

# Alignment of the CMS Silicon Tracker using Millepede II

DPG, Wuppertal

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## Outline:

- Alignment procedure at the CMS experiment
- The problem of weak modes
- Solutions

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# CMS tracker

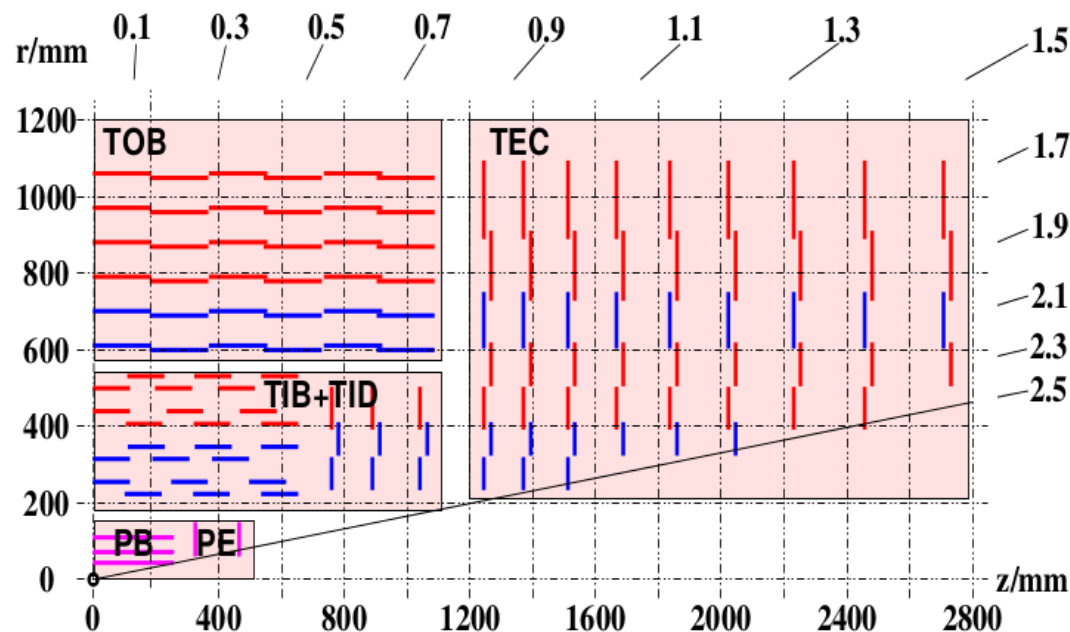
## Silicon tracker :

1440 pixel:

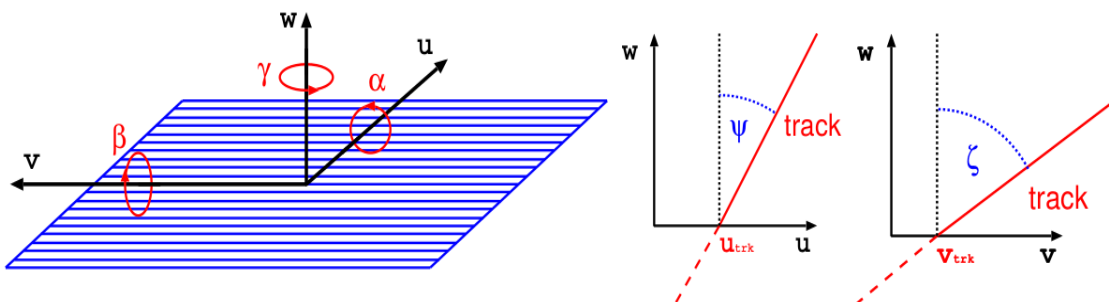
FPIX BPIX

~15118 strip modules (24244 sensors)

TEC TIB TOB TID



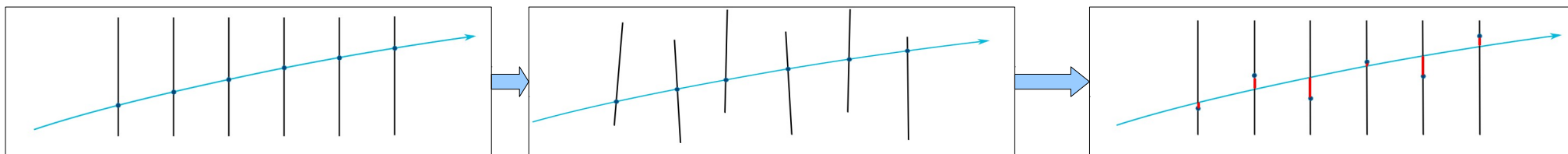
## Displacement of the modules :



We have set of parameters  
for each module :

up to 3 parameters of position  
3 parameters of orientation  
3 parameters of deformation

# Global fit approach



Doing the alignment :

◆ Calculate normalized residuals,  $z_{ij}$  :

$$z_{ij} = \frac{u_{ij;m} - u_{ij;p}(\boldsymbol{\tau}_j, \mathbf{p})}{\sigma_{ij}},$$

$j$  – track numbers ,  $i$  – track's hits numbers ,  $u_{ij;m}$  – measured value of the hit position,  
 $u_{ij;p}$  – predicted value, depend on the vector of track parameters  $\boldsymbol{\tau}_j$  and the alignment parameters  $\mathbf{p}$   
that describe the position, orientation and deformation

- ◆ Alignment algorithm minimizes the  $\chi^2$  function:  $\chi^2(\boldsymbol{\tau}, \mathbf{p}) = \sum_j \left( \sum_i z_{ij}^2(\boldsymbol{\tau}_j, \mathbf{p}) \right),$
- ◆ Linearize the minimization problem
  - ◆ Minimization of the function  $\chi^2(\boldsymbol{\tau}, \mathbf{p})$  , leads to a system of linear equations
  - ◆ Solution of the linear equations is a complex mathematical problem
    - 200000 equations
    - Number of  $\boldsymbol{\tau}$  parameters – in the order of millions!

Organization of **MillePede II** program which is using for alignment :

- Considers all correlations between track and alignment parameters
- The minimization of the function  $\chi^2$  leads to a **system of linear equations**
- In the global case we have ( from normal equations of least squares ) :

$$\mathbf{Ca} = \mathbf{b},$$

where **C** - symmetric n-by-n matrix (measured),  
**a** - vector of alignment (global) and track (local) parameters  
**b** - products of residuals and global derivatives (measured)

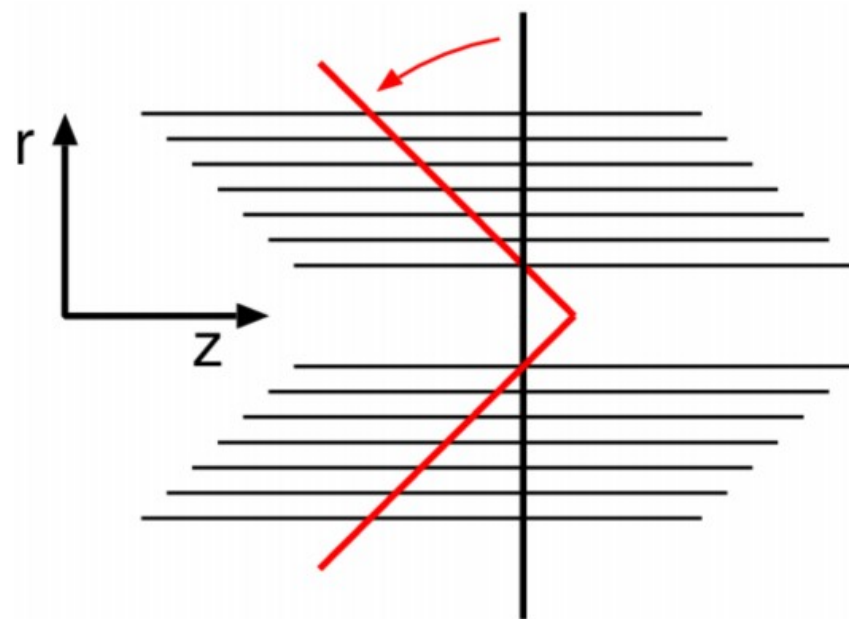


- Individual tracks are independent of each other → block structure of the equation matrix
- Structure can be exploited to reduce the matrix size
- Will get  $\mathbf{C}'\mathbf{a}' = \mathbf{b}' \rightarrow \mathbf{a}' = \mathbf{C}'^{-1}\mathbf{b}'$
- Have smaller equation system, where  $\mathbf{a}'$  contains only the alignment parameters
- The achieved statistical accuracy on the module alignment is found to be significantly better than  $10 \mu m$

Program also include **constraints** between module parameters

# Weak modes introduction

- Procedure is doing the fit with all tracks parameters and with all modules parameters simultaneously
- There are situations when coherent changes of alignment parameters can be compensated by changes of the track parameters → **WEAK MODES**
- Weak modes bias the track measurements:
  - curvature →  $p_t$  bias
- In this analysis we studied the so-called **SAGITTA** and **TWIST** weak modes

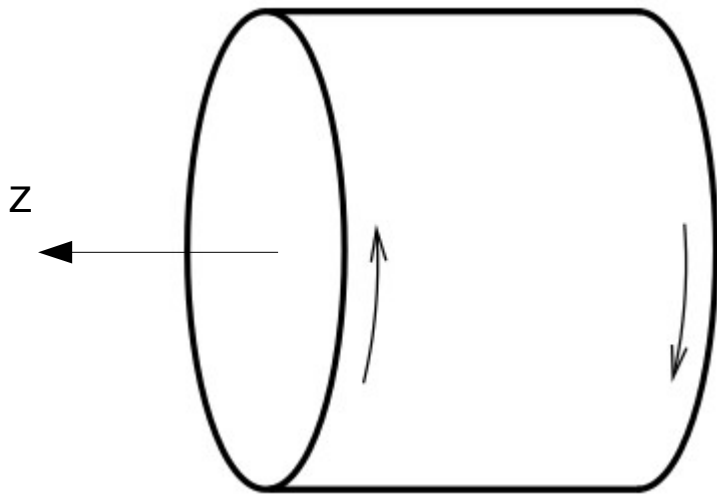


Examples of  $\chi^2$ -invariant deformation

# Twist weak mode

**TWIST** is a deformation of the tracker, in which the modules are moved **coherently** in  $\varphi$  by an amount directly proportional to their longitudinal position :

$$\Delta\varphi = \textit{const} \cdot z$$



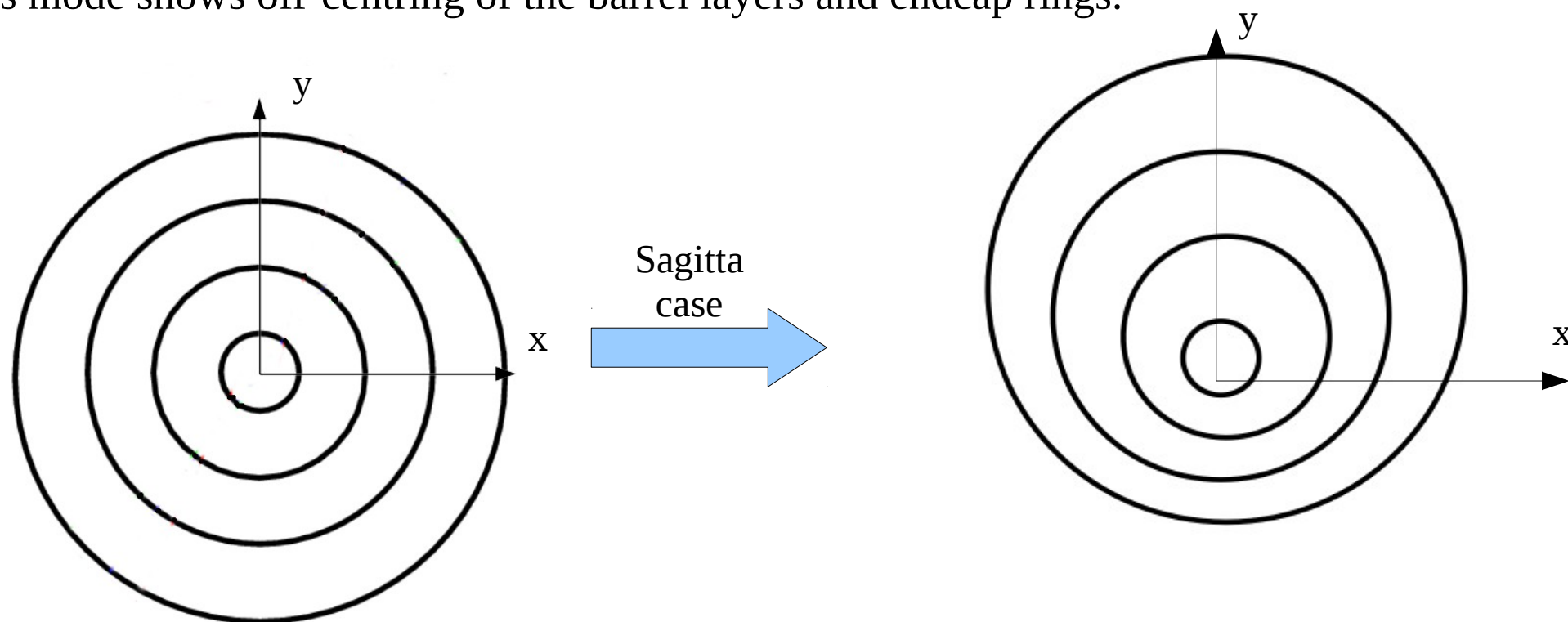
- Modules are moved in  $\varphi$  by an amount proportional to  $z$  position
- Further from the center – large displacements
- Compensated by changes of the track parameters

# Sagitta weak mode

**SAGITTA** is a deformation of the tracker, in which the modules are moved **coherently** in **Y** by an amount directly proportional to their **radius  $r$**  (in a cylindrical coordinate system) :

$$\Delta y = \text{const} \cdot r$$

This mode shows off-centring of the barrel layers and endcap rings.



In this case, again deformation of the detector could leave  $\chi^2$  without significant changes

# Weak modes control

To control weak modes it is crucial to include additional data samples:

- ◆ Cosmic ray tracks that break the cylindrical symmetry (help in Twist & Sagitta cases)
- ◆ Straight tracks without curvature, recorded when the magnetic field is off (help in Twist & Sagitta cases)
- ◆ Knowledge about the production vertex of tracks
- ◆ Knowledge about the invariant mass of a resonance (Z boson) (help in Sagitta & Twist cases)

Procedure to controle weak modes:

- ◆ Took recent well converged alignment with 2012 data and use misalignment scenario (Sagitta/Twist)
- ◆ Using all data samples (cosmic, collision data , 0T, 4T)
- ◆ Apply alignment procedure
- ◆ Compare obtained geometry with legacy alignment object



# Misalignment recovery with twist mode

Twist ( $\Delta\phi = c \cdot z$ )

Twist  
misalignment

Re-alignment

- no 0T cosmic data
- no  $Z_{\mu\mu}$

PXB

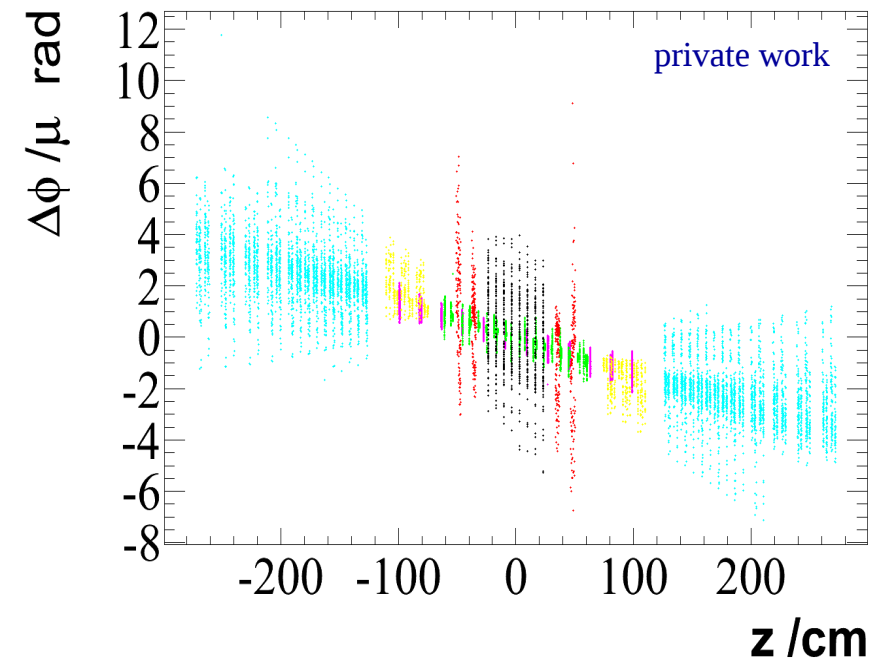
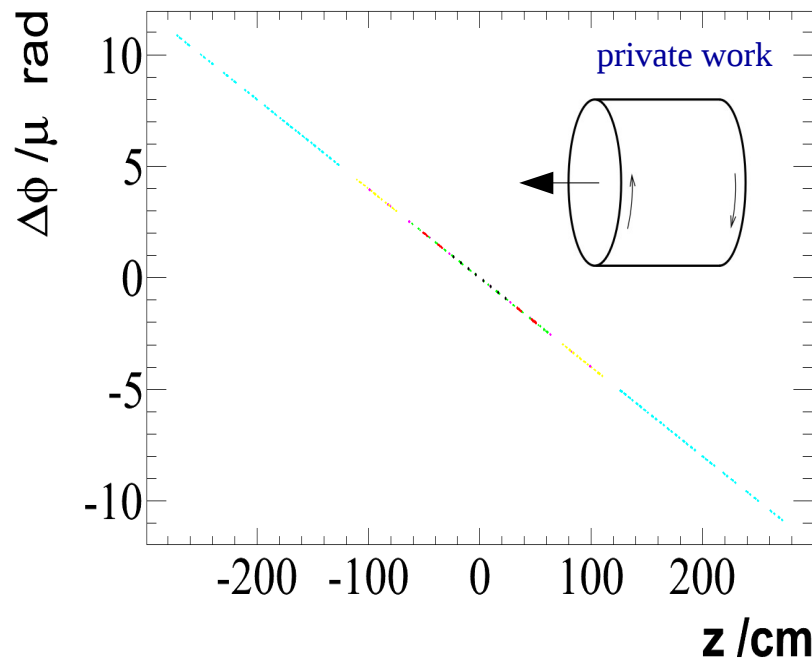
PXF

TIB

TID

TOB

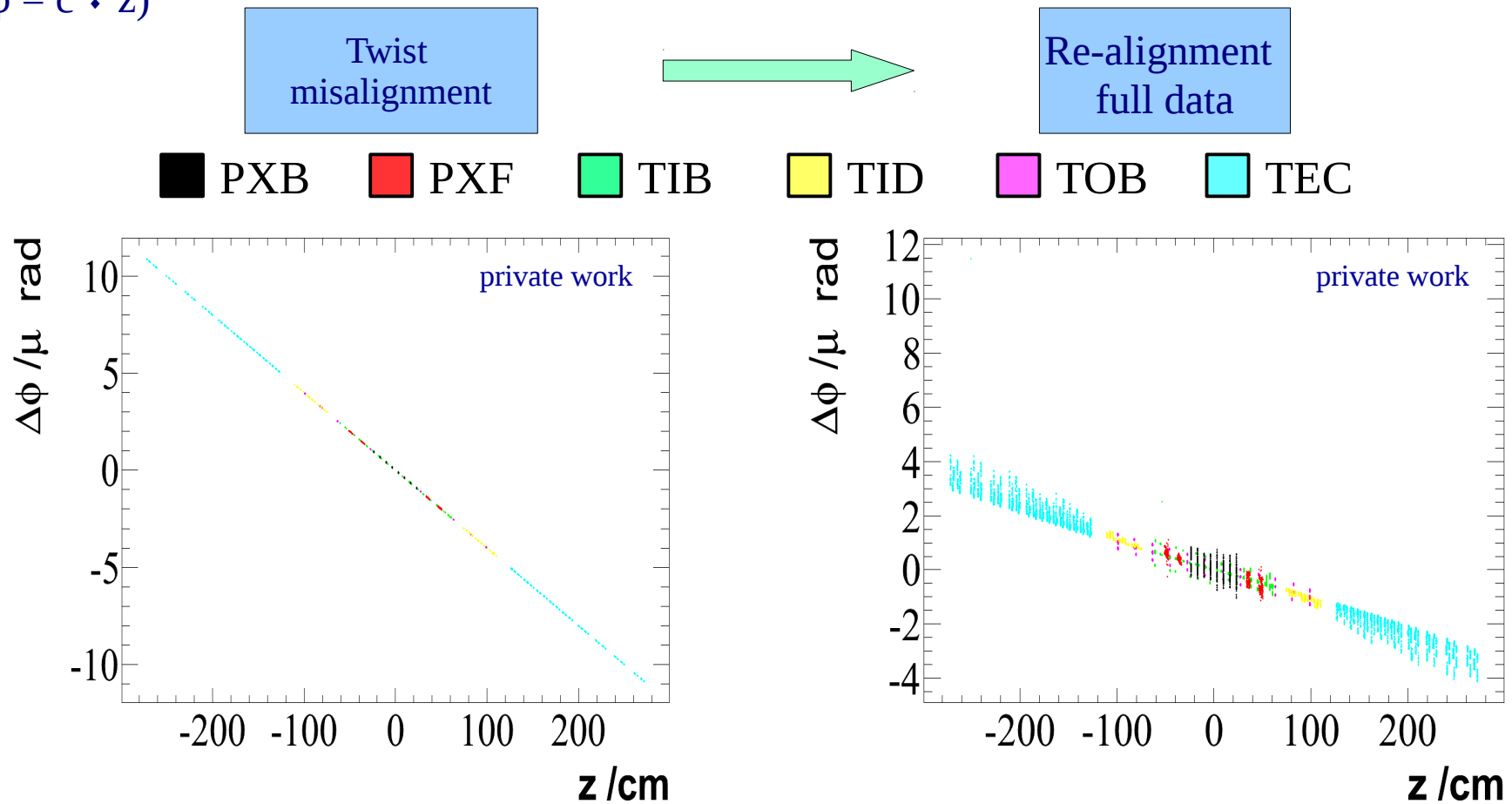
TEC



- Alignment procedure makes some improvements
- Need to use different data types

# Misalignment recovery with twist mode

Twist ( $\Delta\phi = c \cdot z$ )

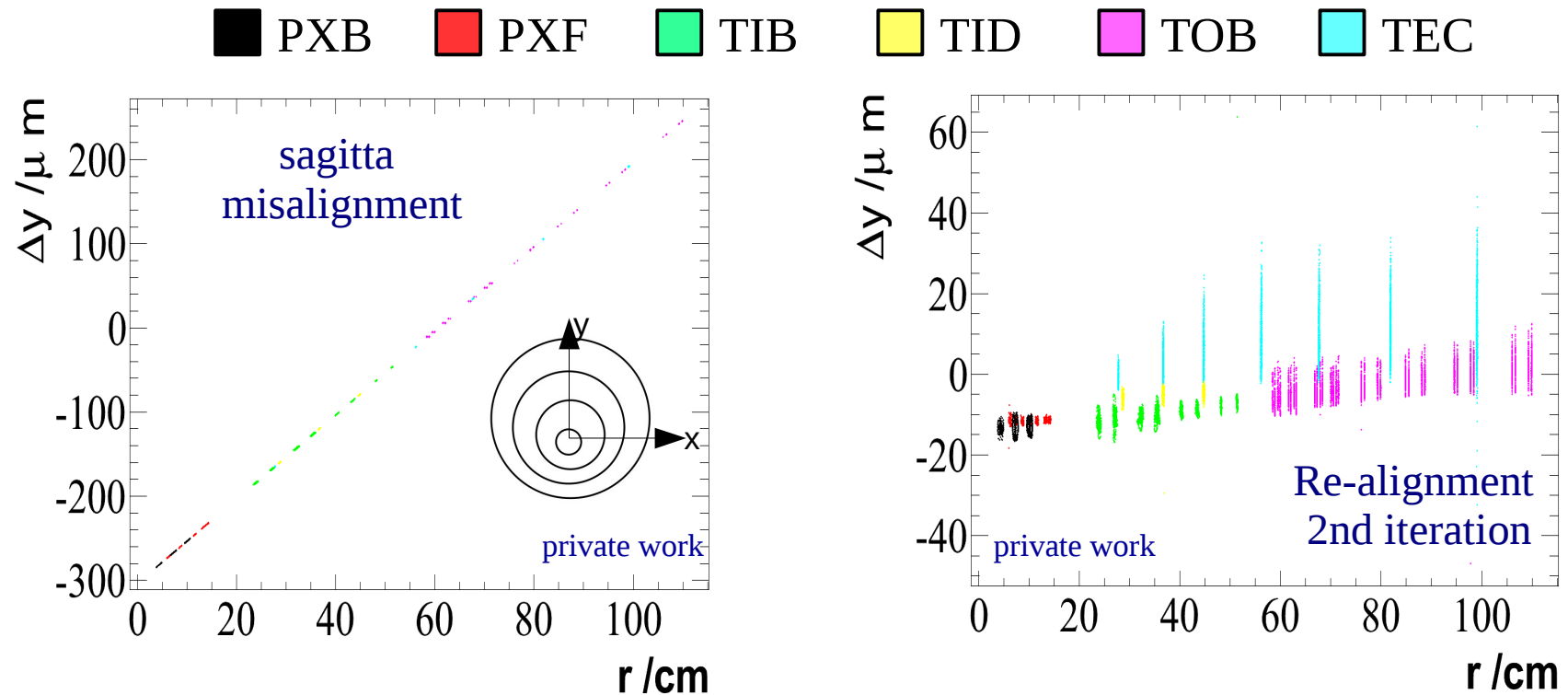


- Alignment procedure is sensitive to the twist weak mode with different data types
- Good recovery of twist mis-alignment with  $Z_{\mu\mu}$ , 0T cosmic data

# Misalignment recovery with sagitta mode

sagitta ( $\Delta y = c \cdot r$ )

*Use all data, include 0T cosmic &  $Z_{\mu\mu}$  !*

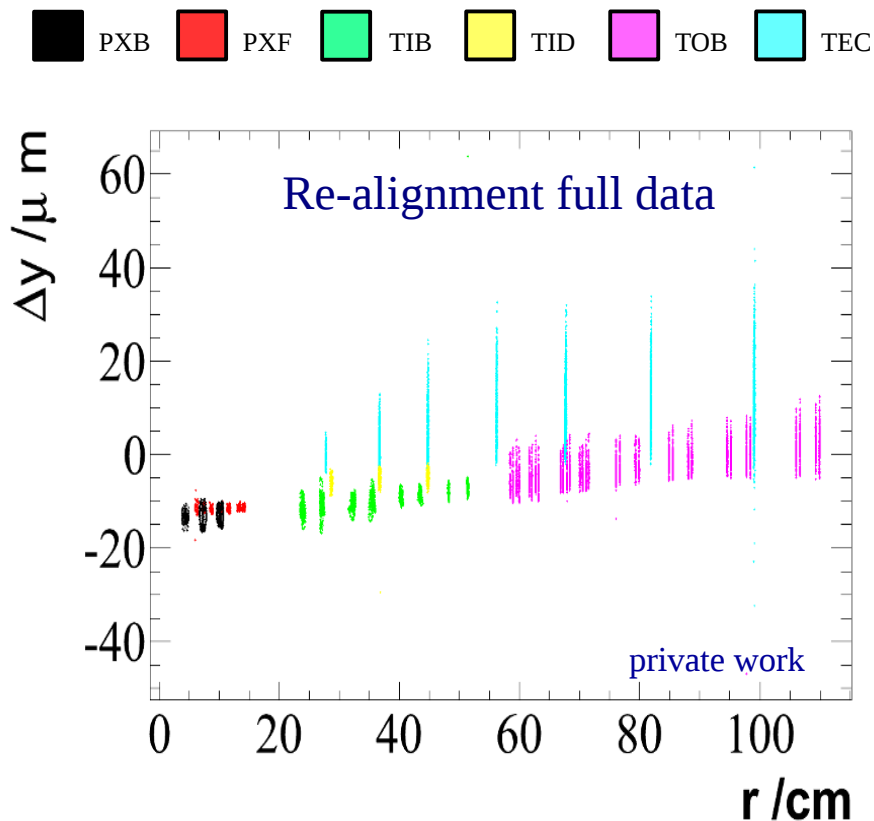


- In the case when mis-alignment is not linear – have to use few iterations
- Alignment procedure is sensitive to the sagitta weak mode
- Good recovery of sagitta mis-alignment

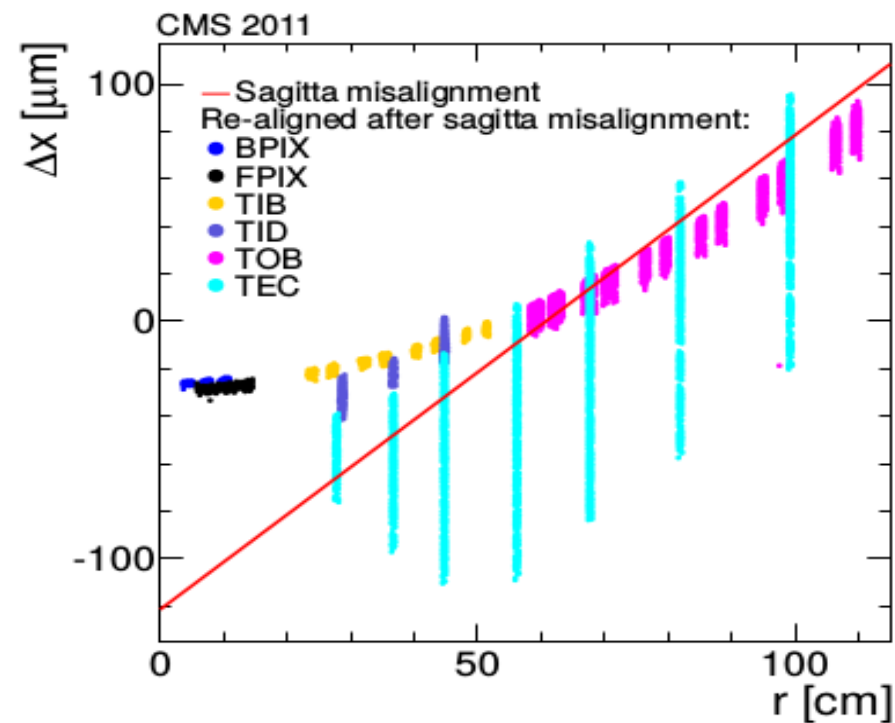
# Comparison with previous alignment studies

sagitta ( $\Delta y = c \cdot r$ )

Based on 2012 Alignment



Based on 2011 Alignment



Significant improvements with respect to the realignment of 2011 Alignment

## Summary

- Track-based alignment procedure of the CMS tracking detector has been explained
- Certain systematic distortions of the tracker geometry are very challenging for track-based alignment ("weak modes")
- Weak modes can be controlled by combining different data types in the alignment fit
  - ➔ good recovery has been demonstrated for the twist and sagitta weak-modes
  - ➔ further weak modes will be studied
- Significant improvements with respect to similar study with 2011 alignment

## Plans

- Study additional weak modes
- Apply legacy alignment methodology to Run II data

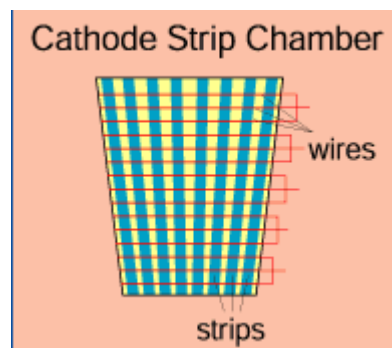
# Back Up

Data type N(events)

|                                   |            |
|-----------------------------------|------------|
| ➤ 0 T collisions :                | 320000     |
| ➤ 0 T cosmics :                   | 857970     |
| ➤ CRAFT cosmics:                  | 1 073 931  |
| ➤ interfill cosmics:              | 1 946 573  |
| ➤ interfill cosmics (peak mode) : | 1 770 243  |
| ➤ isolated $\mu$ :                | 14 788 959 |
| ➤ minimum bias:                   | 1 952 099  |
| ➤ $Z \rightarrow \mu\mu$ :        | 2 419 834  |

# Back Up

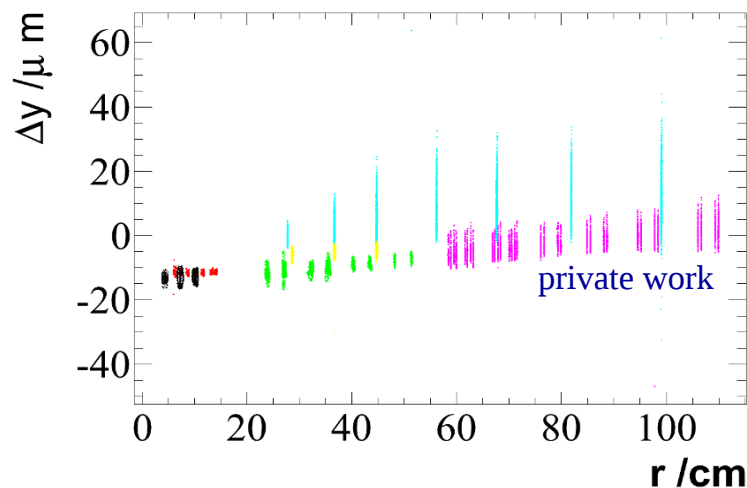
sagitta ( $\Delta y = c \cdot r$ )



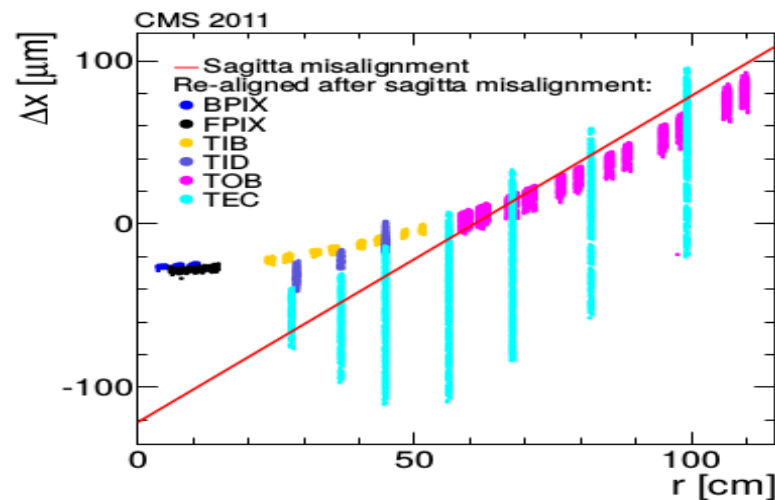
TEC consist of overlapping strip modules. Hence it has low sensitivity in Y directions.

Based on 2012 Alignment

PXB PXF TIB TID TOB TEC



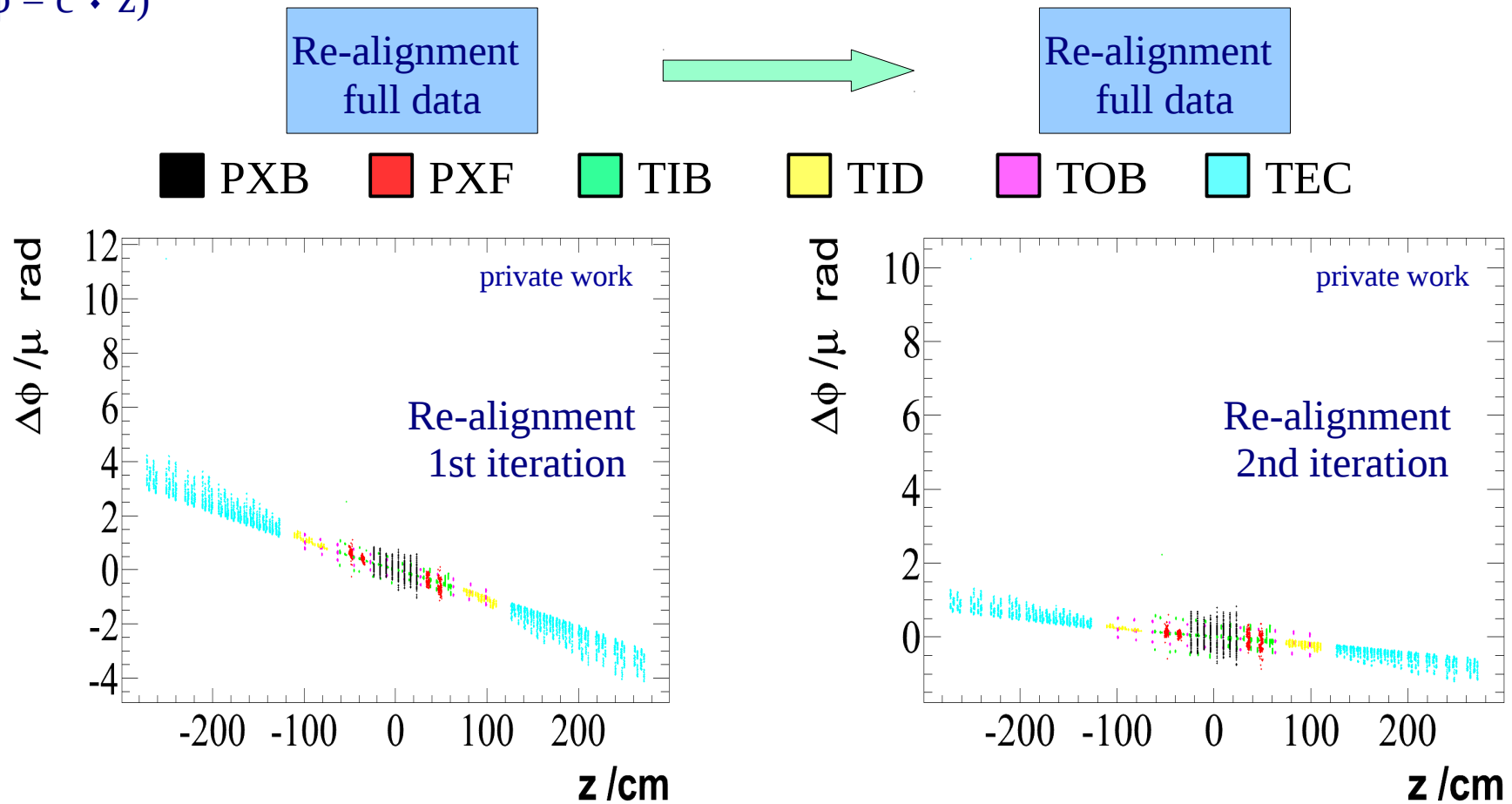
Based on 2011 Alignment



Significant improvements with respect to the realignment of 2011 Alignment

# Back Up. Misalignment recovery with twist mode

Twist ( $\Delta\phi = c \cdot z$ )



- Alignment procedure is sensitive to the twist weak mode with different data types
- Good recovery of twist mis-alignment with  $Z_{\mu\mu}$ , 0T cosmic data