Photoionization of metastable heliumlike $\text{C}^{4+}(1s2s\,^3S_1)$ ions:
Precision study of intermediate doubly excited states

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In a joint experimental and theoretical endeavor, photoionization of metastable $\text{C}^{4+}(1s2s\,^3S_1)$ ions via intermediate levels with hollow, double-$K$-vacancy configurations $2s2p$, $2s3p$, $2p3s$, $2p3d$, $2p4s$, and $2p4d$ has been investigated. High-resolution photon-ion merged-beams measurements were carried out with the resolving power reaching up to 25 800 which is sufficient to separate the leading fine-structure components of the $2s2p\,^3P$ term. Many-body perturbation theory was employed to determine level-to-level cross sections for $K$-shell excitation with subsequent autoionization. The resonance energies were calculated with inclusion of electron correlation and radiative contributions. Their uncertainties are estimated to be below ±1 meV. Detailed balance confirms the present photoionization cross-section results by comparison with previous dielectronic-recombination measurements. The high accuracy of the theoretical transition energies together with the present experimental results qualify photoabsorption resonances in heliumlike ions as new, greatly improved energy-reference standards at synchrotron radiation facilities.

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I. INTRODUCTION

Structure and dynamics of the helium atom and heliumlike ions are among the most interesting subjects in fundamental atomic physics research. Two-electron atoms and ions are the simplest many-electron systems and thus, next to the hydrogen atom and the associated isoelectronic sequence, provide the best possible framework for theory to describe their physical properties with high accuracy. Yet, their three-particle nature and the competition between electron-electron and electron-nucleus interactions make them sufficiently complex to challenge detailed theoretical treatments [1].

From an application-oriented point of view, two-electron ions provide an important tool for plasma diagnostics. Plasma parameters such as electron density and temperature can be derived from line-intensity ratios in the radiation emitted when heliumlike ions are present ([2–5] and references therein). This is especially important for astrophysics and fusion research.

Doubly excited states of helium, i.e., atoms with an empty $K$ shell, are particularly intriguing. Such states were discovered by Compton and Boyce [6] by spectral analysis of the radiation emitted from a low-pressure helium discharge tube and then observed again by Krüger [7] using a hollow-cathode discharge lamp. Both experiments found lines at wavelengths smaller than what is required for ionizing a helium atom and the authors discussed their assignment to $2\ell 2\ell' \rightarrow 1s2\ell$ transitions with $\ell, \ell' = s, p$. About thirty years later Madden and Codling [8] measured the absorption spectrum of helium by using synchrotron radiation and found unambiguous evidence for the population of doubly excited autoionizing levels. Their experiment initiated widespread and long-lasting interest in multiply excited states of atoms. Obviously, the doubly excited levels in the helium atom had been populated.
by absorption of a single photon,
\[ \gamma + \text{He}(1s^2) \rightarrow \text{He}(n\ell n'\ell') \]  
(1)
i.e., both K-shell electrons had been excited simultaneously
forming \( n, n' \geq 2 \) \( ^1P \) levels with energies sufficient to sub-
sequently eject an electron via Auger decay. Later, even the produc-
tion of doubly excited \( ^3P \) levels by spin-forbidden absorption of a single photon by \( \text{He}(1s^21S) \) was observed ([9,10] and references therein).

There are several different spectroscopic notations for hel-
iumlike resonant states. A review of resonance parameters and
notations has been presented by Rost et al. [11]. In the
present publication the classification of doubly excited heli-
iumlike levels is that of the most comprehensive collection of
spectroscopic information on atoms in various charge states,
the NIST Atomic Spectra Database [12].

Photoexcitation producing doubly excited levels in helium
atoms has been investigated in great detail [13–16] by em-
ploying the increased photon flux available at second- and
third-generation synchrotron light sources. Obtaining similar
results for heliumlike atomic ions is much more demanding.
The reason for this is the decrease of the pertinent cross
sections with increasing atomic number and the difficulty
to provide sufficiently high ion-target densities [17]. In the
sections with increasing atomic number and the difficulty
of intermediate configurations \( n\ell n'\ell' = 2s2p, 2s3p, 2p3s, 2p3d, 2s4p, 2p4s, \) and \( 2p4d \). Similar experiments have only
been reported for the neutral helium atom [36], but at a much
lower resolving power compared to the present experiment
and without providing absolute cross sections.

While the present work is part of the research effort dealing
with the physics of heliumlike atoms and ions it is also part
of an experimental endeavor to study K-shell photoexcita-
tion and photoionization of carbon ions in different charge
states. Investigations have been carried out previously on \( C^+ \) [37–39], \( C^2+ \) [40], and \( C^{3+} \) [41]. Photoionization of atomic
ions via the production of a K-shell vacancy and the related
aspects of precision measurements of the atomic structure
of autoionizing states have been discussed in a recent re-
view [42]. Since that review, new experimental investigations
on photoionization of atomic ions involving K-shell electrons
have been published for \( O^+, O^{2+} \) [43], \( O^{4+}, O^{5+} \) [44],
\( Ne^+ \) [45], \( O^- \) [46], \( F^- \) [47], \( Fe^{2+}, Fe^{3+}, Fe^{22+}, \) and \( Fe^{23+} \) [48].

This paper is organized as follows. After this introduction,
Sec. II provides a brief description of the experimental tech-
nique and the measurement of absolute photoionization cross
sections. The energy calibration and related uncertainties are
discussed in detail. The calculations performed in the present
context by employing many-body perturbation theory with
complex rotation with inclusion of contributions to the level
energies arising from the quantization of the electromagnetic
field are described in Sec. III. In the main section, Sec. IV,
the experimental and theoretical results are presented. They are
compared with one another and with information available in
the literature. In particular, the principle of detailed balance is
exploited to demonstrate the consistency of the present results
on photoionization of \( C^{4+}(1s2s3S) \) with previous measure-
ments of absolute cross sections for dielectronic recombi-
nation of \( C^{5+}(1s2S) \) ions. The paper ends with a summary and
an outlook.
II. EXPERIMENT

The measurements were carried out using the PIPE (Photon-Ion Spectrometer at PETRA III) end station of beamline P04 [49] at one of the world’s brightest synchrotron radiation sources, PETRA III, at DESY in Hamburg. The experimental setup and procedures have been described previously [45,50]. Photoionization of singly charged carbon ions employing PIPE has been reported [38,39]. Therefore, the description of experimental details is restricted to issues directly related to the present C4+ photoionization measurements.

A. Experimental arrangement and procedures

Heliumlike C4+ ions were produced from methane gas in the hot plasma of an electron-cyclotron-resonance (ECR) ion source. The plasma chamber was set to a potential of +6 kV and the ions were extracted towards ground potential. With electrostatic lenses and steerers the ion beam was transported to the entrance aperture of a double-focusing 90° dipole bending magnet which separated the beam components according to their mass-over-charge (m/q) ratio. Behind the exit aperture of the magnet, the selected 12C4+ ion beam component was transported further, again with electrostatic elements including a quadrupole triplet, parallel steering plates, and a hemispherical 50° deflector, the merger, which steered the ion beam onto the photon-beam axis and focused it to the center of a 50 cm long drift tube, the photon-ion merged-beam interaction region. With voltages on the order of 500 V on the drift tube, ions produced inside are energy-labeled [45,50] and thus, the length of the interaction region is defined. This is necessary for the measurement of absolute cross sections for photon-ion interactions.

The C5+ photoionization products were separated from the C4+ parent ion beam with a second double-focusing 90° dipole bending magnet, the demerger. The parent ion beam was collected in a large Faraday cup located inside the magnet chamber while the product ions passed through apertures and were directed to a 180° out-of-plane hemispherical deflector which focused the product ion beam onto a single-particle detector. The out-of-plane deflection serves to suppress detector background arising from stray ions, electrons, or photons. It cannot discriminate, though, against C5+ ions produced in collisions with residual gas along their path between the merger and the demerger. Suppression of such background is only possible by keeping the vacuum pressure in that section as low as possible. In the interaction region a pressure of 3 x 10^-10 mbar could be maintained.

A further source of background is the occurrence of dark counts. The detector employed in the PIPE experiments is based on a well-optimized geometry [51,52] with a metallic ion-electron converter plate and a channel electron multiplier (CEM) amplifying the few-electron pulses released from the converter plate by impacting ions. All detector components are shielded inside a grounded metal box. The detection efficiency for atomic ions is close to 100% (0.97 ± 0.03). By avoiding sharp edges of electrodes inside the box and by keeping the applied voltages as low as possible (at most ±1200 V) the rate of dark counts has been reduced to the level of 0.02 s^-1.

The 12C4+ ions have an m/q ratio of approximately 3. Beside the desired carbon ions, the ion source also produced H+ with m/q also close to 3. Closer inspection yields m/q ≈ 3.023 u/e for H+ where u is the atomic mass unit and e the elementary charge. For 12C4+ m/q ≈ 2.999 u/e. The 8 per mille difference is by far enough to separate the different ion species by the first dipole magnet.

Magnetic separation of 12C4+ (1s2 3S0) ground-state ions from metastable 12C4+ (1s2 3S1) is not possible. From previous experiments on electron-impact ionization of helium-like ions it is known that the ECR ion source produces 3S metastable ions while the population of 3S ions is negligibly small [34,35]. This can be rationalized by the different lifetimes τ (20.59 ms [53] for C4+(1s2 3S1) and only 3.03 μs [54] for C4+(1s2 3P1)) and the different statistical weights of the 3S and 3P levels. Most important in this context is the approximate flight time of about 15 μs of 12C4+ ions from the source to the interaction region leaving little room for the survival of C4+ (1s2 3S0) excited levels. Hence, the parent C4+ beam consisted of ions in the 1s2 3S ground level with a fraction 1 − f and the 1s2 3S metastable level with a fraction f. From previous experiments in which the same type of ion source was employed, it is known that f was 6% for a beam of N5+ ions [35] and 13.6% for a beam of Li+ ions [34]. The fraction of C4+(1s2 3P1) had to be expected to be somewhere between these two numbers. The exact value is to be determined by comparison with theory and other experiments (see below).

For the measurement of absolute cross sections the ion beam was strongly collimated by variable apertures in front of and behind the interaction region. With these apertures closed to about 1 mm × 1 mm and almost touching the photon beam, the ion beam was optimized for transmission and, by that, very good overlap with the photon beam was enforced. Then the aperture behind the interaction region was opened to a size 3 mm × 3 mm and the beam overlap factor was measured with six independent slit scanners yielding values of around 4400 cm^-1. The ion current in the absolute measurements was about 1 nA. At 359 eV the photon flux was 4.2 x 10^{11} s^-1 at a bandwidth of 16 meV. Under these conditions the background count rate of C5+ ions produced by electron-stripping collisions with residual gas components was about 2 s^-1. The maximum signal count rate on the 2s2p 3P2 resonance was slightly below 4 s^-1. Unfortunately, the dominant (∼90% ground-level) fraction of the parent beam contributed to the electron-stripping background but not to the 3P resonance signals which can exclusively be reached from the 3S metastable component of the parent beam. As it turned out, only about 10% of the total C4+ beam contributed to the measured signal.

Without considering the uncertainty arising from the determination of the metastable-ion fraction f, the apparent cross sections measured with a mixed beam of ground-state and metastable C4+ ions have a systematic uncertainty of 15% [50]. The apparent cross sections were normalized to 100% parent metastable ions by multiplication with f^-1. The fraction f was determined by comparison with the results of the present theory. By employing the principle of detailed balance photoionization cross sections can be converted to
photorecombination data for dielectric recombination of C$^\text{3+}(1s)$ ions, obtained at a heavy-ion storage ring, shows excellent agreement (within 3%; see below) thus confirming the present determination of $f$.

B. Energy calibration and related uncertainties

The photon energy was calibrated against the position of the $2p \rightarrow 4s$ excitation resonance in neutral Ar at 244.39 eV and the position of the lowest vibrational level ($\nu = 0$) reached in the N $1s \rightarrow \pi^+$ excitation of the neutral N$_2$ molecule at 400.88 eV. The reference energies originate from electron-energy-loss spectroscopy (EELS) experiments [55–60] for which low uncertainties in the energy determination are quoted. However, one has to keep in mind that quoted uncertainties tend to be optimistic. In a previous detailed study of characteristic energies in neutral Ne and Ne$^+$ [45] numerous experimental results were compared with one another. Discrepancies between different experiments much larger than the quoted uncertainties were found which casts doubt on the low error bars quoted in the literature. In the case of the Ne $1s \rightarrow 3p$ transition energy at 867.29 eV a realistic uncertainty is presently 0.2 eV in contrast to quoted error bars as low as 20 meV.

Reference energies relevant to the present calibration are (1) for the Ar ($2p \rightarrow 4s$) transition:

- $244.39 \pm 0.01$ eV [55],
- $244.37 \pm 0.02$ eV [58],
- $244.390 \pm 0.004$ eV [60];

(2) for the N$_2$ ($N\ 1s \rightarrow \pi^+, \nu = 0$) transition:

- $400.86 \pm 0.03$ eV [56],
- $400.70 \pm 0.05$ eV [57],
- $404.88 \pm 0.02$ eV [58],
- $404.865 \pm 0.02$ eV [59] [relative to the Ar ($2p \rightarrow 4s$) transition at 244.39 eV].

The energy determination in EELS experiments is directly associated with the measurement of the voltage on the electron spectrometer. Thus the accuracy of transition energies from EELS experiments critically depends on the sensitivity and accuracy of the voltmeter used. The EELS experiments on the Ar calibration line agree with one another within their quoted uncertainties. This is not the case for the N$_2$ reference. Discrepancies have been explained by the different accuracies of digital voltmeters used in the different experiments. However, deficiencies in the performance of the voltmeter should be reflected in the quoted uncertainty of a measured transition energy. Obviously, this was not the case in the experiments by Hitchcock and Brion [57] and Sodhi and Brion [58] whose numbers differ by 180 meV while uncertainties of only 50 meV and 20 meV have been quoted. Thus, it cannot be excluded that the error bar on the 400.88 eV N$_2$ ($N\ 1s \rightarrow \pi^+, \nu = 0$) transition energy quoted by Sodhi and Brion [58] is greater than 20 meV.

The situation calls for improved calibration standards for synchrotron-radiation experiments in the photon-energy range from approximately 300 eV to 1 keV. The uncertainties of the existing standards are so large, that state-of-the-art atomic structure theory cannot at all be tested for light few-electron atomic systems by using synchrotron radiation in that energy range. A possible way to deal with this problem is to calibrate experiments by comparing measured resonance positions with accurate theoretical energies. In the energy range of interest, moderately highly charged ions of light elements have to be considered. These are not typically available at synchrotron radiation sources. However, by using the PIPE setup it would be possible to transfer high-quality calibration standards to the gas-phase standards that are presently in use at synchrotrons worldwide.

In a recent publication, level energies for $1s^22\ell$ and $1s2f/2\ell'$ states in lithiumlike ions with atomic numbers $Z$ between 6 and 17 have been presented [61,62] with uncertainties of a few meV, much smaller than the typical calibration uncertainties of common neutral-gas standards. Transitions in lithiumlike C$^\text{3+}$ through Ne$^\text{7+}$ would adequately cover the energy range of interest with resonances at about 300 eV to more than 900 eV. These ions can be produced with the ECR source presently installed at PIPE so that a recalibration of gas standards is now possible. Also the present investigation with heliumlike C$^\text{4+}$ ions has the potential for providing new calibration standards.

For comparison of the present experimental data with the present theoretical calculations it is desirable to have a theory-independent calibration. Therefore, the “conventional” calibration procedure is discussed further. Based on past experience it was assumed that the deviation between nominal energies provided by the beamline and the true photon energy is a linear function of the nominal photon energy. In order to determine this linear function the Ar and N$_2$ resonance energies mentioned above were employed. With the uncertainties of the reference energies, the linear function has uncertainties which increase when it is extrapolated to energies beyond 400 eV. By assuming an increased uncertainty of 30 instead of 20 meV for the $N\ 1s \rightarrow \pi^+, \nu = 0$ transition found at the uncertainty-weighted average energy of 400.86 eV that results from the existing measurements [56–59], the possible error of the energy axis set by the calibration standards is 40 meV at 440 eV.

There are additional sources of uncertainty in the present calibration for the measurements with C$^\text{4+}$ ions. Since the ions move against the direction of the photon beam the resulting Doppler shift has to be corrected for. The Doppler-corrected energy $E_D$ for interacting ion and photon beams determined from the photon energy $E_{\text{Lab}}$ in the laboratory frame is

$$E_D = \frac{E_{\text{Lab}}}{\gamma(1 + \beta \cos \theta)},$$

with the Lorentz factor

$$\gamma = \frac{1}{\sqrt{1 - \beta^2}},$$

and the angle $\theta$ between the two beam directions. The ion velocity $v = \beta c$ where $c$ is the vacuum speed of light can be inferred from $\gamma$ which is related to the kinetic energy of the ions,

$$E_{\text{kin}} = qeU_{\text{acc}} = (\gamma - 1)m\beta^2 c^2.$$
of the ion. Thus, \( \beta \) can be calculated as

\[
\beta = \sqrt{1 - \frac{1}{(1+x)^2}},
\]

with \( x \) defined as \( qe U_{\text{acc}}/(m_0 c^2) \). The Doppler-corrected photon energy is then given by

\[
E_D = \frac{E_{\text{Lab}}}{(1+x) + \sqrt{2x + x^2 \cos \theta}}.
\]

For counterpropagating beams the angle is \( \theta = 180^\circ \) and \( \cos \theta = -1 \). This results in

\[
E_D = [(1+ x) + \sqrt{2 + x}] E_{\text{Lab}}.
\]

Uncertainty arises from the unknown plasma potential in the ion source which changes the effective acceleration voltage. The plasma potential in an ECR ion source is expected to be no more than 50 V [63]. Thus, the uncertainty of the Doppler correction is at most about 3 meV at \( E_{\text{Lab}} = 360 \) eV in the present case. Assuming that the potential of the ion source plasma chamber was measured with an uncertainty smaller than 0.5\% an additional possible error of 2 meV results. If the angle \( \theta \) between the two beams deviates from 180\(^\circ\), the Doppler correction changes accordingly. Given the tight collimation of the ion beam with \( \pm 0.6 \) mm at the entrance to the interaction region and \( \pm 1.1 \) mm at the exit, the maximum deviation in the angle \( \theta \) is 3.4 mrad corresponding to a shift of about 2 meV at 360 eV.

Another source of uncertainty is the stability of the photon source geometry which is taken into account in the calculation and the control of the actual photon energy in real time. Depending on temperature and operation mode of the synchrotron ring the electron beam may change its position in the ring. Changes in the position of the stored electron beam have an immediate impact on the photon energy transported to the experiment. Despite the tremendous achievements in beam position stability, drifts of several meV within a few hours of seemingly stable ring and monochromator operation have been observed. Moreover, the carefully calibrated in-vacuum angular encoders of both pre-mirror and grating introduce an uncertainty which has been reduced to a level of about 10 ppm which corresponds to 4 meV at 400 eV. During a period of one top-up ring-filling cycle the electron beam can move adding another 1 meV periodic energy shift. In summary, calibration uncertainties (one standard deviation) of the present measurements are estimated to be 40 meV at 360 eV and 50 meV at 440 eV.

III. THEORY

In the present theoretical treatment the energies of the doubly and singly excited states are calculated with relativistic many-body perturbation theory in an all-order formulation including single and double excitations, as described by Salomonson and Öster [67]. This means that all types of excitations that can be formed in a pure two-electron system are accounted for. The \( \text{C}^{4+} \) ion is placed in a spherical box within which a discrete radial grid is used. Diagonalization of the discretized hydrogenlike Dirac Hamiltonian gives a discrete basis set, complete on the grid chosen. The basis set is then used to construct correlated wave functions to all orders in the perturbation expansion of the electron-electron interaction. Here, both the Coulomb and the Breit interaction are accounted for. The perturbation expansion is constructed from an extended model space [68] whenever a state is dominated by two or more nearly degenerate configurations. An example is the \( 2s2p^3P_1 \) state which in \( jj \) coupling has major contributions both from the \( 2s2p_{1/2} \) and the \( 2s2p_{3/2} \) configuration. This is a common scenario in \( jj \) coupling, but also for doubly excited states in general. A multiple expansion of the electron-electron interaction is used, making the method applicable to many-electron atoms in general. The present calculations include all contributing partial waves up to \( \ell_{\max} = 10 \).

When perturbation theory is applied to autoionizing states it is obvious that the use of a discrete basis set will cause problems close to the poles in the energy denominator. A complex scaling of the radial coordinates can, however, solve this problem. The present treatment follows the method employed by Lindroth [69] for the calculation of doubly excited levels in the helium atom, and later for a number of Be-like ions (see, e.g., [70]). The method yields complex energies for the autoionizing states, where the imaginary part corresponds to the half-life time (due to Coulombic decay) of the state. The decay rates due to photon emission are calculated from the dipole matrix elements between the doubly excited states and the \((1s\ell n)^3L_J \) states with \( n \leq 4 \). For details see Ref. [71]. The radiative decay rates of all the considered doubly excited states are completely dominated (with contributions of more than 99\%) by photoemission events that require one-electron transitions only.

For the underlying theory of complex rotation (CR) and many-body perturbation theory (MBPT) the reader is referred to a review by Lindroth and Argenti [72] and references therein.

The additional contributions to the energies originating from the quantization of the electromagnetic field are treated with the procedure implemented in the MCDFGME (Multi Configuration Dirac Fock and General Matrix Element) code, developed by Desclaux and Indelicato [73–76]. It includes one-electron one-loop corrections (self-energy [77–81] and vacuum polarization) and two-loop one-electron corrections (two-loop self-energy, mixed self-energy, and vacuum polarization diagrams, Källén and Sabry potential contributions; see Ref. [82] and references therein) in the Coulomb field of the nucleus, although the result is almost completely (99\%) determined by the one-photon self-energy and the one-loop vacuum polarization. It also includes vacuum polarization due to the electronic potential, retardation beyond the Breit interaction, and the effect from the electron-electron interaction on the self-energy evaluated with the so-called Welton method [74,75,83]. The latter method has recently been tested against the model operator approach developed by Shabaev and collaborators [84,85] for \( Z = 18 \) [86] and relative differences of less than 0.01\% were found.

The normal mass shift is taken care of by the correction factor \( M/(m_e + M) \) where \( m_e \) is the electron rest mass and \( M \) the rest mass of the \( ^{12}\text{C} \) nucleus. The mass-polarization effect (specific mass shift) has also been considered. Calculations of this effect with the nonrelativistic formula resulting in shifts of
TABLE I. Contributions to the calculated total energies of the fine-structure components of C$^{4+}$ (2s 2p 3P) and the lowest-energy triplet state C$^{4+}$ (1s 2s 3S$_1$). The results are given in atomic units for $^{12}$C. To convert to eV multiply by $E_h \times M/(m_e + M)$ where $E_h$ is the Hartree energy in eV, $m_e$ is the electron rest mass, and $M$ the rest mass of the $^{12}$C nucleus. The factor $M/(m_e + M)$ accounts for the normal mass shift. With $E_h = 27.21138602(17)$ eV and $m_e = 5.48579900070(16) \times 10^{-4}$ u according to the 2014 CODATA (Committee on Data for Science and Technology) recommended values [64], the conversion factor is 27.21014177 eV.

<table>
<thead>
<tr>
<th>Contribution</th>
<th>2s 2p 3P$_0$</th>
<th>2s 2p 3P$_1$</th>
<th>2s 2p 3P$_2$</th>
<th>1s 2s 3S$_1$</th>
</tr>
</thead>
<tbody>
<tr>
<td>Coulomb interaction 1st order</td>
<td>-8.207838</td>
<td>-8.207217</td>
<td>-8.205970</td>
<td>-21.118982</td>
</tr>
<tr>
<td>Breit interaction 1st order</td>
<td>0.000075</td>
<td>0.000011</td>
<td>0.000019</td>
<td>0.000569</td>
</tr>
<tr>
<td>$\Delta$ Coulomb and Breit all orders</td>
<td>-0.027206</td>
<td>-0.027193</td>
<td>-0.027188</td>
<td>-0.312848</td>
</tr>
<tr>
<td>$\Delta$ One-electron radiative corrections</td>
<td>0.000114</td>
<td>0.000116</td>
<td>0.000119</td>
<td>0.001009</td>
</tr>
<tr>
<td>$\Delta$ Two-electron radiative corrections</td>
<td>-0.000006</td>
<td>-0.000007</td>
<td>-0.000009</td>
<td>-0.000032</td>
</tr>
<tr>
<td>Drake [65,66]</td>
<td></td>
<td></td>
<td></td>
<td>-21.430301</td>
</tr>
</tbody>
</table>

The entries in Table II include the information defining the spectroscopic notation for each of the resonances. As an example, the first row in the table is for the 2s 2p 3P$_0$ resonance. For the calculation of resonance cross sections the resonance parameters have to be known, i.e., the resonance energy $E_{res}$, the natural width $\Gamma$, the Fano asymmetry parameter $q$ [87], and the resonance strength $S$. In the framework of complex scaling the $q$ values can be obtained from transition matrix elements between the initial state and the resonances as described in Ref. [88]. To determine the contribution of a given resonance to the photoabsorption cross section, knowledge about the associated strength $S_{abs}$ is required. It is provided in the sixth column of Table II. Resonant photoabsorption determines the total intermediate population of the resonant level. This level can decay by autoionization or by photoemission. In the present experiments, the observed final channel was that of net single ionization of C$^{4+}$; i.e., the C$^{5+}$ final products were registered. To obtain the ionization cross section, the relative probability for Auger decay has to be known. This branching ratio follows from

$$B_a = \frac{A_a}{A_a + A_r}, \quad (9)$$

where, in the present context, $A_a$ is the Auger decay rate from the specified intermediate doubly excited level to C$^{5+}$ (1s) + e, and $A_r$ is the radiative-decay rate of the intermediate doubly excited level to all levels of the type C$^{4+}$ (1snrt) with $n \leq 4$ included in the calculation. The radiative decay rates have been calculated in first-order perturbation theory. The effect on the $q$ values from radiation damping, see, e.g., Ref. [89], is expected to be negligible here. The possibility that neighboring resonances can affect each other, as recently discussed in Refs. [90,91], has not been considered. The parameters thus obtained are provided in Table II. The resonance strength for ionization is given by

$$S_{ion} = B_a S_{abs}. \quad (10)$$

Thus, the cross section $\sigma_{II}$ for photoionization of C$^{4+}$ (1s 2s 3S$_1$) in the photon energy range 350 to 450 eV at most 0.1 meV. It is safe to say that mass-polarization shifts of the investigated levels of the heliumlike C$^{4+}$ ion are well below the 1 meV level.

Table I illustrates the importance of different contributions to the level energies of the 2s 2p 3P$_i$ fine-structure components. The C$^{4+}$ ion has a sufficiently low atomic number, $Z = 6$, so that Coulomb correlation dominates by many orders of magnitude over the Breit interaction and the radiative contributions. The one- and two-particle radiative corrections are given separately showing that the latter, which are less well known, contribute with only a few tenths of a meV to the transition energies and thus do not affect the comparison between theory and experiment on a measurable level. The contributions to the 1s 2s 3S$_1$ state, the lowest of the triplets, are also given in Table I.

The present calculation can be compared to the results obtained previously by Drake [65]. He used highly correlated nonrelativistic wave functions of Hylleraas type to calculate the ionization energy. Relativistic and radiative corrections are subsequently treated as perturbations. The calculation is thus very different from the one presented here. Yet, the two calculations agree to within 0.5 meV. The number at the bottom of Table I is obtained by adding Drake’s result for the ionization energy to the well-established values for hydrogenlike systems by Johnson and Soff [66].

IV. RESULTS

The main results of the present theoretical approach are displayed in Table II. Calculations were carried out for all levels within the configurations 2s 2p, 2s 3p, 2p 3s, 2p 3d, 2s 4p, 2p 4s, and 2p 4d that can be reached by electric-dipole non-spin-flip excitation from the metastable C$^{4+}$ (1s 2s 3S$_1$) level. Only 3P excited levels with total angular momenta $J = 0, 1, 2$ are dipole allowed and spin-conserving. Hence, within the selected configurations 21 levels had to be considered for calculating the cross sections for single photoionization of C$^{4+}$ (1s 2s 3S$_1$). These resonances were expected to provide sizable contributions to the total ionization which also includes direct removal of the 2s electron from the metastable 1s 2s 3S level. The threshold for direct 2s ionization is 93.131 eV [12] where the cross section $\sigma_{2s}$ is about 0.57 Mb. In the energy range of present interest, 350 to 450 eV, $\sigma_{2s}$ drops from 0.0306 Mb to 0.0159 Mb. Removal of the 1s electron from metastable C$^{4+}$ (1s 2s 3S$_1$) does not occur at energies lower than 460.608 eV [12].
can be represented as \[ \sigma_{\text{PI}} = \sigma_{2s} + \sum_{k=1}^{21} \frac{2S_{2s,n,k}}{\pi (q_k^2 - 1)^{1/2}} \left[ \frac{(q_k + \epsilon_k)^2}{(1 + \epsilon_k^2)} - 1 \right], \] (11)

that is, a sum over all intermediate autoionizing doubly excited levels characterized by their individual resonance parameters which are labeled with the resonance number \( k = 1, 2, \ldots, 21 \). The reduced energy \( \epsilon_k \) of the \( k \)th resonance is given by

\[ \epsilon_k(E) = \frac{2(E - E_{\text{res},k})}{\Gamma_k}. \] (12)

Figure 1 shows the theoretical cross section function in the energy range 350 to 450 eV. Seven groups of resonances are identified by their electronic configurations. The natural (unconvoluted) cross section covers a very large range from about \( 3 \times 10^{-3} \) to more than 1500 Mb. This can only be made visible in a logarithmic plot which optically overemphasizes the small direct-ionization cross section \( \sigma_{2s} \). Destructive interference of the resonant and direct ionization pathways results in deep but narrow dips in the cross-section function. The natural widths range from \( 4.14 \times 10^{-4} \) to \( 9.69 \times 10^{-3} \) eV spanning more than an order of magnitude. Clearly, such narrow features cannot be resolved in an overview plot like the present one.

For the illustration of details within the groups of \( 2(nl'c) \) resonances, Fig. 2 shows the natural (unconvoluted) cross sections of the \( 2s4p \) resonance levels [panel (a)], of the \( 2p4s \) resonance levels [panel (b)], and of the \( 2p4d \) resonance levels [panel (c)]. The cross sections for photoabsorption and for photoionization are shown by dotted and solid lines, respectively. According to Eq. (9) they differ by the factor \( B_a \), the branching ratio for autoionization. For doubly excited
states of low-$Z$ ions, it is common knowledge that Auger decay probabilities typically exceed radiative decay rates by far. This is confirmed by Fig. 2(a) for the $2s4p$ resonance group, where the branching ratios $B_s$ are approximately 0.85 (see Table II, 9th column, where the dominance of Auger over radiative decay becomes all the more obvious for the $2s2p$ and $2s3p$ resonances). However, the situation is very much different for the $2p4s$ (as well as $2p3s$) resonances where $B_s$ is between 0.10 and 0.19 indicating that radiative stabilization now dominates the decay. For the $2p4d$ (and a little less pronounced also for $2p3d$) resonances this effect is dramatic. Branching ratios $B_s$ between 0.014 and 0.00037 are predicted. Once populated, the $2p4d$ $^3P$ autoionizing levels decay almost exclusively by emission of photons and therefore, they contribute little to the ionization cross section. This is the reason why the first photospectroscopy experiments producing doubly excited levels of helium could actually see a signal. It also provides a rationale for the surprise result of Kasthurirangan et al. [25] who saw photon emission from doubly excited $2p3d$ $^3P$ heliumlike Si, S, and Cl ions.

It should be mentioned in the context of Fig. 2 that experimentally resolving the individual fine-structure components that are clearly separated in the theory plots would require a resolving power $E/\Delta E$ greater than 50 000 for the $2s4p$ resonances and greater than 100 000 for the $2p4d$ resonances. This is presently not experimentally achievable. Moreover, with realistic photon-beam energy bandwidths of 100 meV available at about 438 eV the narrow $2p4d$ resonances would produce apparent photoionization cross sections on the order of 0.1 Mb, far too small to be seen in the present experiments given the relatively high background and the small flux of metastable $C^{4+}$ parent ions available.

In Table III the MBPT results for excitation energies, Auger rates, and rates for radiative decay to $1snt$ states from $2s2p$, $2s3p$, $2p3s$, and $2p3d$ levels are compared to data obtained by Goryaev et al. [33] applying the $Z$-expansion method. Previous results [92,93] were superseded by the new improved calculations which, among other refinements, account for relativistic corrections within the framework of the Breit operator. The comparison shows that the excitation energies from the present theory and the calculations of Goryaev et al. agree within deviations of at most 0.05 eV, i.e., within 120 ppm. In this context, one has to keep in mind that the uncertainties of the present, calculated resonance energies are below 1 meV, i.e., definitely less than 3 ppm. The largest deviations are found for the $2p3d$ energies. For the other resonances the maximum deviation is 0.015 eV corresponding to 36 ppm. The comparison is similarly satisfying for the radiative rates with a maximum difference of 11%. For the $2s2p$ and $2s3p$ resonances, deviations of autoionization rates are within 7%. The deviations increase for the $2p3s$ resonances and reach a factor of almost 2.6 and for the $2p3d$ resonances there are differences in the two calculations of Auger decay rates by factors reaching almost up to 50. While the former numbers could not be tested by the present experiments because of the too small signal rates, experimental results were obtained for $2s2p$, $2s3p$, and $2p3s$ resonances. The measurements support the present theoretical resonance strengths which are immediately related to the radiative and Auger decay rates.

By a number of experimental multirange energy scans at a fixed monochromator exit-slit width of 500 μm, an overview of the resonance features within the energy region of interest was obtained at a resolving power of approximately 2800. Three groups of resonances in the energy ranges 358.6 to 359.6 eV, 418.4 to 419.5 eV, and 437.0 to 438.5 eV, respectively, were covered in a single scan. Such scans were repeated several times to improve the signal-counting statistics. For the $2p3d$ and $2p4d$ resonances the predicted resonance strengths for ionization are about three orders of magnitude below the strengths of the $2s2p$ resonances (see Table II). Due to low count rates and relatively high background, those small contributions could not be measured. Figure 3(a) shows the result of the combined overview scans covering the lower-energy region with the three $2s2p$ $^3P_{0,1,2}$ resonances. The photon energy scale was calibrated following the procedure described at the end of Sec. II. The experimental overview-scan spectrum was obtained on a relative scale and had initially only been normalized to the ion current and the photon flux. The whole spectrum was then multiplied by a constant factor such that the area of the measured $2s2p$ peak shown in Fig. 3(a) matches the theoretical ionization strength. The solid red line in this figure

![Graphs showing cross sections and photon energies](image)
was obtained by convoluting the theoretical spectrum with a Gaussian (the experimental response function at 500 μm monochromator exit-slit width does not appear to be strictly Gaussian, though). The set of three Gaussian-convoluted Fano profiles [94] for the three $^3P$ resonances that resulted from the present theory was adjusted to the experimental data in a least-squares fit by leaving a global energy shift and the Gaussian width free for the fit while all other parameters were kept fixed at the values provided in Table II. By this procedure the Gaussian width was determined to be 116 meV and the overall shift of the experimental to the theoretical $2s2p$ spectrum was found to be $-1.4$ meV. In Fig. 3 all data are shown without any energy shifts; i.e., the experimental and theoretical spectra are based on independent energy scales.

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<th>$E_{\text{GVU}}$ (eV)</th>
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The data displayed in Fig. 3(b) are yet another factor of about 3 down from those in Fig. 3(b). The statistical scatter is relatively large. Nevertheless, signals from the $2s4p\ ^3P_{0,1,2}$ and $2p4s\ ^3P_{0,1,2}$ resonances could be observed. Again, as in Fig. 3(b), the experimental spectrum is shown on the experimentally determined energy scale without any further manipulation beyond the normalization to the $2s2p$ peak at 359 eV. A fit of the experimental data similar to the one applied to the peak at 359 eV suggests a bandwidth of 165 meV and a global shift of the $2s4p\ ^3P_{0,1,2}$ and $2p4s\ ^3P_{0,1,2}$ resonance group relative to the present theory by $+19.0$ meV. As in Fig. 3(a), the theoretical data, convoluted with a 167-meV FWHM Gaussian, and the experimental cross sections are shown on their individual, intrinsically determined energy scales. It is worth noting that the “apparent” cross sections (after convolution with the experimental response function) are down in peak height from the (also convoluted) $2s2p$ resonances by approximately a factor of 15.

The data displayed in Fig. 3(c) are yet another factor of about 3 down from those in Fig. 3(b). The statistical scatter is relatively large. Nevertheless, signals from the $2s4p\ ^3P_{0,1,2}$ and $2p4s\ ^3P_{0,1,2}$ resonances could be observed. Again, as in Fig. 3(b), the experimental spectrum is shown on the experimentally determined energy scale without any further manipulation beyond the normalization to the $2s2p$ peak at 359 eV. A fit of the experimental data similar to the one applied to the peak at 359 eV suggests a bandwidth of 165 meV and a global shift of the $2s4p\ ^3P_{0,1,2}$ and $2p4s\ ^3P_{0,1,2}$ resonance group relative to the present theory by $+40.5$ meV. Accordingly, the present theoretical results were convoluted with a 165-meV FWHM Gaussian for the comparison with the experimental data. The differences of the derived photon-energy spreads are observed at a constant monochromator exit-slit width of 500 μm. They are in the range of a resolving power of $2950 \pm 450$. Excursions from a constant resolving power can partly be attributed to statistical fluctuations and partly to the fact that the photon optics were optimized for the resonance group at 359 eV and the quality of the settings drifted away from the optimum with increasing photon energy.

![Figure 3](image-url)

**FIG. 3.** Cross sections for photoionization of C$^{1+}$($1s2s\ ^3S$) ions. The experimental data with statistical error bars were obtained at constant 500 μm exit-slit width of the monochromator. The solid red line is the present theory convoluted with Gaussians of appropriate FWHM values: (a) 116 meV, (b) 167 meV, (c) 165 meV. The experimental spectrum was obtained in one multirange scan and normalized such that the area of the first peak matches the theoretical resonance strength.
The measurement was carried out with a C4 and its combined statistical and systematic, i.e., total, uncertainty. Each experimental data point is shown with its statistical uncertainties. The measurement was carried out with a C4 ion beam containing an unknown fraction \( f \) of metastable ions. The solid red line represents the present theoretical cross section after convolution with a 16-meV FWHM Gaussian and multiplication by a factor 0.105 which is interpreted to represent \( f \).

The theoretical and experimental results in Fig. 3 are remarkably consistent. Only a single constant factor was applied to the experimental scan spectrum for normalization to the theoretical resonance positions. The individual contributions of the fine-structure components are shown by thin dotted lines. The experimental spectrum was shifted up by 1.4 meV to match the theoretical resonance positions. Each experimental data point is shown with its statistical uncertainty of only 0.003 out of 0.105, i.e., \( f = 0.105 \pm 0.003 \). The absolute uncertainty is about 15% due to the systematic uncertainties of the cross-section measurements.

In an absolute measurement only a defined part (about 50 cm) of the total photon-ion interaction length (approximately 175 cm between the merger and the demerger) is used. Hence, the count rates in an absolute measurement are reduced, typically by a factor of 2, and it takes more time to accumulate counts for a desired level of statistics. A more economical way to get better statistics is to use the whole interaction path length and, by that, obtain a relative cross section which can then be normalized to a measurement like the one displayed in Fig. 4. For best possible energy resolution, the fixed-focus constant \( c_f \) [49,95] of the variable-line-spacing (VLS) grating in use at beamline P04 was carefully adjusted, and thus the bandwidth could be reduced to 13.91 ± 0.23 meV. Under these conditions, 19 sweeps over the energy range 359.00 to 359.17 eV with 1 meV step size and 10 s dwell time each were accumulated during several days. The resulting relative spectrum was normalized to the absolute measurement shown in Fig. 4 and then normalized again to 100% metastable C4+ ions via 2s2p \( ^3P \) resonances. The experimental cross sections with statistical error bars are normalized to 100% metastable parent ions. The solid red line represents the present theoretical cross section after convolution with a 13.91-meV FWHM Gaussian which corresponds to a resolving power \( E/\Delta E = 25 \ 800 \). The individual contributions of the fine-structure components are shown by thin dotted lines. The experimental spectrum was shifted up in energy by 1.4 meV.
13.91-meV FWHM Gaussian. At the resulting resolving power of more than 25 000 the presence of the $2s2p^3P_0$ contribution starts to show as a hump on the low-energy side of the $3P_1$ peak. No photoionization experiment with ions resolving fine-structure components of a given deep-inner-shell resonance term at such high energies has been reported so far [42].

Similar energy-scan measurements were performed for the $2s3p^3P_{0,1,2}$ resonances in the energy range 418.689 to 418.821 eV at 1 meV step size and at comparable resolving power. The energy spread was determined by a fit yielding 17.0 ± 0.5 meV. The experimental peak area was normalized to the known ratio of the $2s3p$ and $2s2p$ photoionization resonance strengths (see Fig. 3) considering the absolute apparent strength of the $2s2p$ resonance (see Fig. 4) and the fraction $f$ of metastable parent ions in the beam. The absolute cross section for (100%) metastable $C^4+(1s2s^3S)$ ions thus obtained is compared in Fig. 6 with the theoretical data convoluted with a 17-meV FWHM Gaussian. As for the $2s2p$ resonances (see Fig. 5) theory and experiment are in excellent agreement with the assumption of a metastable fraction $f$ of 10.5%. At the present resolving power the $2s2p^3P_2$ resonance is partly resolved from the $2s2p^3P_{0,1}$ sum peak.

The present results on photoionization of $C^4+(1s2s^3S)$ can be tested against previous absolute measurements on dielectronic recombination (DR) of $C^5+(1s^2)$ ions [96]. The basis for the comparison is the principle of detailed balance [41,97–103] which, in turn, is based on time-reversal symmetry in atomic processes. The present theoretical calculations of $C^4+$ photoionization cross sections also provide all the information to directly infer $C^5+$ DR cross-section contributions proceeding via resonances that are associated with $C^4+(n\ell n'\ell' 3P_{0,1,2})$ doubly excited states with $n = 2$ and $n' = 2, 3$,

$$e + C^5+(1s^2S_{1/2}) \rightarrow C^4+(n\ell n'\ell' 3P_{0,1,2}) \downarrow$$

$$C^4+(1s2s^3S_1) + \gamma.$$  \hspace{1cm} (13)

The present calculations were extended to include all the ten possible $C^4+(2\ell 2\ell'')$ product levels with $\ell, \ell' = s, p$ that can be populated in DR of $C^5+(1s^2S)$ ions.

The $C^5+$ DR experiment [96] was one of the first such measurements carried out at a heavy-ion storage ring with an electron cooling device. In the early stage of the DR measurements the energy resolution at energies around 270 eV was limited and amounted to approximately 2.5 eV. Later developments of cold electron targets [104] have substantially improved the energy resolution in DR experiments. With present state-of-the-art experimental equipment the energy spread in the electron-ion center-of-mass system could be as low as 0.25 eV at 270 eV with a longitudinal electron beam temperature $k_B T_e = 2 \times 10^{-5}$ eV [104,105] where $k_B$ is Boltzmann’s constant.

The only existing DR experiment [96] covered the electron energy range 260 to 380 eV which includes all resonances associated with $2\ell n\ell'$ configurations with $n = 2, 3, 4, \ldots; \ell, \ell' = 0, 1, \ldots, n - 1$. At the limited resolution of the experiment, structure within these configurations could only be observed for $n = 2$. All the ten possible levels within the $2\ell 2\ell''$ configurations contribute to DR of $C^5+(1s^2S_{1/2})$. However, only the $2s2p^3P$ term can contribute to photoionization of $C^4+(1s2s^3S)$ due to the selection rules for electric dipole transitions. Hence, for a comparison of DR of $C^5+(1s^2S_{1/2})$ with photoionization of $C^4+(1s2s^3S)$ and for exploiting time-reversal symmetry, only the $2s2p^3P_{0,1,2}$ resonances can be considered.

Time-reversal symmetry and the principle of detailed balance relate the cross section $\sigma_{DR}$ for the DR process characterized by Eq. (13) with the cross section $\sigma_{PI}$ for the photoionization process described by Eq. (2) with $n, n' = 2$ on a level-to-level basis,

$$\sigma_{DR} = \frac{g_f}{g_i} \frac{E_\gamma^2}{2m_e c^2 E_e},$$  \hspace{1cm} (14)

where the quantities $g_i = g_i(1s2s^3S) = 3$ and $g_f = g_f(1s^2S) = 2$ are the statistical weights of the initial and final levels of the photoionization process, respectively, and $m_e$ is the electron rest mass. The photon energy $E_\gamma$ and the electron energy $E_e$ in the processes related to one another by time reversal are connected by the condition

$$E_e = E_\gamma - I_{bind},$$  \hspace{1cm} (15)

where $I_{bind} = 93.131$ eV is the ionization energy of the $C^4+(1s2s^3S_1)$ level [12]. Thus, the $2s2p^3P$ resonance group found at about 359 eV in photoionization of $C^4+(1s2s^3S_1)$ has to appear at about 266 eV in DR of $C^5+(1s^2S_{1/2})$.

The DR cross section for a transition $i \rightarrow j \rightarrow f$ from an initial state $i$ to a final state $f$ via an intermediate resonance $j$
FIG. 7. Cross sections for dielectronic recombination (DR) of C\(^{5+}\)(1s \(^7\)S\(_{1/2}\)) in the energy range of 2\(\ell\)2\(\ell\)\('\) resonances. The solid circles with statistical error bars show the results obtained at a heavy-ion storage ring [96]. The solid red line represents the present DR calculation. The energies of the individual resonance contributions are shown by vertical bars and the associated doubly excited levels are indicated. The individual-term contributions to the theoretical DR spectrum are also shown and can be identified by their resonance energies. In particular, the solid black line shows the contribution of the 2s\(^2\)1\(^1\)S\(_0\) resonance and the dashed blue line represents the contribution of the 2s2p\(^3\)P term. For comparison, the contributions of the 2s\(^2\)1\(^1\)S resonance (small gray shaded peak) and the 2s2p\(^3\)P term (gold shaded peak) were inferred from the decay rates provided by Goryaev et al. [33].

\[
\sigma_{\text{DR}} = \frac{\hbar^3 \pi^2 g_j A_j(j \rightarrow i) \sum_f A_i(j \rightarrow f')}{m_e 2E_i g_i} \times \frac{\Gamma(j)}{2\pi (E_i - E_{\text{res}})^2 + \Gamma(j)^2/4},
\]

(16)

where \(g_j\) is the statistical weight of the resonant level \(j\), \(A_j(j \rightarrow i)\) is the Auger decay rate of the resonant level \(j\) to the initial level \(i\), \(A_i(j \rightarrow f')\) comprises all radiative rates for the resonant level \(j\) decaying to all bound levels \(f'\), \(E_{\text{res}}\) is the resonance energy at which level \(j\) is populated, and the total natural width of the resonance \(j\).

\[
\Gamma(j) = \hbar \left[ \sum_i A_i(j \rightarrow i') + \sum_{f'} A_i(j \rightarrow f') \right].
\]

(17)

Figure 7 shows the DR results [96] obtained in the range of the 2\(\ell\)2\(\ell\)\('\) resonances. The data have been corrected for an error in the original analysis (see [107]). The ten possible levels within the 2\(\ell\)2\(\ell\)\('\) manifold are partly resolved. Three peaks can be identified. The energies and Auger rates of the 2\(\ell\)2\(\ell\)\('\) resonances have been calculated within the framework of the present theory and the radiative decay rates have been calculated within first order perturbation theory as previously discussed. All the quantities needed for Eq. (16) are thus at hand. The result is displayed in Fig. 7, where the energy positions are indicated by the vertical bars and labeled by the spectroscopic notation of the associated levels. A detailed account of the present method for calculating DR cross sections has been provided previously by Tokman et al. [71]. The individual cross-section contributions of all levels have also been calculated. Solid lines in Fig. 7 represent the contributions of the six terms associated with 2\(\ell\)2\(\ell\)\('\) configurations. The sum of these contributions (the red line) provides a very good representation of the experimental peak areas. Deviations of experimental and theoretical resonance energies illustrate the limitations of the early DR measurement for which an energy uncertainty of ±1 eV at 500 eV was quoted.

In addition to the present theoretical calculations and for comparison, the DR contributions arising from the 2s\(^2\)1\(^1\)S and 2s2p\(^3\)P resonance terms were inferred from the decay rates provided by Goryaev et al. [33] using Eq. (16). Since the DR spectrum was measured at an electron-energy bandwidth of approximately 2.5 eV the cross sections obtained via Eqs. (14) and (16) have to be convoluted with a 2.5-eV FWHM Gaussian to simulate the experimental conditions of the DR measurement. Excellent agreement of the resonance energies and resonance strengths resulting from the two independent theoretical approaches is observed. On the scale of the figure, the theoretical results cannot be distinguished from one another.

Table IV provides DR resonance strengths \(S\) in kb eV for DR of C\(^{5+}\)(1s \(^7\)S\(_{1/2}\)) leading to the 2s2p\(^3\)P and 2s\(^2\)1\(^1\)S resonance terms together with the sum \(\Sigma = S(2s2p\(^3\)P) + S(2s\(^2\)1\(^1\)S)\) determined by different experiments and theoretical calculations. It is obvious from Fig. 7 that the first peak in the spectrum is a blend of contributions from the 2s\(^2\)1\(^1\)S and 2s2p\(^3\)P terms. The present theoretical calculations and the strengths inferred from the decay rates determined by Goryaev et al. [33] provide a very consistent picture with almost identical results. Both theoretical approaches predict a strength \(S(2s\(^2\)1\(^1\)S) = 1.4\) kb eV which is only about 5% of the strength \(S(2s2p\(^3\)P) \approx 28\) kb eV.

The DR experiment with C\(^{5+}\) ions [96] yields \(S(2s2p\(^3\)P) + S(2s\(^2\)1\(^1\)S) = 28.7\) kb eV. The uncertainty of this strength is directly related to the systematic uncertainty of the experimental cross sections which has been quoted to be ±15%. By subtraction of the theoretically predicted, relatively small contribution \(S(2s\(^2\)1\(^1\)S) = 1.4\) kb eV the strength \(S(2s2p\(^3\)P) = 27.3\) kb eV results. This number has to be compared with the strength \(S(2s2p\(^3\)P) = 28.1\) kb eV inferred from the present photoionization experiments and by exploiting the principle of detailed balance. The difference is less than 3% and thus very much smaller than the systematic uncertainties of both the present photoionization and the previous DR cross-section measurements.

The comparison of the present photoionization experiments with the independent measurement of DR cross sections provides additional support for the present photoionization data and, hence, also for the fraction \(f = 0.105\) of 1s2s\(^3\)S metastable ions in the C\(^{4+}\) beam that was employed in the photoionization experiments.

V. SUMMARY AND OUTLOOK

In the present photoionization experiments with metastable C\(^{4+}\)(1s2s\(^3\)S) ions the energy range of doubly excited (empty-K-shell) 2s2p, 2s3p, 2p3s, 2p3d, 2s4p, 2p4s,
TABLE IV. DR strengths in kb eV of the $2s2p\,^3P$ and $2s^2\,^1S$ resonance terms in $^{12}\text{C}^{++}$ and their sum $\Sigma$ determined in different experiments and by theoretical calculations. For the fraction $f$ of C$^{++}$ ($1s2s\,^3S$) metastable ions in the parent ion beam used in the present photoionization experiments the comparison with the present photoionization theory suggests $f = 0.105$. This fraction has been used for calculating the entry of $S(2s2p\,^3P) = 28.1$ kb eV in the last row. Comparison with the DR measurement instead would have resulted in $f = 0.108$.

<table>
<thead>
<tr>
<th>Data Source</th>
<th>$S(2s2p,^3P)$</th>
<th>$S(2s^2,^1S)$</th>
<th>$\Sigma$</th>
</tr>
</thead>
<tbody>
<tr>
<td>Present MBPT and CR theory</td>
<td>28.1</td>
<td>1.4</td>
<td>29.5</td>
</tr>
<tr>
<td>Inferred from Z-expansion decay rates $[33]$</td>
<td>28.3</td>
<td>1.4</td>
<td>29.7</td>
</tr>
<tr>
<td>DR experiment with C$^{++}$ ions $[96]$</td>
<td>$28.7 - 1.4 = 27.3$</td>
<td></td>
<td>28.7</td>
</tr>
<tr>
<td>Present time-reversed photoionization</td>
<td>28.1</td>
<td></td>
<td></td>
</tr>
</tbody>
</table>

and $2p4d$ resonances was investigated and absolute apparent cross sections were measured with a mixed beam of ground-state and metastable $^3S$ ions. By comparison with the results of the present relativistic many-body perturbation theory (RMBPT) it was possible to determine the fraction $f = 10.5\%$ of the $^3S$ metastable component. The resolving power (25 800) of the experiment was sufficient to resolve fine structure in the doubly excited $2s2p\,^3P_{0,1,2}$ and $2s3p\,^3P_{0,1,2}$ resonance manifolds. The experimental resonance energies agree with the present RMBPT calculations within the experimental uncertainties. The maximum deviation of experiment from theory is about 40 meV at 437 eV. Given the fact that the fraction $f$ of the metastable component of the ion beam was not independently determined, the present experiment could only test the relative sizes of the dominant theoretical resonance contributions. Apart from a constant factor that is associated with the metastable fraction $f$ theoretical and experimental photoionization cross sections are in excellent agreement. A test of the theoretical C$^{++}$ ($2s2p\,^3P$) photoionization resonance contribution on an absolute scale is possible by comparison with measured absolute cross sections for dielectronic recombination of C$^{5+}$ and exploiting the principle of detailed balance for time-reversed processes. This comparison shows that a fraction $f = 0.105$ is consistent with both the C$^{4+}$ photoionization and the previous, independent C$^{5+}$ dielectronic recombination experiments.

A particular effort was made to determine the resonance energies with high accuracy. For this purpose, the RMBPT calculations were performed including all orders. Relativistic effects as well as QED up to the level of second-order contributions were included. The energies thus obtained are estimated to have a maximum uncertainty of 0.001 eV. Such an accuracy makes the heliumlike C$^{4+}$ ion a promising candidate for being used as a primary reference standard for the soft-x-ray region with an uncertainty that is roughly a factor of one hundred better than the present neutral-gas standards in the energy range 300 to 1000 eV. It will be interesting to study ionization of heliumlike ions in the $1s2s\,^3S_1$ level along the associated very fundamental isoelectronic sequence. Further work on this topic is underway.

The very high accuracy of resonance energies obtained by the present theoretical treatment may be exploited in an envisaged effort at the PIPE setup to generate new secondary reference standards for the calibration of soft-x-ray beamlines at synchrotron-radiation sources. A viable scenario would be the transfer of the calibration obtained from the very accurately calculated photoionization-resonance energies of selected heliumlike (and lithiumlike [61,62]) ions to the neutral-gases-based calibration standards that are presently used. Full advantage of the existing high-precision theoretical resonance energies can only be taken if an improved control of the photon-beam energy at beamline P04 can be realized which includes the stable positioning of the electrons circulating in the PETRA III storage ring.

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