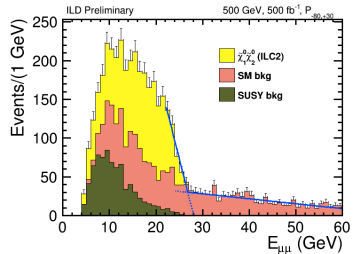
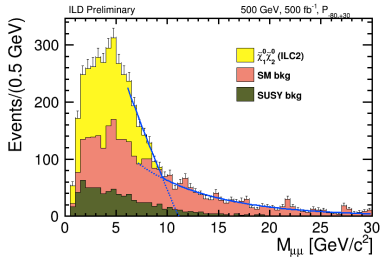


Soft tracks - no problem for ILC

- Event generation Whizard 1.95, hadronisation Pythia 6.422
- Detailed ILD-specific software for simulation and reconstruction (Mokka & Marlin)
- Beam spectrum and ISR included
- Analysis by Jacqueline Yan (KEK)

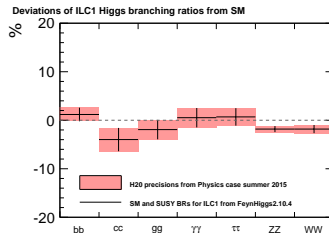
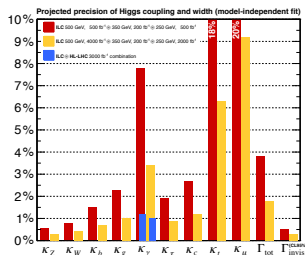
# Mass extraction e.g. $e^+e^- \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0 \mu^+ \mu^-$

- Kinematics: Maximum invariant mass gives the mass splitting. Then maximum of di-electron energy gives the absolute masses since initial state known



# Summary of preliminary analysis results

- Mass precisions  $\sim 0.2\%$  after 20y ( $4 \text{ ab}^{-1}$ )
- Polarised  $\sigma \times \text{BR}$  for  $\tilde{\chi}_2^0$  and  $\tilde{\chi}_1^\pm$  decays with 0.5-2% precision depending on channel and polarisation
- Include Higgs mass ( $\Delta=15 \text{ MeV}$ ) and branching ratios from ILC 20y operation
- Additional relaxable assumption: gluino mass from HL-LHC with 10% precision



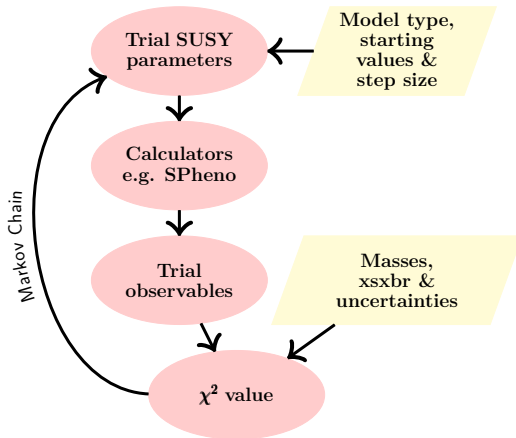
# Fitting parameters to observables with Fittino

Fittino minimises

$$\chi^2 = \left( \frac{\mathcal{O}(ILC) - \mathcal{O}(theory)}{\Delta \mathcal{O}(ILC)} \right)^2$$

(arXiv:hep-ph/0412012)

SUSY SPheno3.3.9beta,  
Higgs mass and BRs  
FeynHiggs2.10.2,  
DM MicrOMEGAs and  
Astrofit

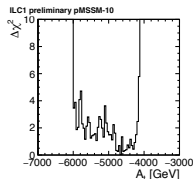
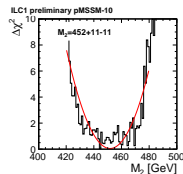
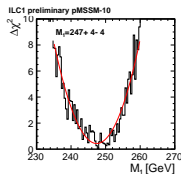




# Fitted model and results: ILC1

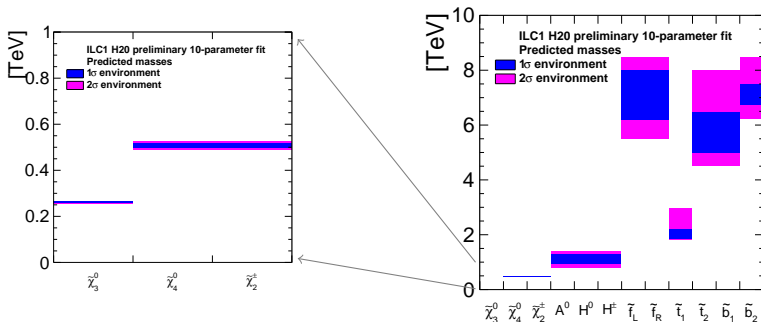
- Underlying model GUT model (NUHM2)
- Fit 10-parameter pMSSM:

parameter	ILC1 pMSSM-10
$M_1$	250
$M_2$	463
$M_3$	1270
$m_A$	1000
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$\mu$	115
$\tan \beta$	10
$M_{\tilde{t}_L}$	4820
$M_{\tilde{t}_R}$	1670
$M_{\text{other sfermions}}$	7150
$A_{t=b=\tau}$	-4400



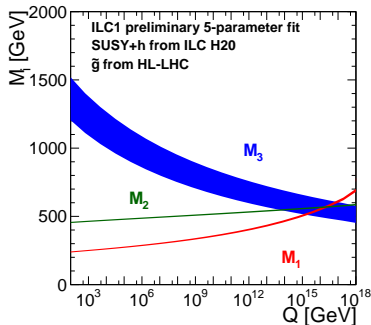
# Predicted masses and dark matter relic density

- Heavier neutralino/chargino masses predicted with 2% uncertainty  
⇒ Sets the scale for energy upgrade e.g. to  $\sqrt{s} \sim 1$  TeV
- Dark matter relic density  $\Omega_{ILC1}/\Omega_{Planck} = 0.054 \pm 0.001$   
⇒ Strong hint that non-SUSY DM or non-thermal production exists



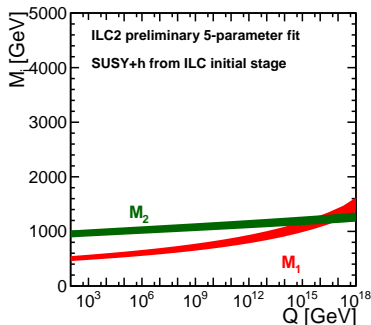
# Test of gaugino mass $M_i$ unification

- Take determined parameters at 1 TeV
- Run up to GUT scale with two-loop RGEs, but fix other parameters to model values
- Underlying theory is NUHM2 and indeed find unification of  $M_i$  at  $10^{16}$  GeV



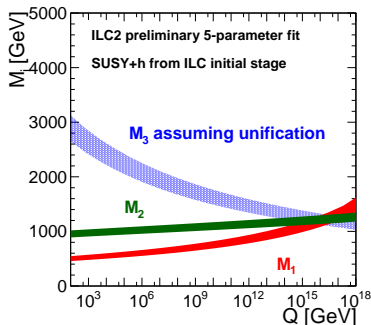
# Extrapolation of $M_3$ assuming unification

- Drop assumption about gluino measurement  $\rightarrow$  ILC2, plot precisions after 5y of ILC
- Find  $Q_U$  s.t.  
 $M_1(Q_U) = M_2(Q_U)$



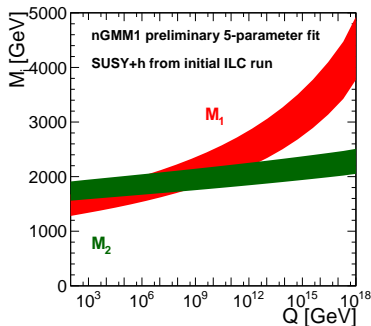
# Extrapolation of $M_3$ assuming unification

- Drop assumption about gluino measurement  $\rightarrow$  ILC2, plot precisions after 5y of ILC
- Find  $Q_U$  s.t.  
 $M_1(Q_U) = M_2(Q_U)$
- If assume  $M_{1/2} = M_3$  and take NUHM2
- Then  $M_3 = 2630 \pm 220$  GeV at 1 TeV
- And prediction for gluino mass  $m_{\tilde{g}} = 2870 \pm 210$  GeV



# Distinguishing between SUSY breaking scenarios

- Motivation for mirage breaking scenario Baer et al 1610.06205
- Gaugino mass parameter unification scale
$$\mu_{mir} = M_{GUT} e^{-8\pi^2/\alpha}$$
- Analysis of scenario with mass gaps  $\sim 5$  GeV ongoing
- Here: expected 4-5y ILC run precisions, clear difference to ILC1 scenario
- Prediction for gluino mass is possible



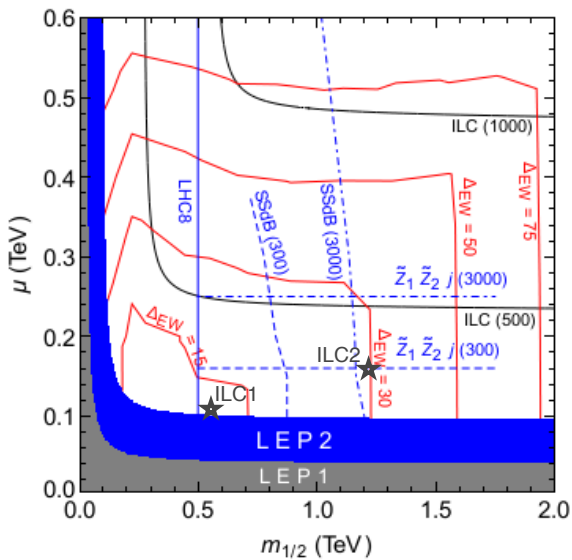
# Summary

- ILC would measure properties of higgsinos to percent-level precision, with 20-year ILC run / threshold scans to sub-% precision in the ILC1 benchmark
- These precise measurements allow for extracting some weak scale parameters, in particular the bino and wino mass parameters with a few % uncertainty after 20 years
- Can obtain predictions for unobserved sparticles, motivating building future lepton and hadron colliders
- Would get a strong hint of the existence of non-SUSY DM
- Can have sensitivity to GUT unification hypotheses

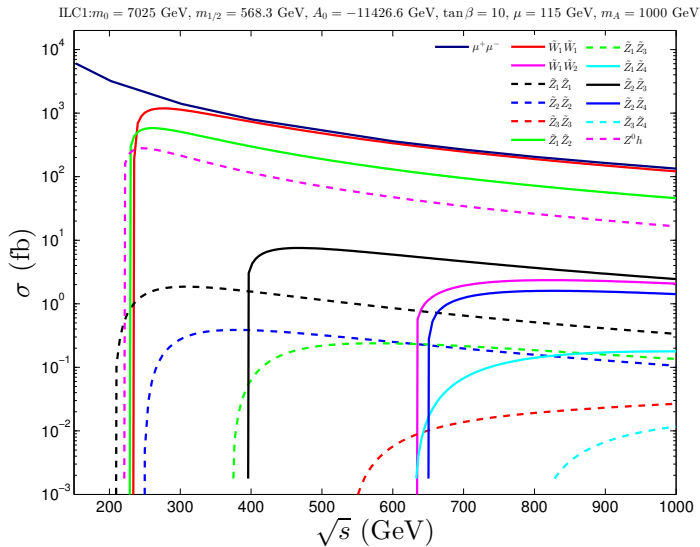


BACKUP





# ILC1 unpolarised cross sections

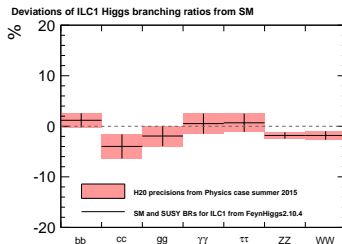
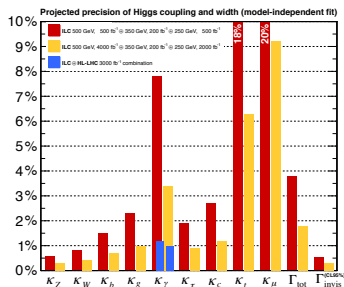


## Fit observables

- ▶ mass  $\tilde{\chi}_1^0, \tilde{\chi}_2^0, \tilde{\chi}_1^\pm$
- ▶ xsxbr of  $\tilde{\chi}_1^+ \tilde{\chi}_1^- \rightarrow q\bar{q}' l \nu_l$  ( $l=e, \mu$ )  
for  $\mathcal{P}(e^- = \mp 80\%, e^+ = \pm 30\%)$
- ▶ xsxbr of  $\tilde{\chi}_1^0 \tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0 ll$  ( $l=e, \mu$ )  
for  $\mathcal{P}(e^- = \mp 80\%, e^+ = \pm 30\%)$
- ▶ Higgs mass
- ▶ Higgs BRs  $h \rightarrow bb, h \rightarrow cc, h \rightarrow \tau\tau, h \rightarrow gg, h \rightarrow \gamma\gamma,$   
 $h \rightarrow ZZ^*, h \rightarrow WW^*$

# Higgs mass and BR measurements from ILC

- ▶ ILC will measure  $m_h$  to 25 MeV (15 MeV) precision in initial (H20) run
- ▶ Precision Higgs measurements included in the fit



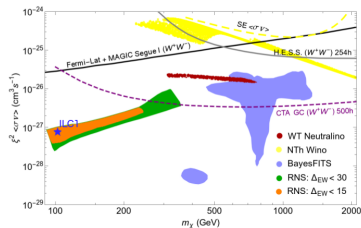
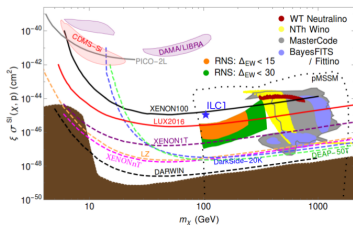
$(h \rightarrow \gamma\gamma \text{ from LHC+ILC})$

# ILC1 pMSSM-10 preliminary fit results

parameter	ILC1 pMSSM-10	Fitted values (higgsinos, Higgs, $m_{\tilde{g}}$ )
$M_1$	250	$247 \pm 4$ GeV
$M_2$	463	$452 \pm 11$ GeV
$M_3$	1270	$1280 \pm 130$ GeV
$m_A$	1000	$1110 \pm 116$ GeV
$\mu$	115	$[115, 122]$ GeV
$\tan \beta$	10	$[8, 18]$
$M_{\tilde{t}_L}$	4820	$[4500, 8000]$ GeV
$M_{\tilde{t}_R}$	1670	$[1500, 2000]$ GeV
$M_{\text{other sfermions}}$	7150	$[5500, 8500]$ GeV
$A_{t=b=\tau}$	-4400	$[-6600, -4200]$ GeV

# Dark matter predictions

- Dark matter relic density  $\Omega_{ILC1}/\Omega_{Planck} = 0.054 \pm 0.001$   
 $\Rightarrow$  Strong hint that non-SUSY DM or non-thermal production exists
- Spin-independent WIMP-nucleon scattering cross section  $\sigma^{SI} = 1.5 \times 10^{-8}$  pb
- WIMP annihilation cross section  $\langle \sigma v \rangle = 2.6 \times 10^{-25} \text{ cm}^3 \text{ s}^{-1}$



Figures from 1609.06735

