Search for high-mass diphoton resonances in proton-proton collisions at 13 TeV and combination with 8 TeV search

The CMS Collaboration*

Abstract

The results of a search are presented for the resonant production of high-mass photon pairs, specifically spin-0 and spin-2 resonances with an invariant mass between 0.5 and 4.5 TeV, and with a width, relative to the mass, between $1.4 \times 10^{-4}$ and $5.6 \times 10^{-2}$. The data sample corresponds to an integrated luminosity of 12.9 fb$^{-1}$ of proton-proton collisions collected with the CMS detector in 2016 at a center-of-mass energy of 13 TeV. No significant excess is observed relative to the standard model expectation. The results of the search are combined statistically with those previously obtained in 2012 and 2015 at $\sqrt{s} = 8$ and 13 TeV, respectively, corresponding to integrated luminosities of 19.7 and 3.3 fb$^{-1}$, to derive exclusion limits on scalar resonances produced through gluon-gluon fusion, and on Randall–Sundrum gravitons. The lower mass limits for Randall–Sundrum gravitons range from 1.95 to 4.45 TeV for coupling parameters between 0.01 and 0.2. These are the most stringent limits on Randall–Sundrum graviton production to date.

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*See Appendix A for the list of collaboration members
1 Introduction

The standard model (SM) of particle physics has been highly successful in describing physical phenomena but it is widely considered to be an incomplete theory because of various shortcomings. In particular, the SM suffers from the so-called hierarchy problem [1], which refers to the large difference between the Higgs boson mass of 125 GeV [2] and the highest energy scale up to which the SM must be valid. Many extensions to the SM have been proposed to address the hierarchy problem, including theories with additional space-like dimensions [3] and models with extended Higgs boson sectors [4]. Some of these extensions predict new resonances that decay to a diphoton final state. For example, the Randall–Sundrum (RS) approach [3, 5] to extra dimensions postulates massive excitations of spin-2 gravitons that can decay to two photons. The simplest extension of the SM Higgs boson sector consists of the addition of a doublet of complex scalar fields. In such models [6], some of these additional scalar resonances can decay to a photon pair [7]. According to the Landau–Yang theorem, the spin of a resonance decaying to two photons can only be zero or an integer larger than one [8, 9].

Recently, the ATLAS and CMS Collaborations at the CERN LHC presented results on searches for high-mass diphoton resonances in proton-proton (pp) collisions at a center-of-mass energy of 13 TeV [10, 11]. The results were based on data collected in 2015, corresponding to integrated luminosities of approximately 3 fb$^{-1}$ per experiment. The CMS results included a combined analysis with pp collision data at $\sqrt{s} = 8$ TeV collected in 2012 [12] corresponding to an integrated luminosity of 19.7 fb$^{-1}$. Both collaborations reported the observation of a moderate excess of events compared to SM expectations, compatible with the production of a new resonance with a mass around 750 GeV.

In this Letter, we report on an updated search for spin-0 resonances and RS gravitons produced in pp collisions and decaying to two photons. The data were collected in 2016 with the CMS detector at $\sqrt{s} = 13$ TeV and correspond to an integrated luminosity of 12.9 fb$^{-1}$. The analysis procedures are very similar to those presented in Ref. [11] for the 2015 data. A combined analysis of the 8 TeV data set of Ref. [12], the 13 TeV data set of Ref. [11], and the 13 TeV data set examined here is performed to improve the sensitivity of the results. Earlier LHC searches for RS gravitons are presented in Refs. [12–28], and for spin-0 particles decaying to two photons in Refs. [12, 29]. These earlier searches are based on pp collisions at either $\sqrt{s} = 7$ or 8 TeV.

2 The CMS detector

The central feature of the CMS apparatus is a superconducting solenoid of 6 m internal diameter, providing a magnetic field of 3.8 T. Within the solenoid volume are a silicon pixel and strip tracker, a lead tungstate crystal electromagnetic calorimeter (ECAL), and a brass and scintillator hadron calorimeter (HCAL). The tracking detectors cover the pseudorapidity range $|\eta| < 2.5$. The ECAL and HCAL, each composed of a barrel and two endcap sections, cover $|\eta| < 3.0$, with the boundary between the barrel and endcaps at around $|\eta| = 1.5$. Forward calorimeters extend the coverage to $|\eta| < 5.0$. The ECAL consists of 75,848 lead tungstate crystals. The barrel section has a granularity $\Delta\eta \times \Delta\phi = 0.0174 \times 0.0174$, with $\phi$ the azimuthal angle, while the endcap sections have a granularity that coarsens progressively up to $\Delta\eta \times \Delta\phi = 0.05 \times 0.05$. Preshower detectors consisting of two planes of silicon sensors interleaved with a total of 3X$_0$ of lead are located in front of the endcap sections. Muons are measured within $|\eta| < 2.4$ by gas-ionization detectors embedded in the steel flux-return yoke outside the solenoid. A more detailed description of the CMS detector, together with a definition of the coordinate system and the relevant kinematic variables, can be found in Ref. [30].
In the barrel section of the ECAL, for photons with energies on the scale of tens of GeV, an energy resolution of about 1% is achieved for unconverted photons and for photons that convert “late”, i.e., just before entering the ECAL. The remaining barrel photons have an energy resolution of about 1.3% up to $|\eta| = 1.0$, rising to about 2.5% for $|\eta| = 1.4$. In the endcaps, the corresponding resolution for unconverted and late-converting photons is about 2.5%, while the remaining endcap photons have a resolution between 3 and 4% [31].

The particle-flow algorithm [32, 33] reconstructs and identifies each individual particle with an optimised combination of information from the various elements of the CMS detector. Particle candidates are classified as either muons, electrons, photons, $\tau$ leptons, charged hadrons, or neutral hadrons.

A two-stage trigger system selects events of interest for the analysis. The level-1 trigger, composed of custom hardware processors, selects events at a maximum readout rate of about 100 kHz using information from the calorimeters and muon detectors. The high-level trigger software algorithms use the full event information to reduce the event rate to less than 1 kHz before data storage.

3 Event simulation

The \textsc{pythia} 8.2 [34] event generator with NNPDF2.3 [35] parton distribution functions (PDFs) is used to produce simulated signal samples of spin-0 and spin-2 resonances decaying to two photons. The samples are generated at leading order (LO), with values of the resonance mass $m_X$ in the range $0.5 < m_X < 4.5$ TeV. Three values of the relative width $\Gamma_X/m_X$ are used as benchmarks: $1.4 \times 10^{-4}$, $1.4 \times 10^{-2}$, and $5.6 \times 10^{-2}$, where $\Gamma_X$ is the width of the resonance. These relative widths correspond, respectively, to resonances much narrower than, comparable to, and significantly wider than the detector resolution. In the context of the RS graviton model, for which $\Gamma_X/m_X = 1.4 \tilde{k}^2$ [36], the relative widths correspond to the dimensionless coupling parameter $\tilde{k} = 0.01, 0.1$, and 0.2. The scalar resonances are produced through gluon-gluon fusion, and RS graviton resonances through both gluon-gluon fusion and quark-antiquark annihilation. In the RS model, the first mechanism accounts for approximately 90% of the production cross section.

The SM background mostly arises from the direct production of two photons, the production of $\gamma$+jets events in which jet fragments are misidentified as photons, and the production of multijet events with misidentified jet fragments. These backgrounds are simulated with the \textsc{sherpa} 2.1 [37], \textsc{madgraph5_aMC@NLO} 2.2 [38] (interfaced with \textsc{pythia} 8.2 for parton showering and hadronization), and \textsc{pythia} 8.2 generators, respectively, using the CT10NLO [39], NNPDF3.0 [40], and NNPDF2.3 PDF sets, again respectively. The \textsc{pythia} tune CUETP8M1 [41] is used.

For both the signal and background samples, the detector response is simulated using the \textsc{geant4} package [42]. The simulated samples incorporate additional pp interactions within the same or a nearby bunch crossing (pileup) and are weighted to reproduce the measured distribution of the number of interactions per bunch crossing.

4 Event selection and diphoton mass spectrum

The trigger requirements, photon identification criteria, and event selection procedures are described in Ref. [11]. Some details are given below. Energy deposits in the ECAL compatible with the shower shape expected for a photon are clustered together to define a photon candi-
date. Variations in the crystal transparency during the data collection period are corrected for using a dedicated monitoring system, and the single-channel response is equalized based on collision data [31]. A multivariate regression technique [31] is used to correct the photon energy for the incomplete containment of the shower in the clustered crystals, the shower losses for photons that convert before reaching the calorimeter, and the effects of pileup. The interaction vertex is selected using the algorithm described in Ref. [43], which combines information on the correlation between the diphoton system and the recoiling tracks, the average transverse momentum ($p_T$) of the recoiling tracks, and, when available, directional information from reconstructed photon conversions. For resonances with a mass above 500 GeV, the fraction of events in which the interaction vertex is correctly assigned is approximately 90%. For each photon candidate, the transverse size of the electromagnetic cluster in the $\eta$ coordinate must be compatible with that expected for a photon from a hard interaction, and the ratio of the associated energy in the HCAL to the photon energy must be less than 0.05.

Photon candidates are required to have $p_T > 75$ GeV and to appear within $|\eta| < 2.5$. Candidates in the transition region between the barrel and endcap detectors ($1.44 < |\eta| < 1.57$), where the acceptance is difficult to model, are rejected. Photon candidates associated with electron tracks that are incompatible with conversion tracks are rejected. Photon candidates are required to be isolated. There are two isolation criteria, both of which are imposed: i) the sum of the scalar $p_T$ of charged hadron candidates from the interaction vertex that lie within a cone of radius $R = \sqrt{(\Delta \eta)^2 + (\Delta \phi)^2} = 0.3$ around the photon candidate must be less than 5 GeV, where charged hadrons identified as conversion tracks associated with the photon candidate are excluded; ii) the pileup-corrected sum of the scalar $p_T$ of additional photon candidates within this same cone must be less than 2.5 GeV.

The identification and trigger efficiencies are measured as functions of photon $p_T$ using data events containing a $Z$ boson decaying to a $\mu^+\mu^-$ pair in association with a photon, or to an $e^+e^-$ pair [31]. The efficiency of the photon selection procedure in the kinematic range considered in the analysis is above 90 and 85% for candidates in the barrel and endcap regions, respectively. The ratio between the efficiencies measured in data and simulation is found to be lower than 1 by 3.5% for photons in the barrel region and by 6.5% for photons in the endcap region. No significant $p_T$ dependence of the efficiency ratios is observed, and a $p_T$-independent correction is applied to the normalization of the simulated event samples to account for this difference.

The photon candidates in an event are grouped into all possible pairs. At least one photon candidate in the pair must have $|\eta| < 1.44$, i.e., lie in the barrel. Photon pairs are divided into two categories. The first category, denoted “EBEB”, contains pairs for which both candidates lie in the barrel. For the second category, denoted “EBEE”, one candidate lies in the barrel and the other in an endcap. The invariant mass $m_{\gamma\gamma}$ of the pair must satisfy $m_{\gamma\gamma} > 230$ GeV for EEBB candidates and $m_{\gamma\gamma} > 330$ GeV for EBEE candidates. The fraction of events in which more than one photon pair satisfies the selection criteria is approximately 1%. In these cases, only the pair with the largest scalar sum of photon $p_T$ is retained.

The selection efficiency for signal events varies between 50 and 70%, depending on the signal hypothesis. Because of the different angular distribution of the decay products, the kinematic acceptance for the RS graviton resonances is lower than for scalar resonances. For $m_X < 1$ TeV the difference is approximately 20%. The two acceptances are similar for $m_X > 3$ TeV.

The event selection procedure described above was finalized on the basis of studies with simulated signal and background event samples prior to inspection of the data in the search region of the diphoton invariant mass distribution, which is defined as $m_{\gamma\gamma} > 500$ GeV.
A total of 6284 (2791) photon pairs are selected in the EBEB (EBEE) category. Of these, 461 (800) pairs have an invariant mass above 500 GeV. According to simulation, the direct production of two photons accounts, respectively, for 90 and 80% of the background events selected in the EBEB and EBEE categories. This prediction is tested in data using the method described in Ref. [44].

The diphoton invariant mass distribution of the selected events is shown in Fig. 1, for both the EBEB and EBEE categories. We perform an independent maximum likelihood fit to the data in each category using the function

$$f(m_{\gamma\gamma}) = m_{\gamma\gamma}^{a+b \log(m_{\gamma\gamma})}. \quad (1)$$

This parametric form is chosen to model the background in the hypothesis tests discussed below. The results of the fits are shown in Fig. 1.

![Figure 1](image-url)  
Figure 1: The observed invariant mass spectra $m_{\gamma\gamma}$ for selected events in the (left) EBEB and (right) EBEE categories. There are no selected events with $m_{\gamma\gamma} > 2000$ GeV. The solid lines and the shaded bands show the results of likelihood fits to the data together with the associated 1 and 2 standard deviation uncertainty bands. The ratio of the difference between the data and the fit to the statistical uncertainty in the data is given in the lower plots.

## 5 Likelihood fit

A simultaneous fit to the invariant mass spectra of events in the EBEB and EBEE event categories is performed to determine the compatibility of the data with the background-only and the signal+background hypotheses. The test statistic is based on the profile likelihood ratio:

$$q(\mu) = -2 \log \frac{L(\mu S + B|\hat{\theta})}{L(\hat{\mu} S + B|\hat{\theta})}, \quad (2)$$

where $S$ and $B$ represent the probability density functions for resonant diphoton production and for the SM background, respectively. The parameter $\mu$ is the so-called signal strength, while
\(\tilde{\theta}\) represents the nuisance parameters of the model, used to account for systematic uncertainties. The \(\hat{x}\) notation indicates the best fit value of the parameter \(x\) for any \(y\) value, while \(\hat{x}_y\) denotes the best fit value of \(x\) for a fixed value \(y\).

To set upper limits on the rate of resonant diphoton production, the modified frequentist method known as CLs \([45, 46]\) is used, following the prescription described in Ref. \([47]\). The compatibility of the observation with the background-only hypothesis is evaluated by computing the background-only \(p\)-value. The latter is defined as the probability, in the background-only hypothesis, for \(q(0)\) to exceed the value observed in data. This quantity, the “local \(p\)-value” \(p_0\), does not take into account the fact that many signal hypotheses are tested.

Asymptotic formulas \([48]\) are used in the calculations of exclusion limits and local \(p\)-values. The accuracy of the asymptotic approximation in the estimation of exclusion limits and significance is studied, using pseudo-experiments, for a subset of the hypothesis tests and is found to be about 10%.

The signal shape in \(m_{\gamma\gamma}\) is determined from the convolution of the intrinsic shape of the resonance and the CMS detector response to photons. The intrinsic shape is taken from the PYTHIA 8.2 generator. A grid of mass points with 125 GeV spacing, in the range 500–4500 GeV, is used. The resulting shapes are interpolated to intermediate points using a parametric description of the distribution. The detector response is determined using fully simulated signal samples of small intrinsic width, corrected through Gaussian smearing to agree with measurements based on \(Z \rightarrow e^+e^-\) data. Nine uniformly spaced mass hypotheses in the range 500–4500 GeV are employed. The signal mass resolution, quantified through the ratio of the full width at half maximum of the distribution, divided by 2.35, to the peak position, is approximately 1.0 and 1.5% for the EEBB and EBEE categories, respectively. The signal normalization coefficients are proportional to the product of the kinematic acceptance and the signal efficiency within the acceptance region. These are computed, for each category, in simulated samples and interpolated to intermediate points using quadratic functions of \(m_X\) and \(\Gamma_X/m_X\).

The background shape in \(m_{\gamma\gamma}\) is described by the parametric function given by Eq. (1). The values of the parameters \(a\) and \(b\) are determined in the fit to data, with separate values for the EEBB and EBEE categories, and are treated as unconstrained nuisance parameters in the hypothesis tests.

The accuracy of the background parameterization is assessed using simulation and is quantified by studying the difference between the true and predicted numbers of background events in several \(m_{\gamma\gamma}\) intervals in the search region. The relative widths of the intervals, defined by \(2(x_1 - x_2)/(x_1 + x_2)\) with \(x_1\) and \(x_2\) the lower and upper bin edges, range between 2 and 15%. Pseudo-experiments are drawn from the mass spectrum predicted by the simulation and are fit with the chosen background model. The total number of events in each pseudo-experiment is taken from a Poisson distribution whose mean is set equal to the observation in data. For each interval, the distribution of the pull variable, defined as the difference between the true and predicted numbers of events divided by the estimated statistical uncertainty, is constructed. If the absolute value of the median of this distribution is found to be above 0.5 in an interval, an additional uncertainty is assigned to the background parametrization. A modified pull distribution is then constructed, increasing the statistical uncertainty in the fit by an extra term, denoted the “bias term”. The bias term is parametrized as a smooth function of \(m_{\gamma\gamma}\), which is tuned in such a manner that the absolute value of the median of the modified pull distribution is less than 0.5 in all intervals. The amplitude of the bias term function is comparable to that of the 1 standard deviation bands in Fig. 1. This additional uncertainty is included in the likelihood function by adding to the background model a component having the same shape as
the signal. The normalization coefficient of this component is constrained to have a Gaussian distribution of mean zero, with a width equal to the integral of the bias term function over the full width at half maximum of the tested signal shape. The inclusion of this additional component has the effect of avoiding falsely positive or falsely negative tests that could be induced by a mismodeling of the background shape, and it reduces the sensitivity of the analysis by at most 10%.

6 Systematic uncertainties

The impact of systematic uncertainties in this analysis is smaller than that of the statistical uncertainties. The parametric background model has no associated systematic uncertainties except for the bias term uncertainty described in the previous section. Since the background shape coefficients $a$ and $b$ [Eq. (1)] are treated as unconstrained nuisance parameters, the associated uncertainties are statistical in nature.

The systematic uncertainties in the signal normalization associated with the integrated luminosity, the selection efficiency, and the PDFs are 6.2, 6.0, and 6.0%, respectively. The uncertainty in the integrated luminosity is estimated from beam scans performed in August 2016, utilizing the methods of Ref. [49]. The uncertainty associated with the PDFs is evaluated by comparing the overall selection efficiency obtained with the CT10 [39], MSTW08 [50], and NNPDF2.3 [35] PDF sets and taking the largest deviation over all tested signal hypotheses. A 1% uncertainty is associated with the level of knowledge of the energy scale and accounts for the uncertainty in the energy scale at the $Z$ boson peak and its extrapolation to higher masses. A 10% uncertainty is assigned to the knowledge of the photon energy resolution, corresponding to the uncertainty in the estimated additional Gaussian smearing determined at the $Z$ boson peak.

7 Results for the 2016 data

The observed and expected 95% confidence level (CL) upper limits on the product of the production cross section ($\sigma_{13\text{ TeV}}$) and branching fraction to two photons ($B_{\gamma\gamma}$) for scalar and RS graviton resonances are shown in Fig. 2. Using the LO cross sections from PYTHIA 8.2, RS gravitons with masses below 1.75, 3.75, and 4.35 TeV are excluded for $\tilde{k} = 0.01$, 0.1, and 0.2, respectively, corresponding to $\Gamma_X/m_X = 1.4 \times 10^{-4}$, $1.4 \times 10^{-2}$, and $5.6 \times 10^{-2}$.

The value of $p_0$ for different signal hypotheses is shown in Fig. 3. The largest excess is observed for $m_X \approx 620$ GeV, and has a local significance of approximately 2.4 and 2.7 standard deviations for narrow spin-0 and RS graviton signal hypotheses, respectively. After taking into account the effect of searching for several signal hypotheses, i.e., searching over a range of widths and masses, the significance of the excess is reduced to less than one standard deviation. No excess is observed in the proximity of $m_X = 750$ GeV.

8 Combination with the 2012 and 2015 data

The results obtained for the 2016 data are combined statistically with those obtained for the data discussed in Ref. [11], namely 19.7 fb$^{-1}$ of proton-proton collisions recorded at $\sqrt{s} = 8$ TeV in 2012 [12] and 3.3 fb$^{-1}$ recorded at $\sqrt{s} = 13$ TeV in 2015. For a portion of the 2015 data (0.6 fb$^{-1}$), the CMS magnet was off (0 T), while for the rest of the 2015 data and for all of the 2012 and 2016 data, the magnet was at its operational field strength (3.8 T). The analysis of the 0 T data from 2015 is described in Ref. [11].
Figure 2: The 95% CL upper limits on the production of diphoton resonances as a function of the resonance mass $m_X$, from the analysis of data collected in 2016. Exclusion limits for the scalar and RS graviton signals are given by the grey (darker) and green (lighter) curves, respectively. The observed limits are shown by the solid lines, while the median expected limits are given by the dashed lines together with their associated 1 standard deviation uncertainty bands. The leading-order production cross section for diphoton resonances in the RS graviton model is shown for three values of the dimensionless coupling parameter $k$ together with the exclusion upper limits calculated for the corresponding three values of the width relative to the mass, $\Gamma_X/m_X$. Shown are the results for (upper) a narrow width, (middle) an intermediate-width, and (lower) a broad resonance.

The procedure followed for the combined analysis of 8 and 13 TeV data is the same as in Ref. [11]. The ratio of the 8 to the 13 TeV production cross section is computed using PYTHIA 8.2, for the two types of signal hypotheses considered: scalar resonances and RS graviton resonances. The cross section ratio decreases from 0.27 and 0.29 at $m_X = 500$ GeV to 0.03 and 0.04 at $m_X = 4$ TeV, for the scalar and RS graviton resonance hypotheses, respectively.

Exclusion limits are set on the 13 TeV production cross section for both models, and background-only $p$-values are computed for the signal hypotheses.

The correlation model between the systematic uncertainties associated with 8 and 13 TeV data is described in Ref. [11]. It assumes all uncertainties to be uncorrelated except for those related to the knowledge of the photon energy scale, which are taken to have a linear correlation of 0.5, and those related to the knowledge of the PDFs, which are taken to be fully correlated. For the combination of the two 13 TeV data sets, the background shape and the associated bias term uncertainties are assumed to be fully correlated between the corresponding categories of the
2015 (3.8 T) and 2016 data. Independent background normalization coefficients are used for the two data sets. The uncertainty in the signal selection efficiency is taken to be uncorrelated between the 2015 and 2016 data. The uncertainty in the knowledge of the integrated luminosity is treated as follows: a 2.3% uncertainty, corresponding to the knowledge of the absolute luminosity scale calibration determined with beam scans, is taken to be fully correlated between the 2015 (3.8 T) and 2016 data, and additional uncertainties of 1.5 and 5.8%, corresponding to the uncertainty in extrapolating the scale calibration to the data collection conditions, are applied, again respectively. Finally, the photon energy scale uncertainties are taken to be fully correlated between the two data sets.

Figure 4 shows the observed and expected 95% CL upper limits on the 13 TeV production cross section of the different signal hypotheses obtained with the combined analysis of the 13 TeV data recorded in 2015 and 2016. The upper limits on the production of scalar resonances decaying to two photons range from about 10 to 0.2 fb, for resonance masses between 0.5 and 4.5 TeV. Compared to the 2016 data alone, the sensitivity is improved by approximately 10 and 20% at the high and low end of the $m_X$ search region, respectively. Using the LO cross sections from PYTHIA 8.2, RS gravitons with masses below 3.85 and 4.45 TeV are excluded for $\tilde{k} = 0.1$ and 0.2, respectively. For $\tilde{k} = 0.01$, graviton masses below 1.95 TeV are excluded, except for the region between 1.75 and 1.85 TeV.

The observed $p_0$ for $\Gamma_X/m_X = 1.4 \times 10^{-4}$ and $5.6 \times 10^{-2}$ obtained with the combined analysis of the 2015 and 2016 data is shown in Fig. 5. The largest excess is observed for $m_X \approx 1.3$ TeV and has a local significance of about 2.2 standard deviations, corresponding to less than 1 standard deviation after accounting for the effect of searching for several signal hypotheses. For
$m_X = 750$ GeV, the 2.9 standard deviation local significance excess observed in the 2015 data is reduced to 0.8 standard deviations.

Figure 4: The 95\% CL upper limits on the production of diphoton resonances as a function of the resonance mass $m_X$, from the combined analysis of data collected in 2015 and in 2016. Exclusion limits for the scalar and RS graviton signals are given by the grey (darker) and green (lighter) curves, respectively. The observed limits are shown by the solid lines, while the median expected limits are given by the dashed lines together with their associated 1 standard deviation uncertainty bands. The leading-order production cross section for diphoton resonances in the RS graviton model is shown for three values of the dimensionless coupling parameter $\tilde{k}$ together with the exclusion upper limits calculated for the corresponding three values of the width relative to the mass, $\Gamma_X/m_X$. Shown are the results for (upper) a narrow width, (middle) an intermediate-width, and (lower) a broad resonance.

The observed and expected 95\% CL upper limits on the 13 TeV signal production cross sections obtained through a combined analysis of the 8 TeV data from 2012 and the 13 TeV data from 2015 and 2016 are shown in Fig. 6. Compared to the combined 13 TeV data, the analysis sensitivity improves by about 10\% at the low end of the $m_X$ range, while the improvement is negligible at the higher end of the range. Thus the lower limits on the mass of RS gravitons obtained by combining the 8 and 13 TeV data coincide with those obtained with the 13 TeV data alone.

The observed $p_0$ for $\Gamma_X/m_X = 1.4 \times 10^{-4}$ and $5.6 \times 10^{-2}$ obtained with the combined 8 and 13 TeV analysis is shown in Fig. 7. The largest excess, observed for $m_X \approx 0.9$ TeV, has a local significance of about 2.2 standard deviations, corresponding to less than 1 standard deviation overall. For $m_X = 750$ GeV, the 3.4 standard deviation local significance excess reported in Ref. [11] is reduced to about 1.9 standard deviations.
A search for the resonant production of high-mass photon pairs has been presented. The analysis is based on a sample of proton-proton collisions collected by the CMS experiment in 2016 at $\sqrt{s} = 13$ TeV, corresponding to an integrated luminosity of 12.9 fb$^{-1}$. Events containing two photon candidates with transverse momenta above 75 GeV are selected. The diphoton mass spectrum above 500 GeV is examined for evidence of the production of high-mass spin-0 and spin-2 resonances.

Limits on the production of scalar resonances and Randall–Sundrum gravitons in the range $0.5 < m_X < 4.5$ TeV and $1.4 \times 10^{-4} < \Gamma_X/m_X < 5.6 \times 10^{-2}$ are determined using the modified frequentist approach, where $m_X$ and $\Gamma_X$ are the resonance mass and width, respectively. The results obtained with the 2016 data set are combined statistically with those obtained in 2012 and 2015, corresponding to integrated luminosities of 19.7 and 3.3 fb$^{-1}$ of data recorded at $\sqrt{s} = 8$ and 13 TeV, respectively.

No significant excess is observed above the predictions of the standard model. Using the leading-order cross sections, Randall–Sundrum gravitons with masses below 3.85 and 4.45 TeV are excluded for values of the dimensionless coupling parameter $\kappa = 0.1$ and 0.2, respectively. For $\kappa = 0.01$, graviton masses below 1.95 TeV are excluded, except for the region between 1.75 and 1.85 TeV. These are the most stringent limits on Randall–Sundrum graviton production to date.
Figure 6: The 95% CL upper limits on the production of diphoton resonances as a function of the resonance mass $m_X$, from the combined analysis of the 8 and 13 TeV data. The 8 TeV results are scaled by the ratio of the 8 to 13 TeV cross sections. Exclusion limits for the scalar and RS graviton signals are given by the grey (darker) and green (lighter) curves, respectively. The observed limits are shown by the solid lines, while the median expected limits are given by the dashed lines together with their associated 1 standard deviation uncertainty bands. The leading-order production cross section for diphoton resonances in the RS graviton model is shown for three values of the dimensionless coupling parameter $\bar{k}$ together with the exclusion upper limits calculated for the corresponding three values of the width relative to the mass, $\Gamma_X/m_X$. Shown are the results for (upper) a narrow width, (middle) an intermediate-width, and (lower) a broad resonance.

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Figure 7: Observed background-only $p$-values for resonances with (upper) $\Gamma_X/m_X = 1.4 \times 10^{-4}$ and (lower) $5.6 \times 10^{-2}$ as a function of the resonance mass $m_X$, from the combined analysis of the 8 and 13 TeV data. The results obtained for the two individual center-of-mass energies are also shown. The curves corresponding to the scalar and RS graviton hypotheses are shown in left and right columns, respectively. The insets show an expanded region around $m_X = 750$ GeV.

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57: Also at Ozyegin University, Istanbul, Turkey
58: Also at Izmir Institute of Technology, Izmir, Turkey
59: Also at Marmara University, Istanbul, Turkey
60: Also at Kafkas University, Kars, Turkey
61: Also at Istanbul Bilgi University, Istanbul, Turkey
62: Also at Yildiz Technical University, Istanbul, Turkey
63: Also at Hacettepe University, Ankara, Turkey
64: Also at Rutherford Appleton Laboratory, Didcot, United Kingdom
65: Also at School of Physics and Astronomy, University of Southampton, Southampton, United Kingdom
66: Also at Instituto de Astrofisica de Canarias, La Laguna, Spain
67: Also at Utah Valley University, Orem, USA
68: Also at Argonne National Laboratory, Argonne, USA
69: Also at Erzincan University, Erzincan, Turkey
70: Also at Mimar Sinan University, Istanbul, Istanbul, Turkey
71: Now at The Catholic University of America, Washington, USA
72: Also at Texas A&M University at Qatar, Doha, Qatar
73: Also at Kyungpook National University, Daegu, Korea