Computing the hadronic vacuum polarization function by analytic continuation

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Abstract

We propose a method to compute the hadronic vacuum polarization function on the lattice at continuous values of photon momenta bridging between the spacelike and timelike regions. We provide two independent demonstrations to show that this method leads to the desired hadronic vacuum polarization function in Minkowski spacetime. We show with the example of the leading-order QCD correction to the muon anomalous magnetic moment that this approach can provide a valuable alternative method for calculations of physical quantities where the hadronic vacuum polarization function enters.

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I. INTRODUCTION

In quantum field theory, the photon receives vacuum polarization corrections, which modify the electromagnetic coupling constant depending on the virtual photon momentum. Quark-loop contributions to these corrections include the effects of the strong interaction through further gluon exchanges. Since the strong interaction becomes nonperturbative at low energies, the calculation of these hadronic vacuum polarization (HVP) effects requires a nonperturbative method to treat Quantum Chromodynamics (QCD).

The QCD corrections from the HVP appear in many physical observables, such as the running QED coupling, the weak mixing angle, the $2P - 2S$ Lamb shift of muonic hydrogen and the muon anomalous magnetic moment, a_μ , [1–3]. In fact, a_μ is a prime example where the HVP correction is very important since there is a discrepancy between the experimental measurement of this quantity and the Standard Model (SM) prediction, and it is the HVP correction that constitutes the dominant theoretical uncertainty. Since it is tempting to interpret the discrepancy as an indication for new physics beyond the SM, it is necessary to have a well-controlled QCD calculation of the HVP function before drawing any definitive conclusions.

At low energies, the running QCD coupling becomes large and perturbative QCD fails to be a reliable tool for determining the HVP function. The established approach of using a dispersion relation together with experimental data for e^+e^- scattering and τ decay depends on the experimental input and hence cannot be considered as an *ab initio* SM calculation. It is therefore highly desirable to perform a nonperturbative calculation of the HVP function from first principles using lattice QCD. Indeed, for the quantities listed above, where the HVP function is an important ingredient, it has been shown that lattice QCD can provide a substantial, if not an essential, contribution; see Refs. [3, 4] for a more detailed discussion.

In standard lattice calculations, the HVP function is calculated at discrete spacelike momenta by performing a Fourier transform of the Euclidean vector-vector correlation function. However, these lattice QCD computations suffer from a generic difficulty, namely, that low momenta are usually not directly accessible. For example, in the determination of the hadronic contribution a_μ^{hvp} to the muon anomalous magnetic moment, momenta close to 0.003 GeV^2 dominate. Accessing the HVP function at small or near-zero momentum is a general concern relevant to many observables to be computed in lattice QCD. To ad-

dress this problem, efforts to approach the low momentum region by using partially twisted boundary conditions have been undertaken [5]. In addition, a number of approaches have been employed to describe the HVP function on the lattice. Here, well-motivated [5–10] or even model-independent [11] functional forms to fit the results computed at discrete lattice momenta have been used. They lead to smooth functions that describe well the desired momentum region, including zero momentum, needed to obtain the renormalized HVP function. A step forward has been discussed in Ref. [12] where it was suggested to calculate the derivative of the relevant correlation function by using the sequential source propagator method to obtain the zero momentum contribution.

In this paper we discuss a more general approach that allows us to compute the HVP function at continuous momenta, both in the spacelike and timelike regions. Following the ideas employed in Ref. [13] this will be achieved by starting with the Euclidean vector-vector correlation function and performing a Fourier transform only in the spatial directions and an integration (in practice, a summation) in the time direction with a factor of $e^{\omega t}$. In this way, we are able to calculate the HVP function at very small spacelike momenta, covering also the zero momentum value and even extending in to the timelike region. We will provide two independent demonstrations that this leads to a sound and valid evaluation of the HVP function. In order to test whether this approach is practical and leads to small errors in the calculation, we apply this method to a realistic lattice calculation of the HVP function and the associated a_μ^{hvp} . We stress that we consider our approach as an alternative way to compute the HVP function in lattice QCD at small momenta. It has the advantage of avoiding assumptions on the analytic form used to describe the HVP function. However, as we will see below, it will not lead to an increased precision in the calculation of physical quantities such as a_μ^{hvp} . Nevertheless, we believe that our new method can serve as a valuable alternative to the presently employed techniques, as we do not have to model the functional form of the HVP in the low-momentum region.

In Sec. II we introduce the analytic continuation method. We then demonstrate in Secs. III and IV the validity of the proposed method using two different approaches. In Sec. V we perform a computation of the HVP function based on twisted mass fermions, and in Sec. VI we show with the example of a_μ^{hvp} how the analytic continuation method is used to determine a physical observable.

II. ALTERNATIVE METHOD

In Minkowski spacetime, the HVP function, denoted as $\Pi^M(k^2)$, is defined by

$$\Pi^M(k^2)(k_\mu k_\nu - g_{\mu\nu} k^2) = i \int d^4x e^{-ikx} \langle \Omega | T \{ J_\mu^M(x) J_\nu^M(0) \} | \Omega \rangle , \quad (1)$$

where $J_\mu^M(x)$ is the hadronic component of the electromagnetic vector current defined in Minkowski spacetime with $x = (x_0, \vec{x})$ and $kx = k_0 x_0 - \vec{k} \cdot \vec{x}$, k is the photon momentum with $k^2 = k_0^2 - \vec{k}^2 > 0$ for timelike and $k^2 < 0$ for spacelike momenta, and $|\Omega\rangle$ is the QCD vacuum state.

In Euclidean spacetime, the HVP function is given in a similar way

$$\Pi^E(K^2)(K_\mu K_\nu - \delta_{\mu\nu} K^2) = \int d^4X e^{iKX} \langle \Omega | T \{ J_\mu^E(X) J_\nu^E(0) \} | \Omega \rangle , \quad (2)$$

with the vector current $J_\mu^E(X)$ defined in the Euclidean spacetime with $X = (\vec{x}, t)$. Here the momenta $K = (\vec{k}, K_t)$ are spacelike with $K^2 = \vec{k}^2 + K_t^2 > 0$.

By modifying Eq. (2) we propose to calculate the HVP function in another way: we perform a Fourier transform only in the spatial directions but integrate in the time direction with a factor of $e^{\omega t}$,

$$\bar{\Pi}(K^2)(K_\mu K_\nu - \delta_{\mu\nu} K^2) = \int dt e^{\omega t} \int d^3\vec{x} e^{i\vec{k}\vec{x}} \langle \Omega | T \{ J_\mu^E(\vec{x}, t) J_\nu^E(\vec{0}, 0) \} | \Omega \rangle . \quad (3)$$

In Eq. (3), the momentum K is given by $K = (\vec{k}, -i\omega)$, with $\vec{k}$ the spatial momentum and ω an input parameter (corresponding here to the photon energy). By varying ω , we can achieve values for $K^2 = -\omega^2 + \vec{k}^2$ that enter both the spacelike and timelike momentum regions. In particular, using Eq. (3) we can compute $\bar{\Pi}(K^2)$ at zero momentum or near $K^2 = 0.003 \text{ GeV}^2$ without an extrapolation in the momentum. A very important restriction is that the energy ω must satisfy

$$-K^2 = \omega^2 - \vec{k}^2 < M_V^2 , \quad \text{or} \quad \omega < E_V , \quad (4)$$

where E_V is the lowest energy in the vector channel and M_V is the corresponding invariant mass. Restricting the values of ω in this way, a mixing between the photon state and the hadronic states is avoided. Furthermore, the divergence caused by $e^{\omega t}$ at infinitely large t is eliminated by a suppression factor $e^{-E_V t}$ arising from the asymptotic time dependence of $\langle J_\mu^E(t) J_\nu^E(0) \rangle$ thus rendering the integral of Eq. (3) convergent. In the following two sections,

we will provide two independent demonstrations that Eq. (3) is the analytic continuation of Eq. (1) from Minkowski spacetime to Euclidean spacetime, and therefore $\bar{\Pi}(K^2)$ defined in Eq. (3) can be used directly to compute the HVP function and physical quantities where the HVP function enters.

III. DEMONSTRATION USING THE PHOTON-VECTOR CURRENT TRANSITION AMPLITUDE

To demonstrate that Eq. (3) is a theoretically valid definition for the HVP function, we follow Ref. [13] closely. We first consider a matrix element in Minkowski spacetime,

$$\langle \gamma(k, \lambda) | J_\nu^M(0) | \Omega' \rangle, \quad (5)$$

where $\langle \gamma(k, \lambda) |$ is a photon state with momentum k and polarization λ . It is normalized according to $\langle \gamma(k, \lambda) | \gamma(k', \lambda') \rangle = (2\omega) \delta_{\lambda\lambda'} (2\pi)^3 \delta^{(3)}(\vec{k} - \vec{k}')$ with ω the photon energy. $|\Omega'\rangle$ denotes the vacuum state in a theory that contains both QCD and QED while $|\Omega\rangle$ is the QCD vacuum state with a trivial QED sector.

Using the Lehmann–Symanzik–Zimmermann reduction formula, we convert the matrix element in Eq. (5) into a correlation function,

$$\langle \gamma(k, \lambda) | J_\nu^M(0) | \Omega' \rangle = i \lim_{k' \rightarrow k} \varepsilon_\mu^{M*}(k, \lambda) k'^2 \int d^4x e^{-ik'x} \langle \Omega' | T \{ A_\mu^M(x) J_\nu^M(0) \} | \Omega' \rangle, \quad (6)$$

where $A_\mu^M(x)$ is the photon field defined in Minkowski spacetime. It is normalized according to $\langle \Omega' | A_\mu^M(x) | \gamma(k, \lambda) \rangle = \varepsilon_\mu^M(k, \lambda) e^{ikx}$, and $\varepsilon_\mu^M(k, \lambda)$ is the polarization vector in Minkowski spacetime. In perturbative QED, we expand $\exp(iS_{\text{int}}) = 1 + iS_{\text{int}} + \dots$, with $S_{\text{int}} = e \int d^4x A_\mu^M(x) J_\mu^M(x)$, where e is the electron charge. At $O(e)$ we have

$$\begin{aligned} \langle \gamma(k, \lambda) | J_\nu^M(0) | \Omega' \rangle = \\ -e \lim_{k' \rightarrow k} \varepsilon_\mu^{M*}(k, \lambda) k'^2 \int d^4x d^4y e^{-ik'x} \langle \Omega' | T \{ A_\mu^M(x) A_\rho^M(y) J_\rho^M(y) J_\nu^M(0) \} | \Omega' \rangle. \end{aligned} \quad (7)$$

The Wick contraction of the photon fields $A_\mu^M(x) A_\rho^M(y)$ yields a propagator $D_{\mu\rho}^M(x, y) = -ig_{\mu\rho} \int \frac{d^4l}{(2\pi)^4} \frac{e^{-il(x-y)}}{l^2 + i\varepsilon}$, which cancels the inverse propagator k'^2 outside the integral and results in

$$\langle \gamma(k, \lambda) | J_\nu^M(0) | \Omega' \rangle = i e \varepsilon_\mu^{M*}(k, \lambda) \int d^4y e^{-iky} \langle \Omega | T \{ J_\mu^M(y) J_\nu^M(0) \} | \Omega \rangle. \quad (8)$$

On the other hand, we can define a matrix element $M_{\mu\nu}(t_0, \vec{k})$ in Euclidean spacetime and examine the $t_0 \rightarrow \infty$ limit,

$$\begin{aligned}
M_{\mu\nu}(t_0, \vec{k}) &= \int d^3\vec{x} e^{i\vec{k}\vec{x}} \langle \Omega' | T \{ A_\mu^E(\vec{x}, t_0) J_\nu^E(0) \} | \Omega' \rangle \\
&\rightarrow \sum_\lambda \langle \Omega' | A_\mu^E(\vec{0}, 0) | \gamma(k, \lambda) \rangle \frac{e^{-\omega t_0}}{2\omega} \langle \gamma(k, \lambda) | J_\nu^E(0) | \Omega' \rangle \\
&= \sum_\lambda \varepsilon_\mu^E(k, \lambda) \frac{e^{-\omega t_0}}{2\omega} \langle \gamma(k, \lambda) | J_\nu^E(0) | \Omega' \rangle, \tag{9}
\end{aligned}$$

where $\omega = |\vec{k}|$ is the on shell photon energy and $A_\mu^E(\vec{x}, 0)$ is normalized according to $\langle \Omega' | A_\mu^E(\vec{x}, 0) | \gamma(k, \lambda) \rangle = \varepsilon_\mu^E(k, \lambda) e^{-i\vec{k}\vec{x}}$. In the second line, we have inserted the ground state $|\gamma(k, \lambda)\rangle$ into the matrix element and neglected the three-photon states, which are suppressed by powers of the QED coupling. We also neglected hadronic states because they are suppressed by a factor of $e^{-(E_V - \omega)t_0}$. Here we have used the restriction $\omega < E_V$.

The matrix element $M_{\mu\nu}(t_0, \vec{k})$ can be analyzed by again using leading-order perturbative QED and integrating out the photon field in Eq. (9). It results in

$$M_{\mu\nu}(t_0, \vec{k}) = e \int dt \int d^3\vec{y} e^{i\vec{k}\vec{y}} D_{\mu\rho}^E(\vec{k}, t_0 - t) \langle \Omega | T \{ J_\rho^E(\vec{y}, t) J_\nu^E(0) \} | \Omega \rangle, \tag{10}$$

where $D_{\mu\rho}^E(X) = \int \frac{d^4K}{(2\pi)^4} e^{iKX} D_{\mu\rho}^E(K)$ is a photon propagator defined in Euclidean spacetime. Using the integral

$$\mathcal{I} = \int_{-\infty}^{\infty} \frac{dK_t}{2\pi} D_{\mu\nu}^E(K) e^{iK_t t} = \int_{-\infty}^{\infty} \frac{dK_t}{2\pi} \frac{\delta_{\mu\nu}}{K_t^2 + \vec{k}^2} e^{iK_t t} = \frac{e^{-|\vec{k}||t|}}{2|\vec{k}|} \delta_{\mu\nu} = \frac{e^{-\omega|t|}}{2\omega} \delta_{\mu\nu}, \tag{11}$$

we can write

$$M_{\mu\nu}(t_0, \vec{k}) = e \int dt \frac{e^{-\omega|t_0-t|}}{2\omega} \int d^3\vec{y} e^{i\vec{k}\vec{y}} \langle \Omega | T \{ J_\mu^E(\vec{y}, t) J_\nu^E(0) \} | \Omega \rangle. \tag{12}$$

We are ultimately interested in the limit of $t_0 \rightarrow \infty$ for which $M_{\mu\nu}(t_0, \vec{k})$ becomes

$$M_{\mu\nu}(t_0, \vec{k}) \rightarrow e \int dt \frac{e^{-\omega(t_0-t)}}{2\omega} \int d^3\vec{y} e^{i\vec{k}\vec{y}} \langle \Omega | T \{ J_\mu^E(\vec{y}, t) J_\nu^E(0) \} | \Omega \rangle. \tag{13}$$

Combining this with Eq. (9) yields

$$\langle \gamma(k, \lambda) | J_\nu^E(0) | \Omega' \rangle = e \varepsilon_\mu^{E*}(k, \lambda) \int dt e^{\omega t} \int d^3\vec{y} e^{i\vec{k}\vec{y}} \langle \Omega | T \{ J_\mu^E(\vec{y}, t) J_\nu^E(0) \} | \Omega \rangle. \tag{14}$$

The left-hand sides of Eqs. (8) and (14) are the matrix elements $\langle \gamma(k, \lambda) | J_\nu^{M(E)}(0) | \Omega' \rangle$, which are equivalent physical observables in Minkowski and Euclidean spacetime. On the other

hand, the right-hand sides of Eqs. (8) and (14) are the HVP functions defined in Eq. (1) and (3), respectively. This demonstrates that these two definitions are equivalent.

In the above analysis, we have looked at the special case that the photon is on shell, with $k^2 = \omega^2 - \vec{k}^2 = 0$. In the case that the photon is off shell, we can replace the matrix element in Eq. (5) by

$$\langle W(k, \lambda) | J_\nu^M(0) | \Omega' \rangle , \quad (15)$$

where W denotes a massive vector boson, which has the same quantum numbers as the photon but a nonzero mass M_W . Therefore, its momentum k satisfies $k^2 = \omega^2 - \vec{k}^2 = M_W^2 > 0$. Using a similar line of arguments as above, we can then show that the analytic continuation, as outlined here, is valid for momenta $k^2 > 0$. In the next section, we will demonstrate that the method is also valid for $k^2 < 0$.

IV. DEMONSTRATION FROM TEMPORAL MOMENTS

In this section we will demonstrate the validity of Eq. (3) using a Taylor expansion and the introduction of temporal moments. This technique has been previously used for lattice [14] and perturbative calculations [15, 16]. In Ref. [4] it has been proposed, to also use it for the calculation of the HVP function and a_μ^{hvp} .

We write Eq. (2) as follows:

$$\begin{aligned} \Pi^E(K^2) F_{\mu\nu}(K) &= \int dt e^{iK_t t} \int d^3\vec{x} e^{i\vec{k}\vec{x}} \langle J_\mu^E(\vec{x}, t) J_\nu^E(\vec{0}, 0) \rangle , \\ &= \int dt e^{iK_t t} C_{\mu\nu}(\vec{k}, t) , \end{aligned} \quad (16)$$

where $F_{\mu\nu}(K) = K_\mu K_\nu - \delta_{\mu\nu} K^2$ is the Lorentz factor.

The temporal moments of the correlation function $C_{\mu\nu}(\vec{k}, t)$ are computed through

$$\begin{aligned} G_{n,\mu\nu}^{\vec{k}} &\equiv \frac{(i)^n}{n!} \int dt t^n C_{\mu\nu}(\vec{k}, t) \\ &= \frac{1}{n!} \frac{\partial^n [\Pi^E(K^2) F_{\mu\nu}(K)]}{\partial (K_t)^n} \Big|_{K_t=0} \\ &= \begin{cases} M_m^{\vec{k}} F_{\mu\nu}(K) \Big|_{K_t=0} + \frac{1}{2} M_{m-1}^{\vec{k}} \frac{\partial^2 F_{\mu\nu}(K)}{\partial K_t^2} \Big|_{K_t=0}, & \text{for } n = 2m, \\ M_m^{\vec{k}} \frac{\partial F_{\mu\nu}(K)}{\partial K_t} \Big|_{K_t=0}, & \text{for } n = 2m + 1, \end{cases} \end{aligned} \quad (17)$$

where m is an integer and the coefficients $M_m^{\vec{k}}$ are given by

$$M_m^{\vec{k}} = \frac{1}{m!} \frac{\partial^m \Pi^E(K^2)}{\partial (K_t^2)^m} \Big|_{K_t=0}. \quad (18)$$

According to the definition in Eq. (17), the temporal moments $G_{n,\mu\nu}^{\vec{k}}$ are nothing but the coefficients in the Taylor expansion of the function $\Pi^E(K^2)F_{\mu\nu}(K)$ at $K_t = 0$. Using a once-subtracted dispersion relation and the optical theorem

$$\begin{aligned} \Pi^E(K^2) - \Pi^E(0) &= -\frac{K^2}{\pi} \int ds \frac{\text{Im}[\Pi(s)]}{s(s+K^2)}, \\ \text{Im}[\Pi(s)] &= \frac{R(s)}{12\pi}, \quad R(s) \equiv \frac{\sigma(e^+e^- \rightarrow \text{hadrons})}{4\pi\alpha(s)^2/(3s)}, \end{aligned} \quad (19)$$

we can relate $M_m^{\vec{k}}$ to the experimental observables $R(s)$,

$$\begin{aligned} M_{m=0}^{\vec{k}} &= \Pi^E(0) - \frac{\vec{k}^2}{12\pi^2} \int ds \frac{R(s)}{s(s+\vec{k}^2)}, \\ M_{m \neq 0}^{\vec{k}} &= (-1)^{m+1} \frac{\vec{k}^2}{12\pi^2} \int ds \frac{R(s)}{s(s+\vec{k}^2)^{m+1}} + (-1)^m \frac{1}{12\pi^2} \int ds \frac{R(s)}{s(s+\vec{k}^2)^m}. \end{aligned} \quad (20)$$

Note that $M_m^{\vec{k}}$ is suppressed by a factor of $(s+\vec{k}^2)^{-m} \leq (M_V^2 + \vec{k}^2)^{-m} = (E_V^2)^{-m}$ with $E_V \equiv \sqrt{M_V^2 + \vec{k}^2}$ again the lowest energy level in the vector channel; see the previous section. We can construct a convergent series through

$$S^{\vec{k}}(\omega^2) = \sum_m M_m^{\vec{k}}(-\omega^2)^m, \quad \text{if } \omega^2 < M_V^2 + \vec{k}^2 \quad (\omega < E_V). \quad (21)$$

Putting Eq. (20) into the series and using $\Pi^E(0) = \Pi^M(0)$, we find that $S^{\vec{k}}(\omega^2)$ satisfies the dispersion relation

$$S^{\vec{k}}(\omega^2) - \Pi^M(0) = \frac{\omega^2 - \vec{k}^2}{12\pi^2} \int ds \frac{R(s)}{s(s - (\omega^2 - \vec{k}^2))}, \quad (22)$$

which means that $S^{\vec{k}}(\omega^2)$ is equivalent to $\Pi^M(k^2)$ at $k^2 = \omega^2 - \vec{k}^2$.

On the other hand, we can construct another series through

$$\begin{aligned} S_{\mu\nu}^{\vec{k}}(\omega) &= \sum_n G_{n,\mu\nu}^{\vec{k}}(-i\omega)^n \\ &= S^{\vec{k}}(\omega^2) \left(F_{\mu\nu}(K) \Big|_{K_t=0} + (-i\omega) \frac{\partial F_{\mu\nu}(K)}{\partial K_t} \Big|_{K_t=0} + \frac{(-i\omega)^2}{2} \frac{\partial^2 F_{\mu\nu}(K)}{\partial (K_t)^2} \Big|_{K_t=0} \right) \\ &= \Pi^M(k^2) \Big|_{k=(\omega, \vec{k})} (K_\mu K_\nu - \delta_{\mu\nu} K^2) \Big|_{K=(\vec{k}, -i\omega)}. \end{aligned} \quad (23)$$

$S_{\mu\nu}^{\vec{k}}(\omega)$ is nothing but $\bar{\Pi}(K^2)(K_\mu K_\nu - \delta_{\mu\nu} K^2)$ as given in Eq. (3). We thus have demonstrated the equivalence between $\bar{\Pi}(K^2)$ and $\Pi^M(k^2)$ using temporal moments of the Euclidean vector correlation function.

In a special case for $\vec{k} = \vec{0}$ and $\mu = \nu = z$, we have

$$G_{0,zz}^{\vec{0}} = 0, \quad G_{2m+1,zz}^{\vec{0}} = 0, \quad G_{2m+2,zz}^{\vec{0}} = \frac{1}{2} M_m^{\vec{0}} \neq 0. \quad (24)$$

The HVP function can then be constructed by

$$\begin{aligned} \bar{\Pi}(-\omega^2) &= -\frac{1}{\omega^2} \sum_m G_{2m+2,zz}^{\vec{0}} (-i\omega)^{2m+2} = -\int dt \frac{e^{\omega t} - 1}{\omega^2} C_{zz}(\vec{0}, t), \\ \bar{\Pi}(-\omega^2) - \bar{\Pi}(0) &= -\int dt \left[\frac{e^{\omega t} - 1}{\omega^2} - \frac{t^2}{2} \right] C_{zz}(\vec{0}, t). \end{aligned} \quad (25)$$

Eq. (25) is the analytic continuation of the formula

$$\Pi(K_t^2) - \Pi(0) = \int dt \left[\frac{e^{iK_t t} - 1}{K_t^2} + \frac{t^2}{2} \right] C_{zz}(\vec{0}, t) \quad (26)$$

given in Ref. [4].

V. COMPUTATION OF $\bar{\Pi}(K^2)$

The analytic continuation method described in the previous sections has been successfully applied in lattice QCD calculations of pion and charmonium radiative decay [17, 18]. In this work we present the first lattice calculation of the HVP function using this technique.

We will use the same ensembles as in Ref. [9]. The gauge configurations are generated using two-flavor maximally twisted mass fermions [19]. The masses of the up and down quarks are equal and heavier than the physical value with the pion mass ranging from 650 MeV to 290 MeV. We employ two lattice spacings $a = 0.079$ fm and 0.063 fm to check for lattice artifacts and two lattice volumes to examine finite-size effects.

On a finite lattice, Eq. (3) takes the form

$$\begin{aligned} \bar{\Pi}(K^2; t_{\max}) (K_\mu K_\nu - \delta_{\mu\nu} K^2) &= \bar{\Pi}_{\mu\nu}(\vec{k}, \omega; t_{\max}), \\ \bar{\Pi}_{\mu\nu}(\vec{k}, \omega; t_{\max}) &= \sum_{t=-t_{\max}}^{t_{\max}-a(\delta_{\mu,t}-\delta_{\nu,t})} e^{\omega(t+a(\delta_{\mu,t}-\delta_{\nu,t})/2)} C_{\mu\nu}(\vec{k}, t), \end{aligned} \quad (27)$$

where $K = (\hat{k}_1, \hat{k}_2, \hat{k}_3, i\hat{\omega})$, with $\hat{k}_i \equiv (2/a) \sin(k_i a/2)$ and $\hat{\omega} \equiv (2/a) \sinh(\omega a/2)$, is the standard lattice definition of four-momentum. The correlator $C_{\mu\nu}(\vec{k}, t)$ is defined by

$$C_{\mu\nu}(\vec{k}, t) = \sum_{\vec{x}} e^{-i\vec{k}(\vec{x}+a\hat{\mu}/2-a\hat{\nu}/2)} \langle J_\mu^E(\vec{x}, t) J_\nu^E(\vec{0}, 0) \rangle, \quad (28)$$

with $J_\mu^E(\vec{x}, t)$ the point-split conserved vector current. The value of $t_{\max}$ can be taken up to $T/2$, where T is the temporal extent of the lattice.

In Eq. (27) we assume that the HVP tensor carries a Lorentz factor $K_\mu K_\nu - \delta_{\mu\nu} K^2$. For some values of ω , we have $K_\mu K_\nu - \delta_{\mu\nu} K^2 = 0$. We denote these special values by ω_0 . At $\omega = \omega_0$ and in the large- $t_{\max}$ limit, the HVP tensor $\bar{\Pi}_{\mu\nu}(\vec{k}, \omega; t_{\max})$ is supposed to be consistent with zero up to Lorentz symmetry breaking effects. In our calculation we do find this to be verified for each spatial momentum $\vec{k}$ and polarization direction $\{\mu, \nu\}$. We therefore calculate $\bar{\Pi}(K^2, t_{\max})$ from

$$\bar{\Pi}(K^2; t_{\max}) = \frac{\bar{\Pi}_{\mu\nu}(\vec{k}, \omega; t_{\max}) - \bar{\Pi}_{\mu\nu}(\vec{k}, \omega_0; t_{\max})}{K_\mu K_\nu - \delta_{\mu\nu} K^2}. \quad (29)$$

A. Classification of the correlators

FIG. 1: Time dependence of the correlators with different polarization directions.

On a lattice with a linear box size L , the spatial momenta $\vec{k}$ in Eq. (28) take discrete values

$$\vec{k} = (2\pi/L)\vec{n}, \quad \text{for } \vec{n} \in \mathbb{Z}^3. \quad (30)$$

In the calculation performed here, we choose the four lowest momentum modes with $\vec{n} = (0, 0, 0)$, $(1, 0, 0)$, $(1, 1, 0)$, and $(1, 1, 1)$ and will use $|n|^2$ to distinguish these different modes.

The rotational symmetry indicates that under a (discrete) rotation $\hat{R} \in SO(3, \mathbb{Z})$ with $SO(3, \mathbb{Z})$ the cubic group, the correlators satisfy

$$C_{ij}(\hat{R}\vec{k}, t) = \Lambda_{ii', jj'}(\hat{R}^{-1}) C_{i'j'}(\vec{k}, t), \quad \text{for } i, i', j, j' = x, y, z, \quad (31)$$

where $\Lambda_{ii', jj'}$ denotes a suitable representation of the group $SO(3, \mathbb{Z})$. In addition to the rotational symmetry, the lattice action is also invariant under parity and time reversal symmetries¹, which yields²

$$C_{\mu\nu}(\vec{k}, t) = \eta_{\mu\nu} C_{\mu\nu}(-\vec{k}, t), \quad C_{\mu\nu}(\vec{k}, t) = \eta_{\mu\nu} C_{\mu\nu}(\vec{k}, -t - a(\delta_{\mu,t} - \delta_{\nu,t})), \quad (32)$$

with a factor $\eta_{\mu\nu}$ defined by

$$\eta_{\mu\nu} = \begin{cases} +1, & \text{for } \mu, \nu = x, y, z, \\ -1, & \text{for } \mu = t, \nu = x, y, z \text{ or the reverse}, \\ +1, & \text{for } \mu = \nu = t. \end{cases} \quad (33)$$

We classify the correlators using the polarization direction $\{\mu, \nu\}$. Furthermore, in each class of correlators we average those which are equivalent under the symmetries of rotation, parity, and time reversal. In Table I we list the classification of the correlators applying the notation of the momentum mode $|n|^2$ and the polarization direction $\{\mu, \nu\}$.

$ n ^2$	0	1				2					3			
$\vec{n}$	(0, 0, 0)	(1, 0, 0)				(1, 1, 0)					(1, 1, 1)			
$\{\mu, \nu\}$	$\{x, x\}$	$\{x, x\}$	$\{x, t\}$	$\{y, y\}$	$\{t, t\}$	$\{x, x\}$	$\{x, y\}$	$\{x, t\}$	$\{z, z\}$	$\{t, t\}$	$\{x, x\}$	$\{x, y\}$	$\{x, t\}$	$\{t, t\}$
N_{sym}	3	6	12	12	6	24	24	48	12	12	24	48	48	8

TABLE I: Classification of the correlators $C_{\mu\nu}(\vec{k}, t)$ with $\vec{k} = (2\pi/L)\vec{n}$. N_{sym} is the number of equivalent correlators under rotation, parity and time reversal.

¹ Note that the twisted mass fermion action violates parity symmetry. In our work we formed a parity average to enforce $O(a)$ improvement.

² Using the local vector current, we will have the relation $C_{\mu\nu}(\vec{k}, t) = \eta_{\mu\nu} C_{\mu\nu}(-\vec{k}, t)$ and $C_{\mu\nu}(\vec{k}, t) = \eta_{\mu\nu} C_{\mu\nu}(\vec{k}, -t)$.

For a given momentum mode $|n|^2$, different polarizations of the vector current can lead to different types of correlators. Taking $|n|^2 = 1$ as an example, we show the time dependence of the correlators $C_{\mu\nu}(\vec{k}, t)$ in Fig. 1. Notice that, although the correlators themselves are different, they lead to the same HVP function, as we will demonstrate below.

B. Finite-size effect

While Eq. (3) requires us to compute an integral in the range from $t = -\infty$ to $t = +\infty$, with a given finite lattice volume we can only perform a summation over t values from $t = -T/2$ to $+T/2$ as shown in Eq. (27). Thus, our calculations are contaminated by finite-size effects, which, however, will vanish in the limit $T \rightarrow \infty$.

Ideally, it would be desirable to use the complete t range on the lattice. However, in practice the correlator $C_{\mu\nu}(\vec{k}, t)$ at $|t|$ close to $T/2$ shows very large fluctuations and often no useful information can be extracted for these large values of $|t|$. Therefore, we define a maximal t value, $t_{\max} = \eta(T/2)$. To be concrete, we will set $\eta = 3/4$ in the following. Note that, in principle, any value of η would provide a well-defined choice for our method.

Of course, on a finite-size lattice the above value of η will induce a finite-size effect. This systematic effect is given by

$$\begin{aligned} \bar{\Pi}(K^2; t > t_{\max})(K_\mu K_\nu - \delta_{\mu\nu} K^2) &= \bar{\Pi}_{\mu\nu}(\vec{k}, \omega; t > t_{\max}) \\ \bar{\Pi}_{\mu\nu}(\vec{k}, \omega; t > t_{\max}) &\equiv \left(\sum_{t=t_{\max}+a-a(\delta_{\mu,t}-\delta_{\nu,t})}^{+\infty} + \sum_{t=-\infty}^{-t_{\max}-a} \right) e^{\omega(t+a(\delta_{\mu,t}-\delta_{\nu,t})/2)} C_{\mu\nu}(\vec{k}, t). \end{aligned} \quad (34)$$

In order to obtain an estimate of this finite-size effect, we will assume that for $t > t_{\max}$, the vector correlator is dominated by the ground state. We believe that this provides a good estimate of the finite-size effects in our calculation for the following reasons: First, for all our ensembles, the contribution given in Eq. (34) is already exponentially suppressed and thus contributes only little to the total vacuum polarization function. Second, even if for $t > t_{\max}$ other states may contribute, they provide only a correction to a correction and thus should not change our conclusions significantly. Note that the situation might change if one uses large ω and makes K^2 approach the hadron production threshold. In this case, the $t > t_{\max}$ contribution becomes dominant. Besides this, both the energy and the amplitude extracted from $C_{\mu\nu}(\vec{k}, t)$ are affected by the finite lattice volume. Such effects should be

treated properly using the Lellouch–Lüscher method [20] and Meyer’s proposal in Ref [21], which is, however, beyond the scope of this work.

In this calculation we evaluate the HVP function only in the region of spacelike or low timelike momenta. Our strategy will be to use as our results the values extracted from $\bar{\Pi}(K^2; t_{\max})$ and then estimate the finite-size effects by computing the contribution of $\bar{\Pi}(K^2; t > t_{\max})$. It turns out that with our current lattice setup, the finite-size effects are comparable to the statistical error and thus cannot be neglected in our calculation. Of course, the calculation can be systematically improved in future work by using larger volumes and higher statistics.

C. Results for $\bar{\Pi}(K^2)$

FIG. 2: $\bar{\Pi}(K^2; t_{\max})$ as a function of K^2 . The ensemble used here features $a = 0.079$ fm, $L^3 \times T/a^4 = 24^3 \times 48$ and $m_\pi = 423$ MeV.

We calculate $\bar{\Pi}(K^2; t_{\max})$ using Eqs. (27) - (29) and show $\bar{\Pi}(K^2; t_{\max})$ as a function of K^2 for different momentum modes and polarization directions in Fig. 2. We demonstrate here that using the analytic continuation method, we are able to calculate $\bar{\Pi}(K^2)$ at continuous

FIG. 3: $\bar{\Pi}(K^2; t_{\max}) + \bar{\Pi}(K^2; t > t_{\max})$ as a function of K^2 . We use the same ensemble as in Fig. 2.

K^2 , which cover both the spacelike and timelike momentum domain. For some momentum modes, e.g., $|n|^2 = 1$, small discrepancies appear in the HVP functions for different polarization directions, which we attribute to finite-size effects. As a next step we evaluate these finite-size effects using Eq. (34). The results for $\bar{\Pi}(K^2; t_{\max}) + \bar{\Pi}(K^2; t > t_{\max})$ are shown in Fig. 3. After adding the contribution of $\bar{\Pi}(K^2; t > t_{\max})$ the results for different polarization directions turn out to be more consistent. This finding suggests that the discrepancies for $\bar{\Pi}(K^2; t_{\max})$ do originate from finite-size effects, which then constitute the dominant systematic effect in our calculation.

Since the HVP functions are consistent among various momentum modes and polarization directions, we can average them to obtain the final result. Here we perform a weighted average with a weight of $1/\sigma_{\text{stat}}^2$, where σ_{stat} is the relative statistical error of $\bar{\Pi}(K^2; t_{\max})$ (or $\bar{\Pi}(K^2; t_{\max}) + \bar{\Pi}(K^2; t > t_{\max})$). Particularly at $K^2 = 0$ we show in the upper panel of Fig. 4 the averaged result for $\bar{\Pi}(0; t_{\max}) = -0.2047(20)$ and in the lower panel the one for $\bar{\Pi}(0; t_{\max}) + \bar{\Pi}(0; t > t_{\max}) = -0.2069(21)$. These results deviate at the 1σ level, demonstrating that finite-size effects are comparable to the statistical error.

Up to now, the analysis is performed for the case of $N_f = 2$ flavors. To allow for a direct

FIG. 4: A comparison of $\bar{\Pi}(K^2; t_{\max})$ at $K^2 = 0$ for different momentum modes $|n|^2$ and polarization directions $\{\mu, \nu\}$. In the upper panel, $\bar{\Pi}(0; t_{\max})$ is calculated directly using Eq. (27), while in the lower panel, $\bar{\Pi}(0; t_{\max})$ is corrected using Eq. (34). We use the same ensemble as in Fig. 2. Averaging the results for different momentum modes and polarization directions, we obtain $\bar{\Pi}(0; t_{\max}) = -0.2047(20)$ and $\bar{\Pi}(0; t_{\max}) + \bar{\Pi}(0; t > t_{\max}) = -0.2069(21)$.

comparison with experimental data, we extend the currently used method to the case of $N_f = 2 + 1 + 1$ flavors [22]. We add the $t > t_{\max}$ contribution to the renormalized HVP function $\Pi_R(K^2) = \bar{\Pi}(K^2) - \bar{\Pi}(0)$ and extrapolate it to the physical pion mass using the modified extrapolation method proposed in Refs. [3, 9]. In the timelike region, especially the region where K^2 approaches the hadron production threshold, it is very difficult to reproduce $\Pi_R(K^2)$ due to the significant finite-size effects. We therefore restrict the calculation of $\Pi_R(K^2)$ to the spacelike region. The experimental results for $\Pi_R(K^2)$ are compiled using Jegerlehner's package **alphaQED** [23], where the dispersion relation is used to relate the experimental data of $R(s)$ (last updated in 2012) to $\Pi_R(K^2)$. As illustrated in Fig. 5, the lattice results for $\Pi_R(K^2)$ are consistent with the experimental data but with presently available statistics the corresponding fluctuations are much larger. Nevertheless, the found agreement is reassuring that the analytic continuation method can describe the vacuum

polarization function in the low-momentum region.

FIG. 5: A comparison between $N_f = 2 + 1 + 1$ -flavor lattice results for $\Pi_R(K^2)$ calculated using the analytic continuation method proposed in this work and the experimental results compiled using the cross sections $R(s)$ as input together with the dispersion relation [23]. The lattice results are shown by the red sparse error bars and the experimental data are shown by the grey condensed error band. We find that the lattice results are consistent with the experimental data but carry larger errors.

VI. DETERMINATION OF a_μ^{hvp}

The lattice result of the HVP function obtained in the previous section can be used to determine a series of physical observables; see, e.g., Ref.[3]. Here we take the leading-order HVP correction to the muon anomalous magnetic moment, a_μ^{hvp} , as an example to study the practical feasibility of the proposed analytic continuation method.

In lattice QCD a_μ^{hvp} can be calculated through

$$a_\mu^{\text{hvp}} = \alpha^2 \int_0^\infty dK^2 \frac{1}{K^2} f\left(\frac{K^2}{m_\mu^2}\right) (\Pi(K^2) - \Pi(0)) , \quad (35)$$

where α is the fine structure constant, m_μ is the muon mass, and $f(K^2/m_\mu^2)$ is a known function [6], which assumes a maximum at $K^2 = (\sqrt{5} - 2)m_\mu^2 \approx 0.003 \text{ GeV}^2$. To control the

chiral extrapolation, we use a modified definition proposed in Refs. [3, 9],

$$a_{\bar{\mu}}^{\text{hvp}} = \alpha^2 \int_0^\infty dK^2 \frac{1}{K^2} f\left(\frac{K^2}{m_\mu^2} \frac{H_{\text{phys}}^2}{H^2}\right) (\Pi(K^2) - \Pi(0)) , \quad (36)$$

with $H = M_V$. Note that when the pion mass approaches its physical value, we have $H = H_{\text{phys}}$. Thus, the modified definition, $a_{\bar{\mu}}^{\text{hvp}}$, reproduces the value of $a_{\bar{\mu}}^{\text{hvp}}$ at the physical pion mass.

In conventional lattice calculations, to perform the integral in Eq. (36), one needs to parametrize the HVP function. Using the analytic continuation method, we can calculate the HVP function for a continuous momentum region $0 < K^2 < K_{\text{max}}^2$, with $K_{\text{max}}^2 = \sum_{i=x,y,z} \hat{K}_i^2$ being the squared spatial momentum. Thus, we can avoid the parametrization in this momentum region. In practice we split Eq. (36) into three parts:

$$\begin{aligned} a_{\bar{\mu}}^{\text{hvp}} &= a_{\bar{\mu}}^{(1)} + a_{\bar{\mu}}^{(2)} + a_{\bar{\mu}}^{(3)} , \\ a_{\bar{\mu}}^{(1)} &= \alpha^2 \int_0^{K_{\text{max}}^2} dK^2 \frac{1}{K^2} f\left(\frac{K^2}{m_\mu^2} \frac{H_{\text{phys}}^2}{H^2}\right) (\Pi(K^2) - \Pi(0)) , \\ a_{\bar{\mu}}^{(2)} &= \alpha^2 \int_{K_{\text{max}}^2}^\infty dK^2 \frac{1}{K^2} f\left(\frac{K^2}{m_\mu^2} \frac{H_{\text{phys}}^2}{H^2}\right) (\Pi(K_{\text{max}}^2) - \Pi(0)) , \\ a_{\bar{\mu}}^{(3)} &= \alpha^2 \int_{K_{\text{max}}^2}^\infty dK^2 \frac{1}{K^2} f\left(\frac{K^2}{m_\mu^2} \frac{H_{\text{phys}}^2}{H^2}\right) (\Pi(K^2) - \Pi(K_{\text{max}}^2)) . \end{aligned} \quad (37)$$

In Eq. (37), $a_{\bar{\mu}}^{(1)} + a_{\bar{\mu}}^{(2)}$ can be calculated directly using the analytic continuation method. As shown by Eq. (27), $\bar{\Pi}(K^2; t_{\text{max}})$ is a linear combination of the correlator $C_{\mu\nu}(\vec{k}, t)$. Putting this definition of $\bar{\Pi}(K^2; t_{\text{max}})$ into Eq. (37), it turns out that $a_{\bar{\mu}}^{(1)}$ and $a_{\bar{\mu}}^{(2)}$ are also linear combinations of $C_{\mu\nu}(\vec{k}, t)$ with coefficients that can be determined by performing the integral in Eq. (37). Similar to the previous section, we can calculate $a_{\bar{\mu}}^{(1)}$ and $a_{\bar{\mu}}^{(2)}$ up to the value of $t_{\text{max}} = \eta(T/2)$ with $\eta = 3/4$ and estimate the finite-size effects using $\bar{\Pi}(K^2; t > t_{\text{max}})$.

The evaluation of $a_{\bar{\mu}}^{(3)}$ still requires a parametrization of $\bar{\Pi}(K^2)$. This will bring in some model dependence in our analysis, which, however, is a small effect, since the total contribution of $a_{\bar{\mu}}^{(3)}$ only amounts to a few percent in case of momentum modes $|n|^2 = 1, 2, 3$. The parametrization of $\bar{\Pi}(K^2)$ used in this calculation is given in the Appendix.

In Fig. 6 we show $a_{\bar{\mu}}^{\text{hvp}}$ as a function of the squared pion mass. The results of $a_{\bar{\mu}}^{\text{hvp}}$ calculated using Eq. (37) are shown by the empty symbols. These results have been averaged among various momentum modes ($|n|^2 = 1, 2, 3$) and polarization directions. For comparison the results determined using our conventional approach of parametrizing the HVP function in

the full momentum range are shown in the same figure, represented by the filled symbols. In the upper panel we show the results of $a_{\bar{\mu}}^{\text{hvp}}(t_{\text{max}})$, which are calculated using the correlator $C_{\mu\nu}(\vec{k}, t)$ covering the range of $-t_{\text{max}} \leq t \leq t_{\text{max}}$. We look at the finite-size effects in $a_{\bar{\mu}}^{\text{hvp}}(t_{\text{max}})$ by comparing the results for different volumes. There are also some deviations between the results from the analytic continuation method and the standard parametrization method. To check for the finite-size effects, we evaluate the contribution to $a_{\bar{\mu}}^{\text{hvp}}$ from $C_{\mu\nu}(\vec{k}, t)$ at $|t| > t_{\text{max}}$ leading to a correction $a_{\bar{\mu}}^{\text{hvp}}(t > t_{\text{max}})$. The corresponding results for $a_{\bar{\mu}}^{\text{hvp}}(t_{\text{max}}) + a_{\bar{\mu}}^{\text{hvp}}(t > t_{\text{max}})$ are shown in the lower panel of Fig. 6. These results are consistent now among different lattice volumes. Besides this, the results from the analytic continuation method also agree with the ones from the standard parametrization method. As can be seen, the results from the analytic continuation method for $a_{\bar{\mu}}^{\text{hvp}}$ show larger fluctuations than the standard ones. However, the analytic continuation method has the conceptual advantage that in the region of low K^2 the parametrization of $\bar{\Pi}(K^2)$ can be avoided. Thus, we think that presently the analytic continuation method can serve as a valuable cross-check of the standard method to analyze the vacuum polarization function.

VII. CONCLUSIONS

The analytic continuation method, proposed in this work, allows us to obtain information on the momentum dependence of the vacuum polarization function at small momenta, which is a significant challenge for lattice QCD calculations. In order to see how the analytic continuation method works in practice, we performed a pilot study for the HVP function to determine the leading-order QCD correction to the muon anomalous magnetic moment. Since at large Euclidean times the HVP is very noisy, we restricted the time summation to a maximum time, $t_{\text{max}} = \eta T/2$, with T the time extent of our lattice. In this work, we have chosen $\eta = 3/4$. Although in the infinite volume limit this would lead to a fully correct definition for $a_{\bar{\mu}}^{\text{hvp}}$, on a finite lattice such a choice induces a finite-size effect. We estimated this finite-size effect by assuming that in the time region excluded by the cut $\eta = 3/4$ the ground state dominates in the vector channel.

In the case of $N_f = 2+1+1$ flavors, adding the so-computed finite-size effects to the results for the HVP function provides a result that agrees with the experimental determination of Ref. [23].

FIG. 6: a_μ^{hvp} as a function of the squared pion mass. For each ensemble, the empty symbols stand for the results calculated using the analytic continuation method. We have averaged a_μ^{hvp} for various momentum modes ($|n|^2 = 1, 2, 3$) and polarization directions. In the upper panel, we show $a_\mu^{\text{hvp}}(t_{\text{max}})$, which are calculated using the correlator $C_{\mu\nu}(\vec{k}, t)$ covering the range $-t_{\text{max}} \leq t \leq t_{\text{max}}$. In the lower panel, we show the total contribution of $a_\mu^{\text{hvp}}(t_{\text{max}}) + a_\mu^{\text{hvp}}(t > t_{\text{max}})$, where an estimate of the finite-size effects has been added. The filled symbols indicate the results for a_μ^{hvp} from the conventional parametrization method.

Going to the case of $N_f = 2$ flavors, we find after the finite-size correction, consistent results for a_μ^{hvp} for all ensembles and also an agreement with results computed earlier by us using the conventional approach. We therefore conclude that applying a cut in the Euclidean time induces indeed a finite-size effect as the dominant systematic error. Thus, when the analytic continuation method is applied on larger lattices in the future, any assumption such as the here employed ground state dominance at large times can be completely avoided, leading to a conceptually clear determination of quantities derived from the HVP function. In this way it would not be required anymore to rely on model dependent parametrizations that enter some of the conventional computations of a_μ^{hvp} . On the negative side, it needs

to be said that, although having a conceptual advantage, the analytic continuation method gives results, at least for a_μ^{hvp} , which have fluctuations that are larger than the conventional method.

However, since all the methods are afflicted with different systematic uncertainties, comparing results from different approaches will provide confidence in the extraction of physical quantities where the HVP function is an essential ingredient. In this sense we believe that the analytic continuation procedure presented in this paper can provide a valuable alternative to other approaches to determine the HVP function. It will be interesting to test the potential of the here proposed analytic continuation method for quantities different from the HVP function, where the momentum dependence at small or even zero momentum is not directly accessible.

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Appendix A: Parametrization of the HVP function

In the conventional approach, in order to compute the integral in Eq. (36), we need a functional form to describe the K^2 dependence of $\Pi(K^2)$. For low K^2 , we use the form

$$\Pi_{\text{low}}(K^2) = \frac{-g_V^2 M_V^2}{K^2 + M_V^2} + (a_0 + a_1 K^2) \ , \quad (\text{A1})$$

where the first term is the dominant contribution from the ground-state vector meson and g_V is the electromagnetic coupling of the vector meson. The polynomial terms account for

residual contributions. $\Pi(0)$ is given by $-g_V^2 + a_0$. For high K^2 , we use the form

$$\Pi_{\text{high}}(K^2) = c + \ln(K^2) \left(\sum_{i=0}^3 b_i (K^2)^i \right), \quad (\text{A2})$$

with a free constant c and a sum over four terms in the polynomial multiplying the logarithmic function. We combine the two expressions using

$$\Pi(K^2) = \frac{1-t}{2} \Pi_{\text{low}}(K^2) + \frac{1+t}{2} \Pi_{\text{high}}(K^2), \quad (\text{A3})$$

where $t = \tanh((K^2 - K_m^2)/\Delta_m^2)$ is a smooth approximation to the step function. The parameters are set as $K_m = 1.3M_V$ and $\Delta_m = 0.3M_V$. We fit the lattice data of $\Pi(K^2)$ to Eq. (A3) with seven free parameters: a_0 , a_1 , c , b_0 , b_1 , b_2 and b_3 .

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